

The Latest Status of the Submillimeter Array (SMA)

Naomi Hirano (ASIAA, Taiwan)

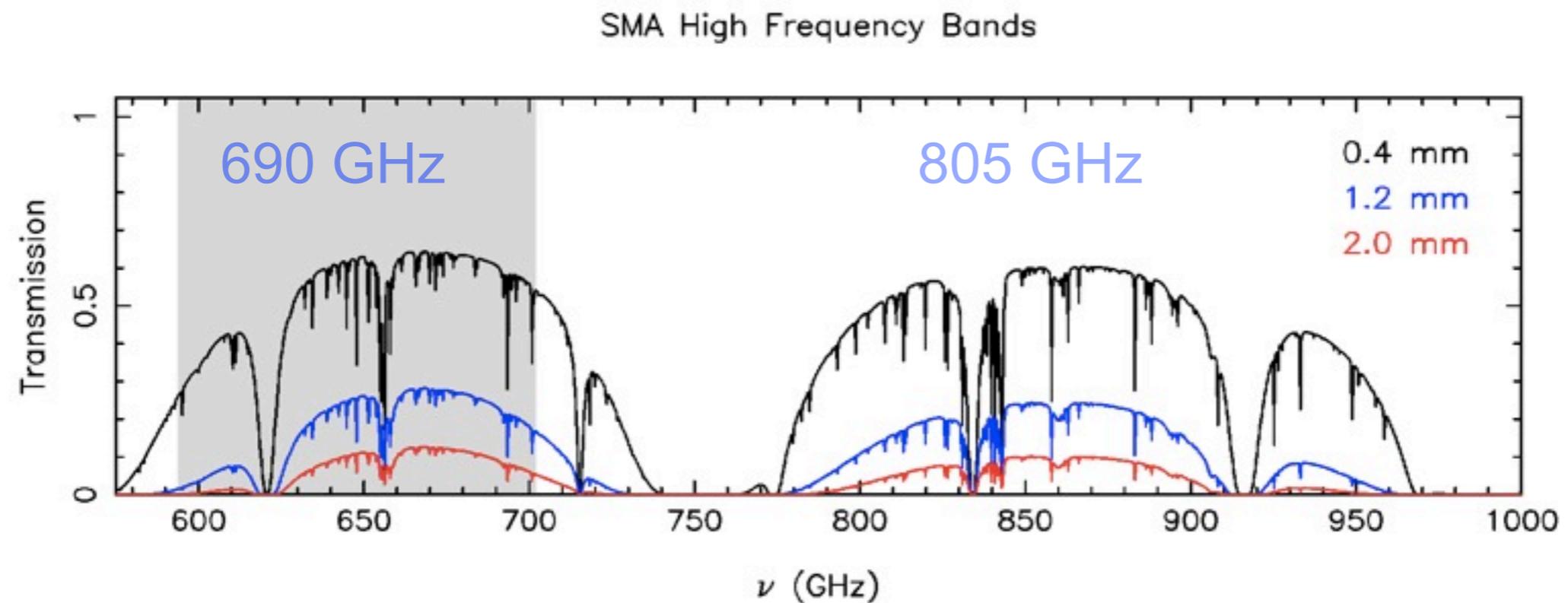
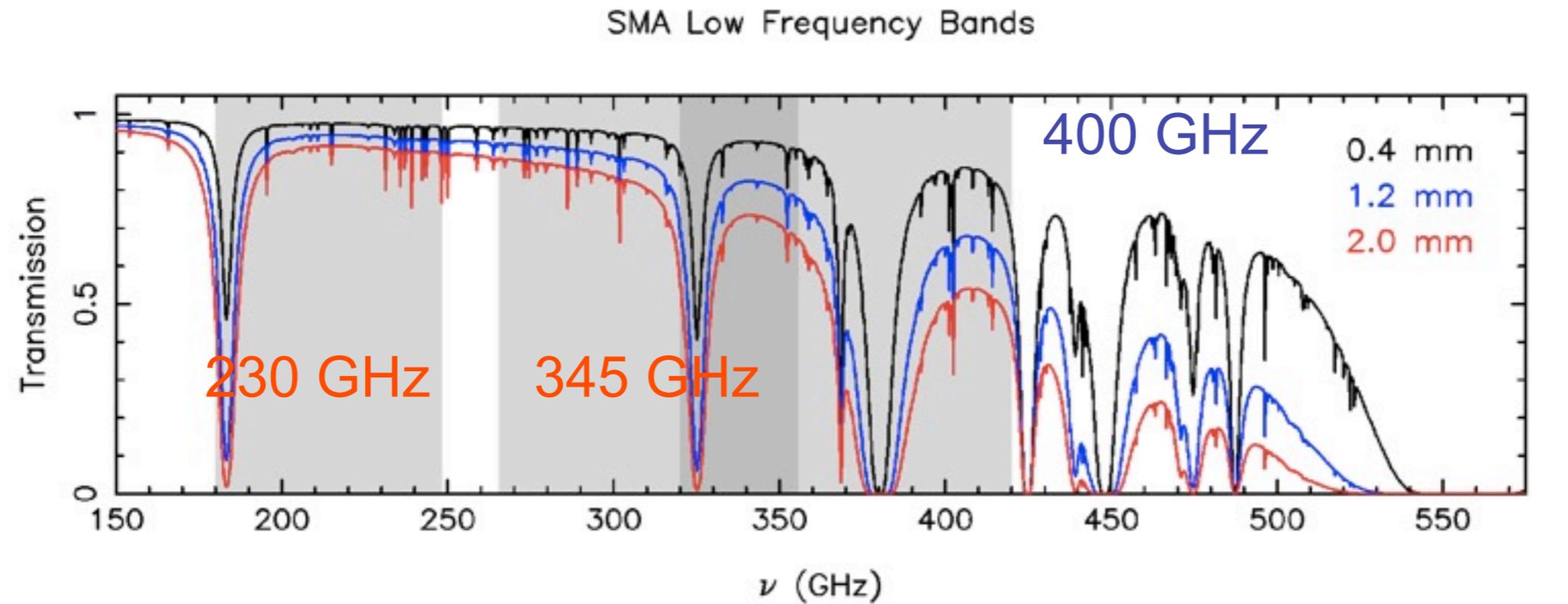




Since 2003



Receiver Bands/Atmosphere



Array configurations

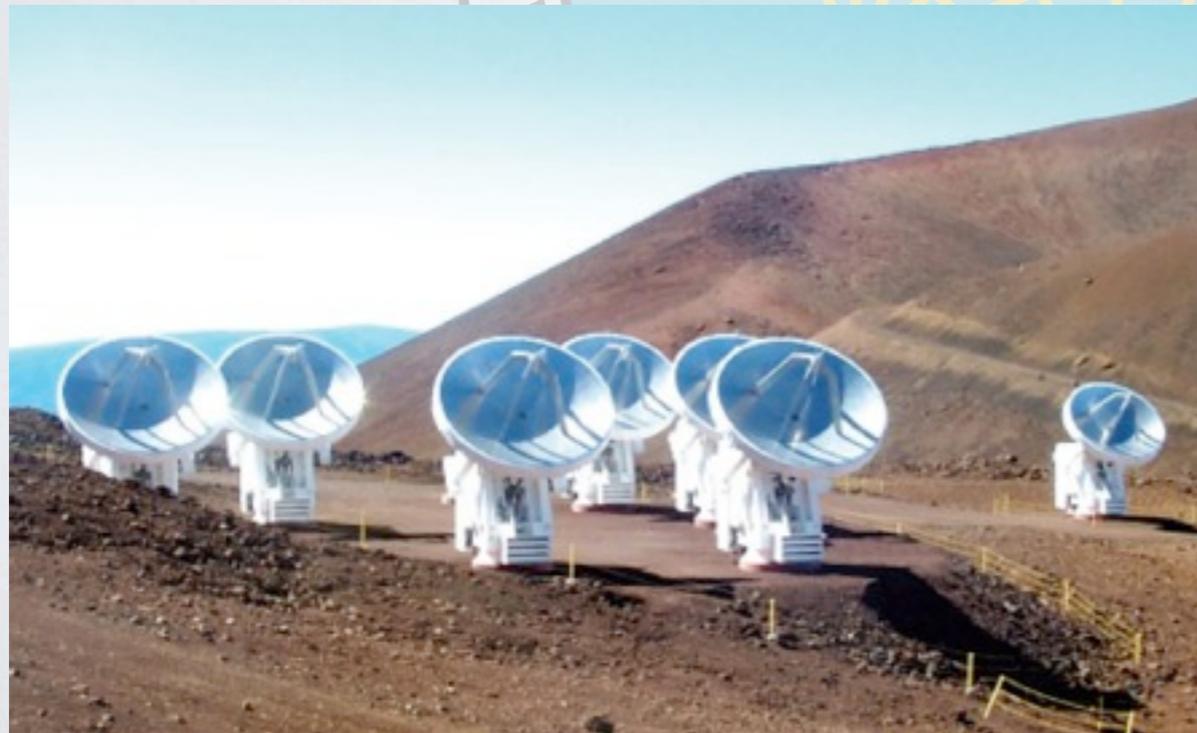


$$\theta \text{ (radian)} \sim \lambda \text{ (wavelength)} / D \text{ (diameter)}$$

Longest baseline : 509 m



× angular resolution: $\sim 0.3''$ @ 345 GHz



**Shortest baseline :
9.5 m**

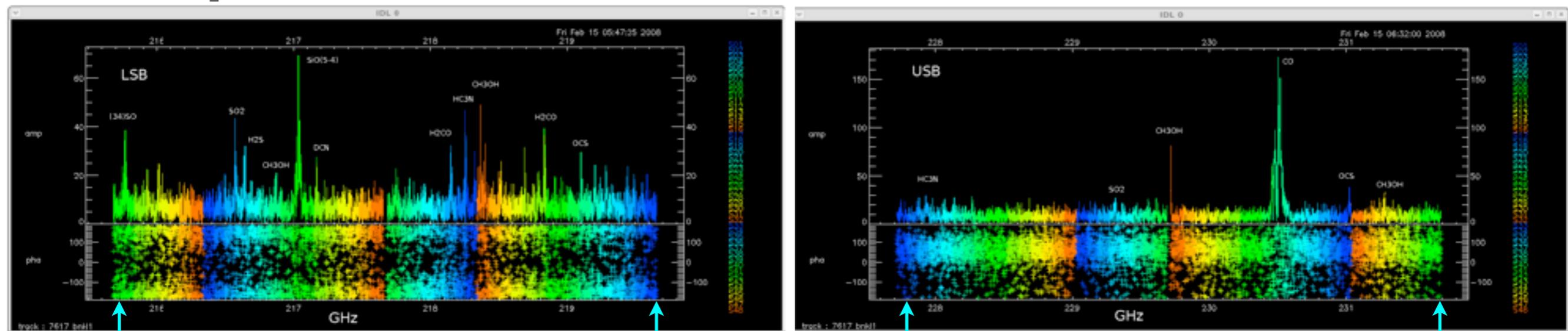
The Capability of wide bandwidth

LSB: 4 GHz = 48 x 82 MHz

USB: 4 GHz = 48 x 82 MHz



The spectra from Orion KL



215.7 GHz

219.7 GHz

227.7 GHz

231.7 GHz

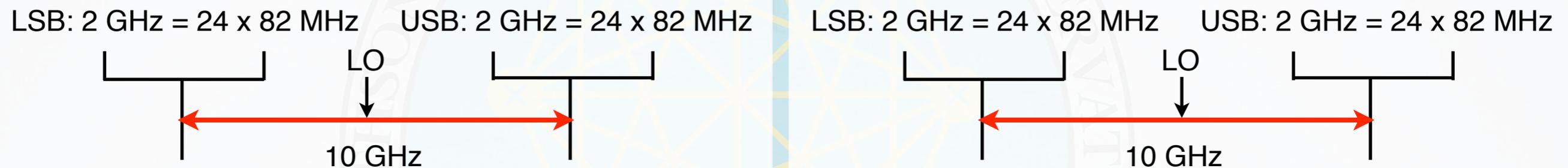
4 GHz + 4 GHz can be observed at the same time.

Dual band mode

2 receivers, 2 GHz (each) mode

low freq. receiver

high freq. receiver



Rx230 or Rx345 & Rx400 (Rx690)

ex.

CO 3-2 & SiO 8-7 in Rx 345 + CO 3-2 & SiO 8-7 in Rx400

CO 2-1 in Rx230 + CO 3-2 in Rx400

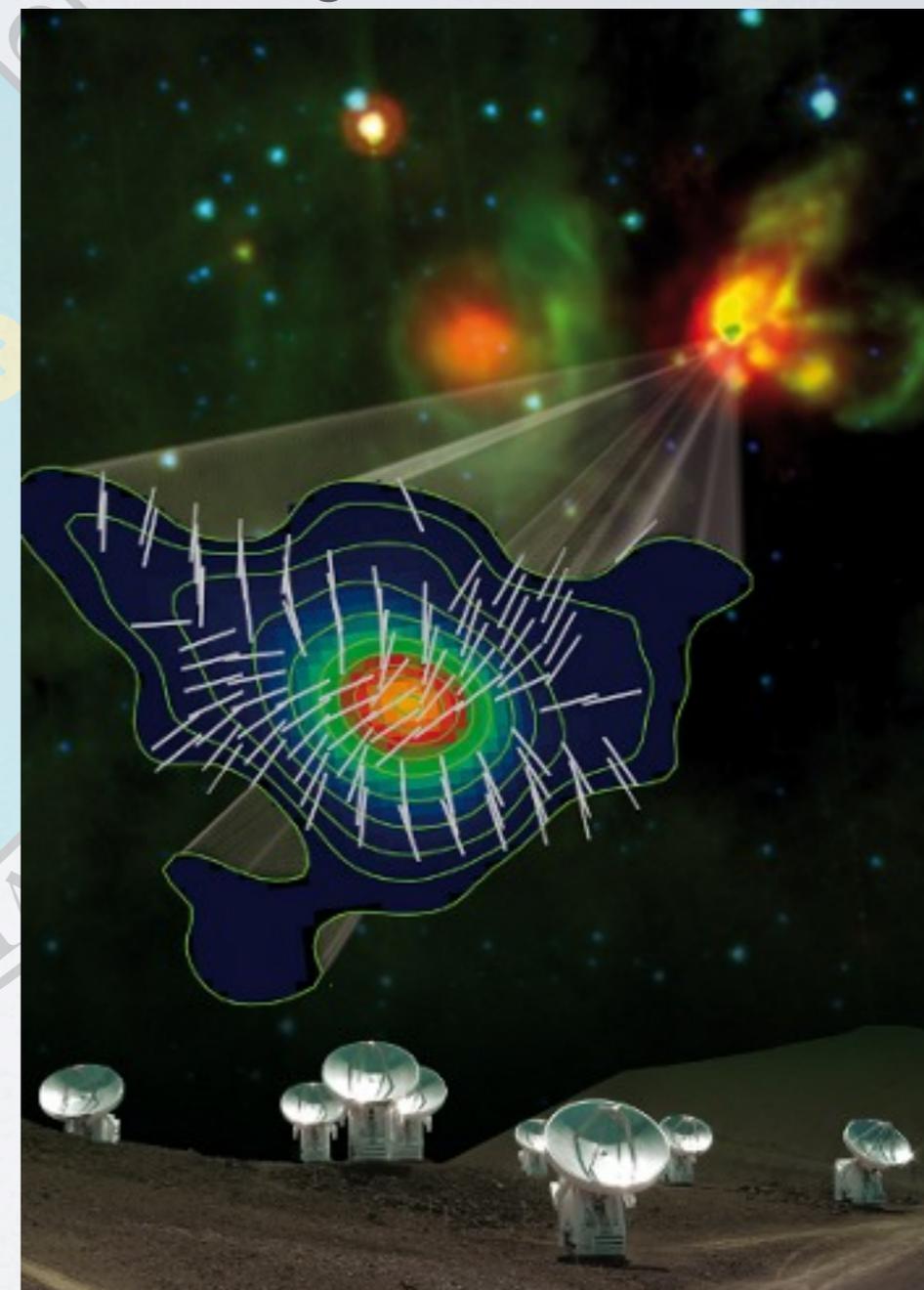
Polarization mode

Continuum & lines

Dual receiver (Rx 345 + Rx 400) mode @ 330–355 GHz is available!!

**using quarter-wave plates
230/(690), 342, 240/400 GHz**

870 μm dust continuum emission
from the massive star forming
region G31.41



Josep Miquel Girart (CSIC-IEEC), Nimesh Patel (Harvard-Smithsonian Center for Astrophysics) and Manel Carrillo

The New Capability for Rx 230 & Rx 345

LSB: 4 GHz = 48 x 82 MHz

USB: 4 GHz = 48 x 82 MHz

LSB: 8 GHz

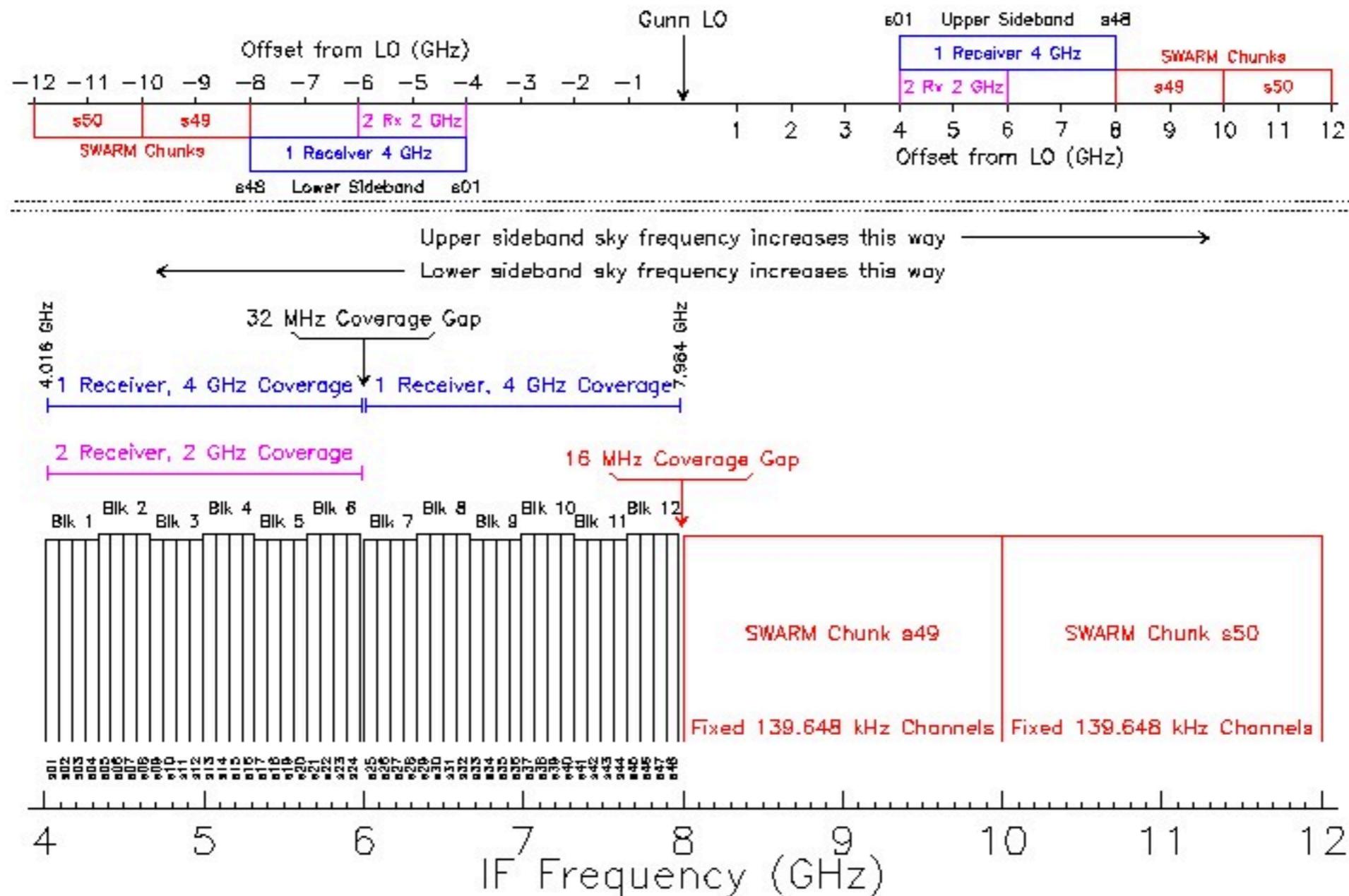
10 GHz

USB: 8 GHz

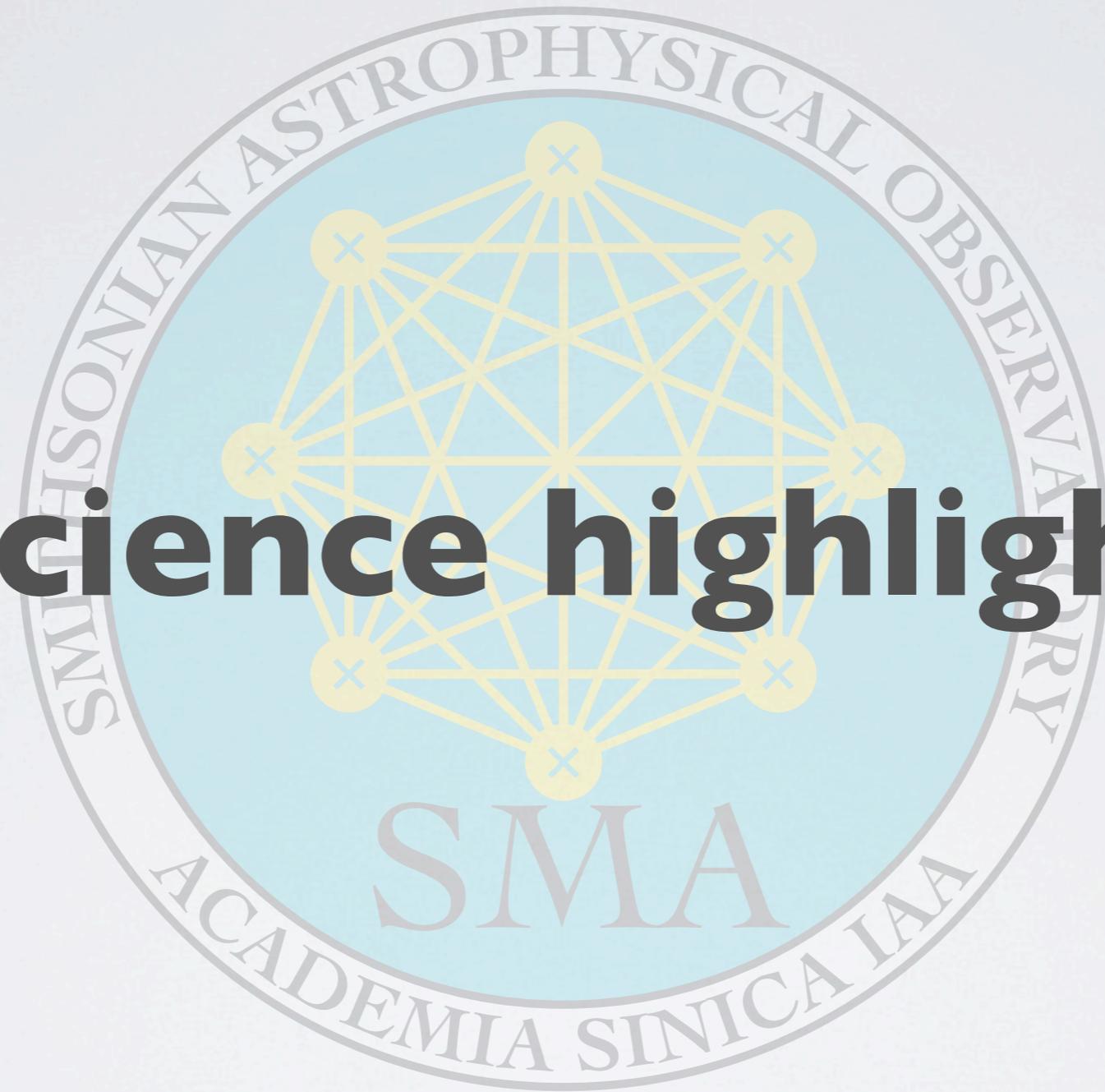
8 GHz + 8 GHz can be observed at the same time.

Coming soon!

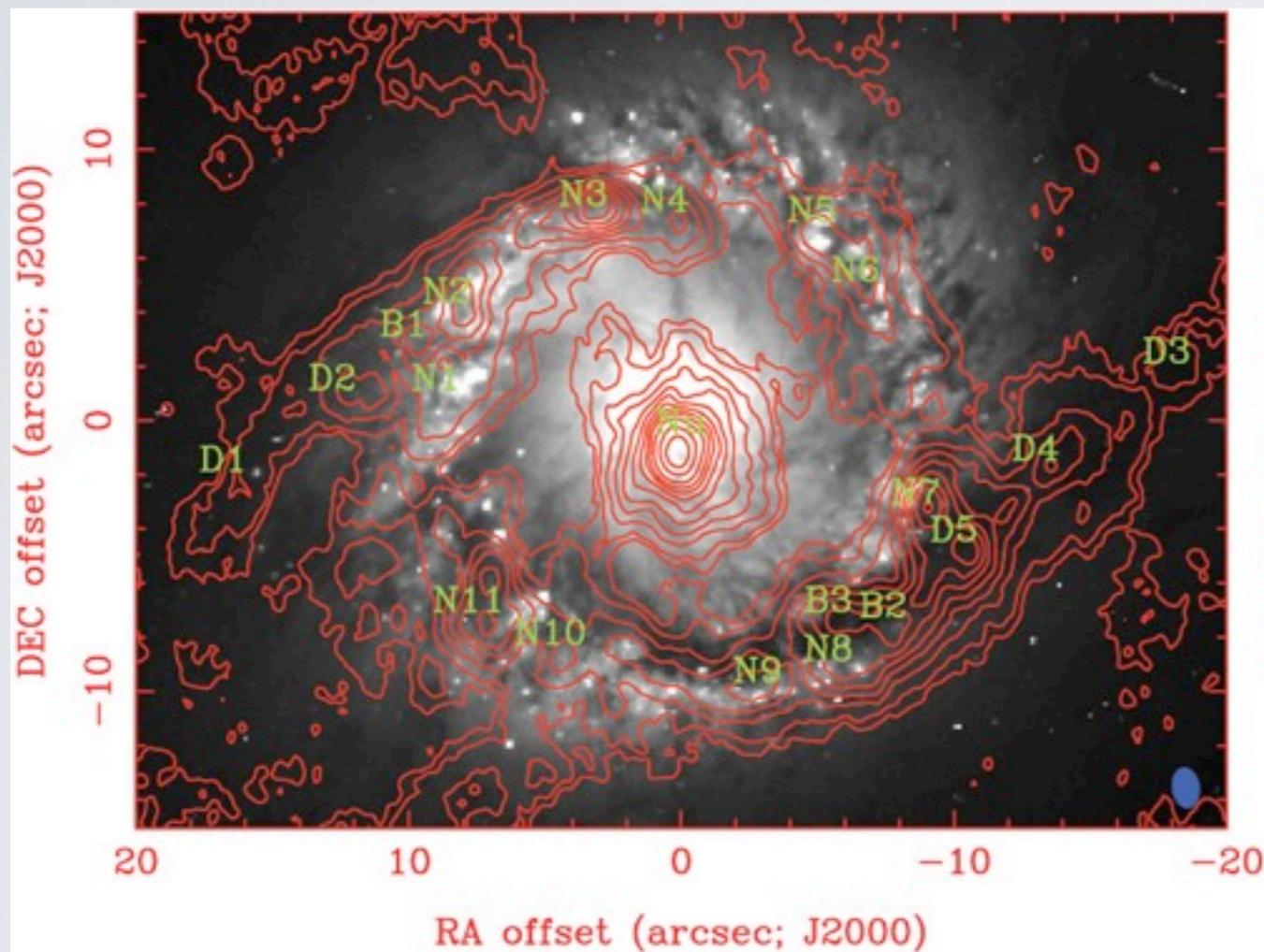
The “Legacy” and the new “SWARM” correlators



Science highlights

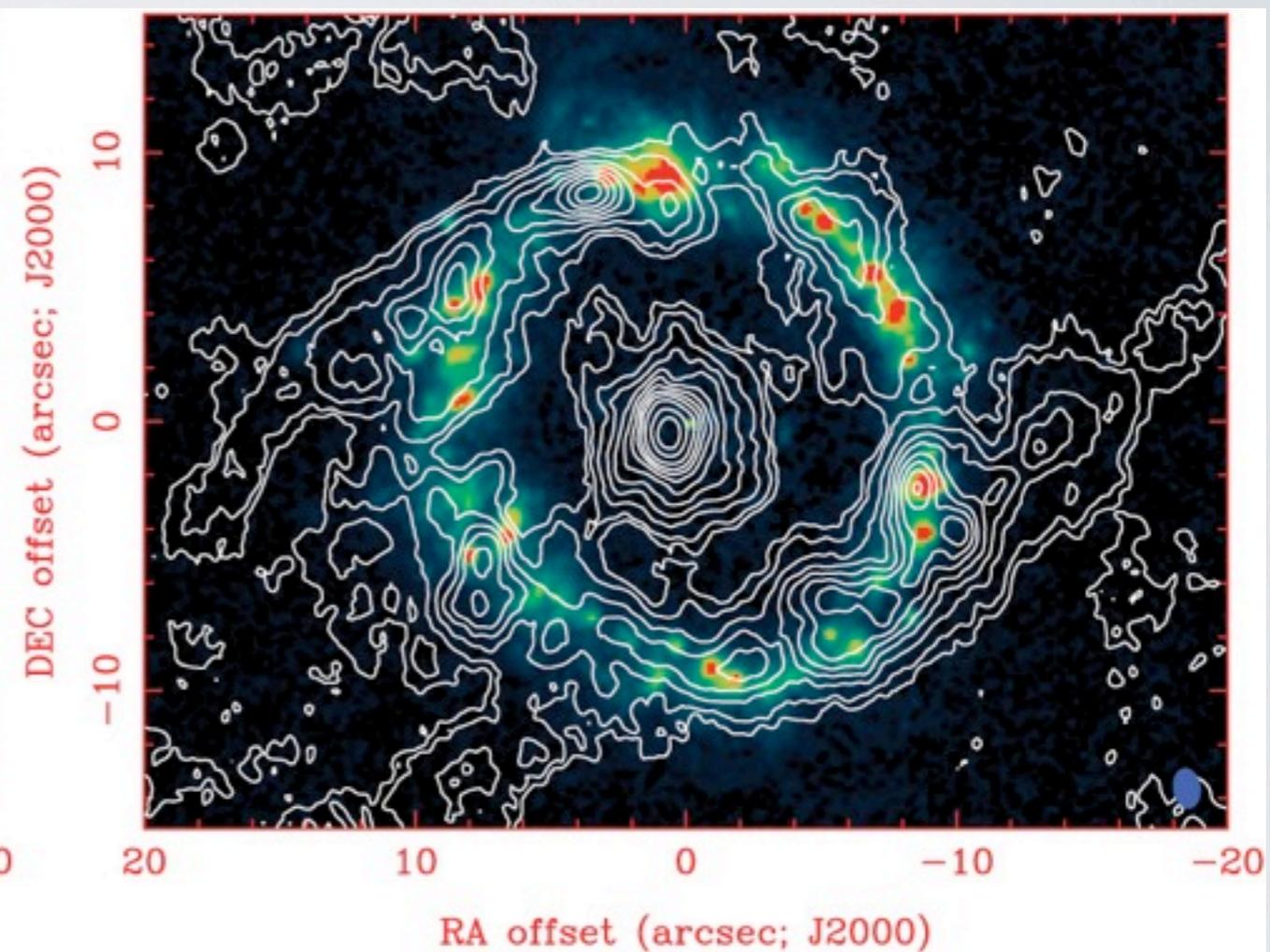


SEYFERT I / STARBURST RING GALAXY NGC 1097



CO 2-1 + HST I-band image

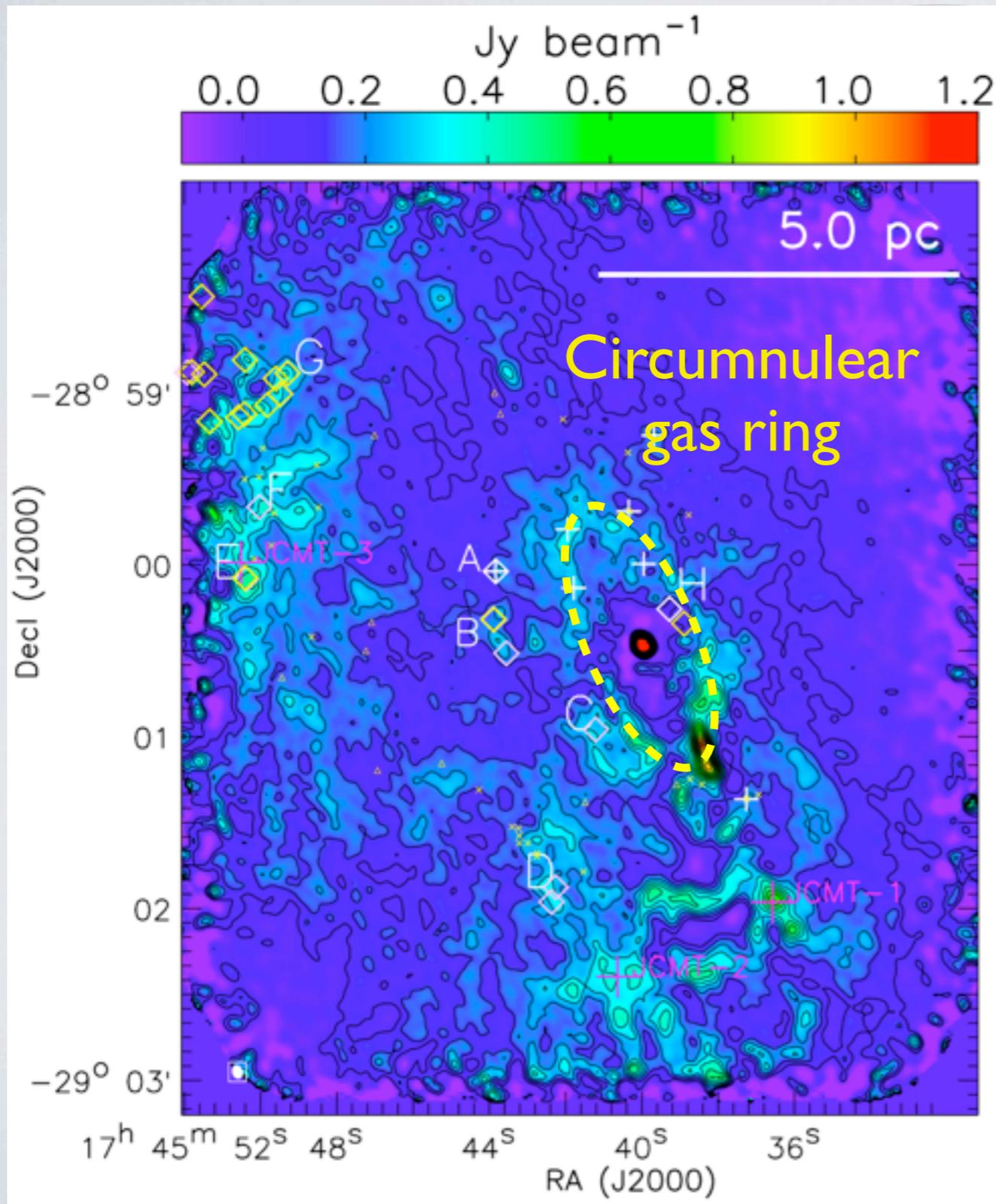
1.5" x 1.0" (105 x 70 pc) resolution



CO 2-1 + HST NICMOS Pa Alpha

Hsieh P.-Y. et al. 2011, ApJ, 736, 129

The First Interferometric 0.86 mm Dust Continuum Image in the Milky Way Center

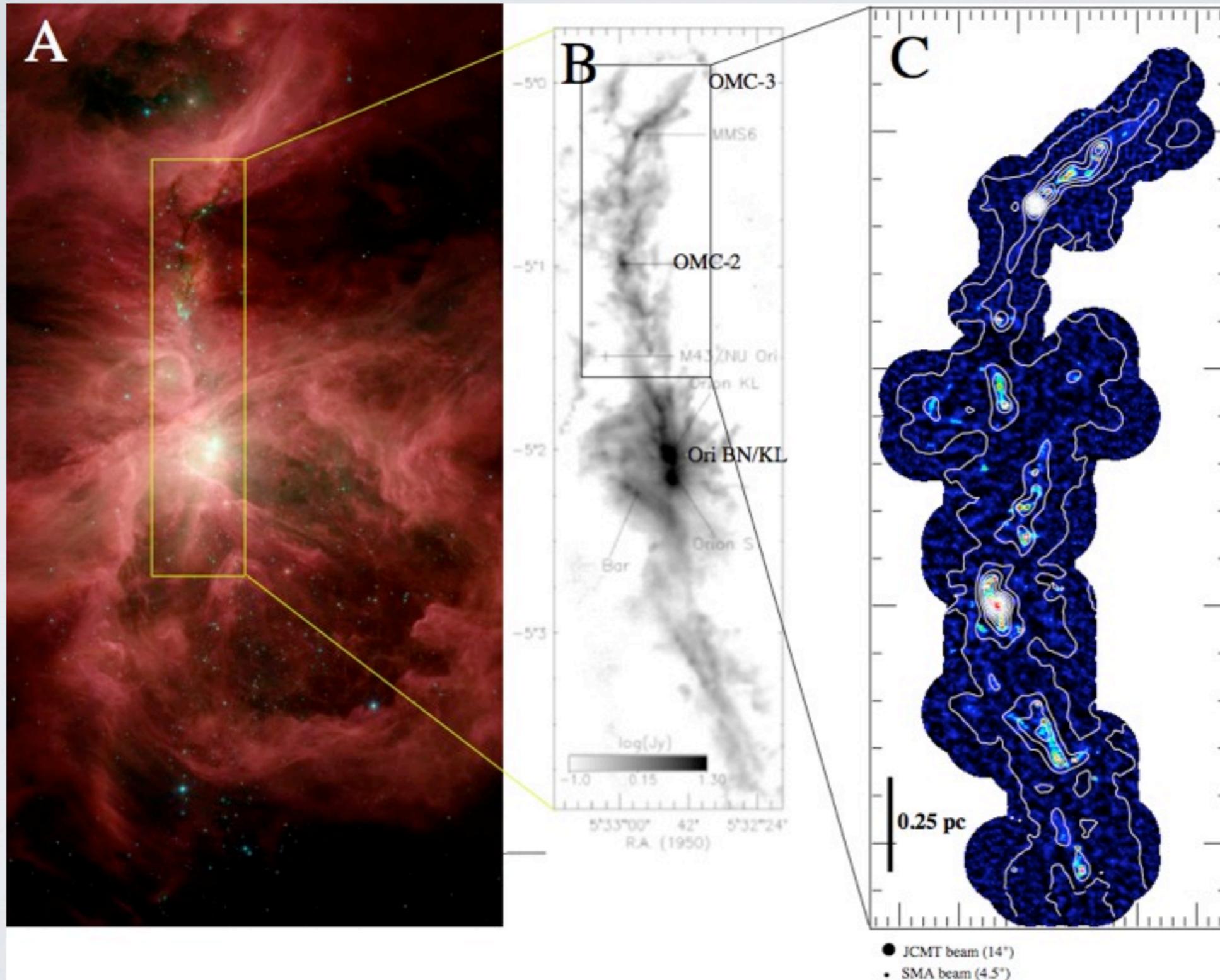


5' x 5' area
157 SMA pointings
+ JCMT
@~5" resolution

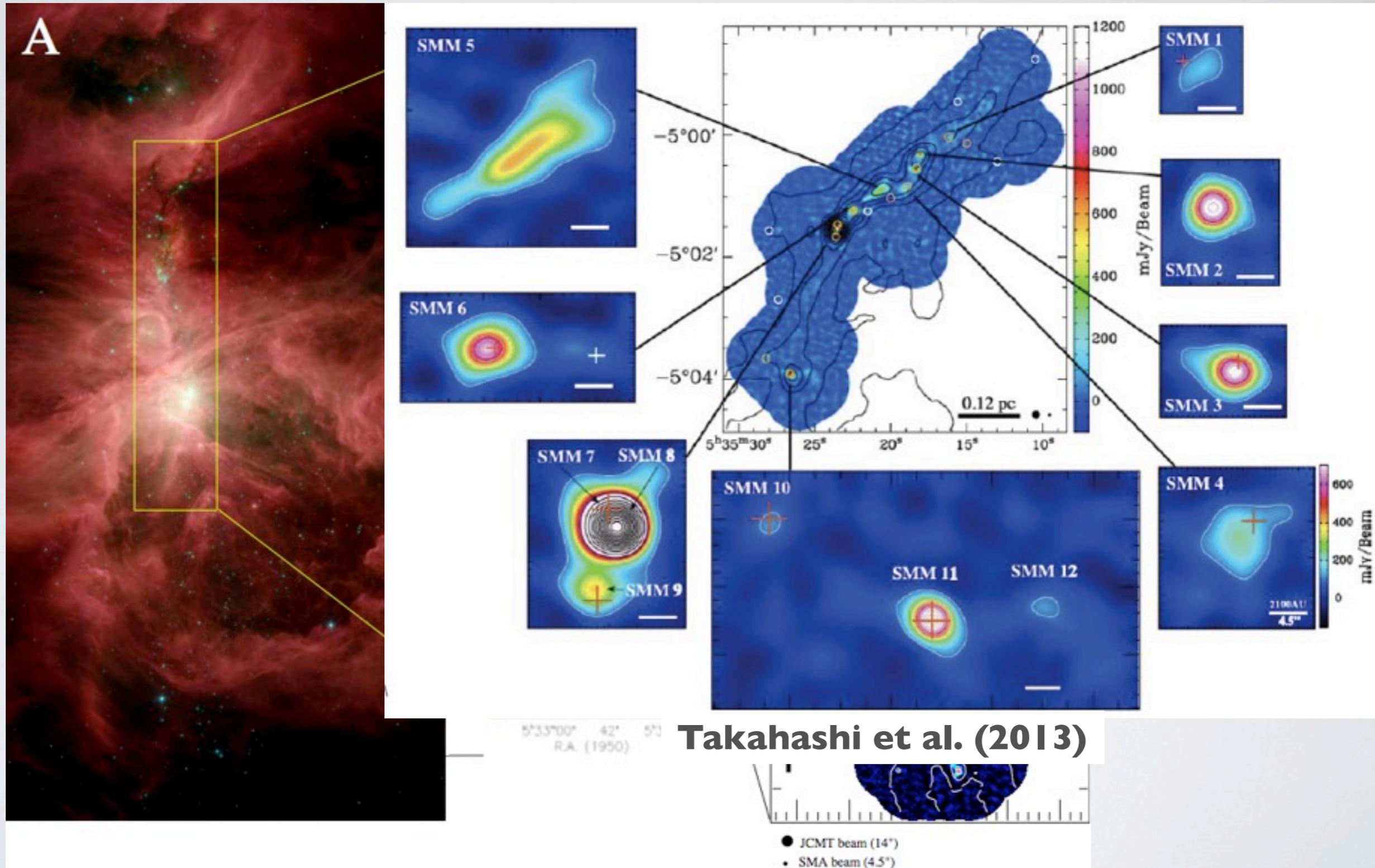
Circumnuclear gas ring
+
arms/streamers

Liu HB et al. 2013, ApJ, 770, 44

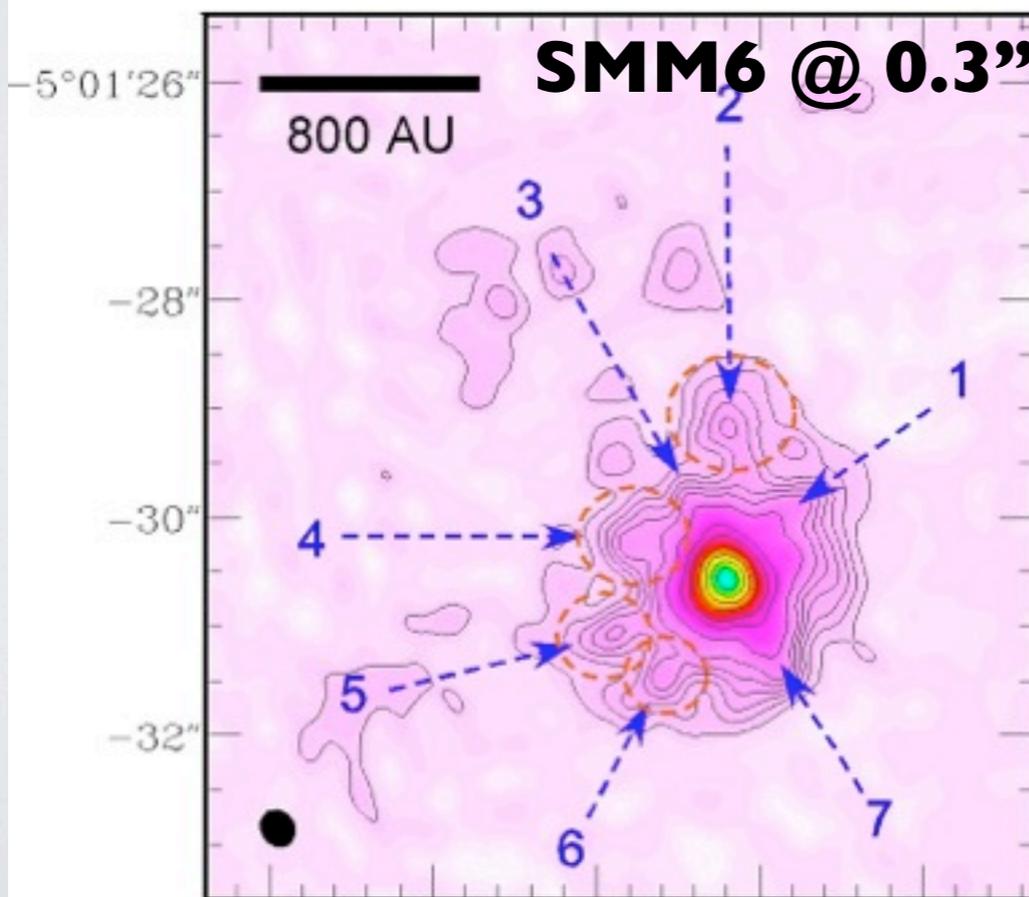
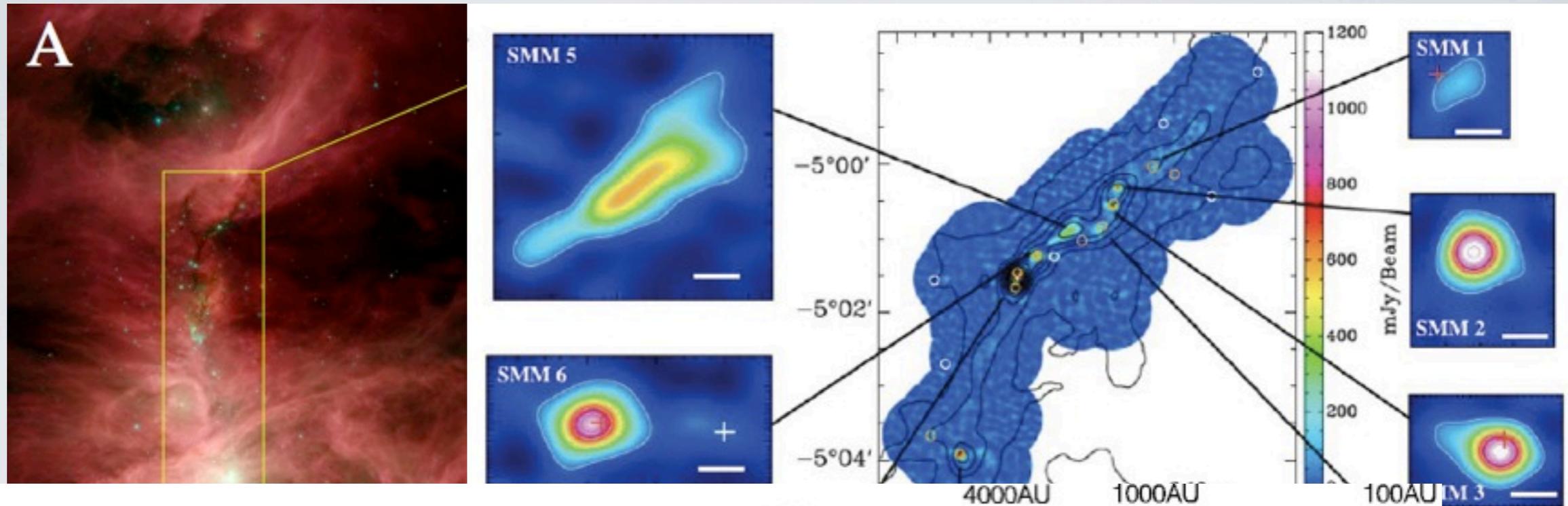
Cores in Orion Molecular cloud



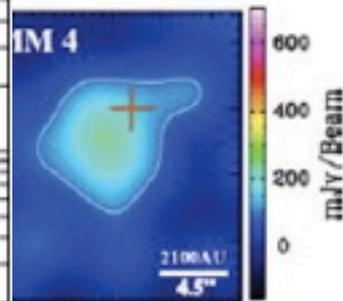
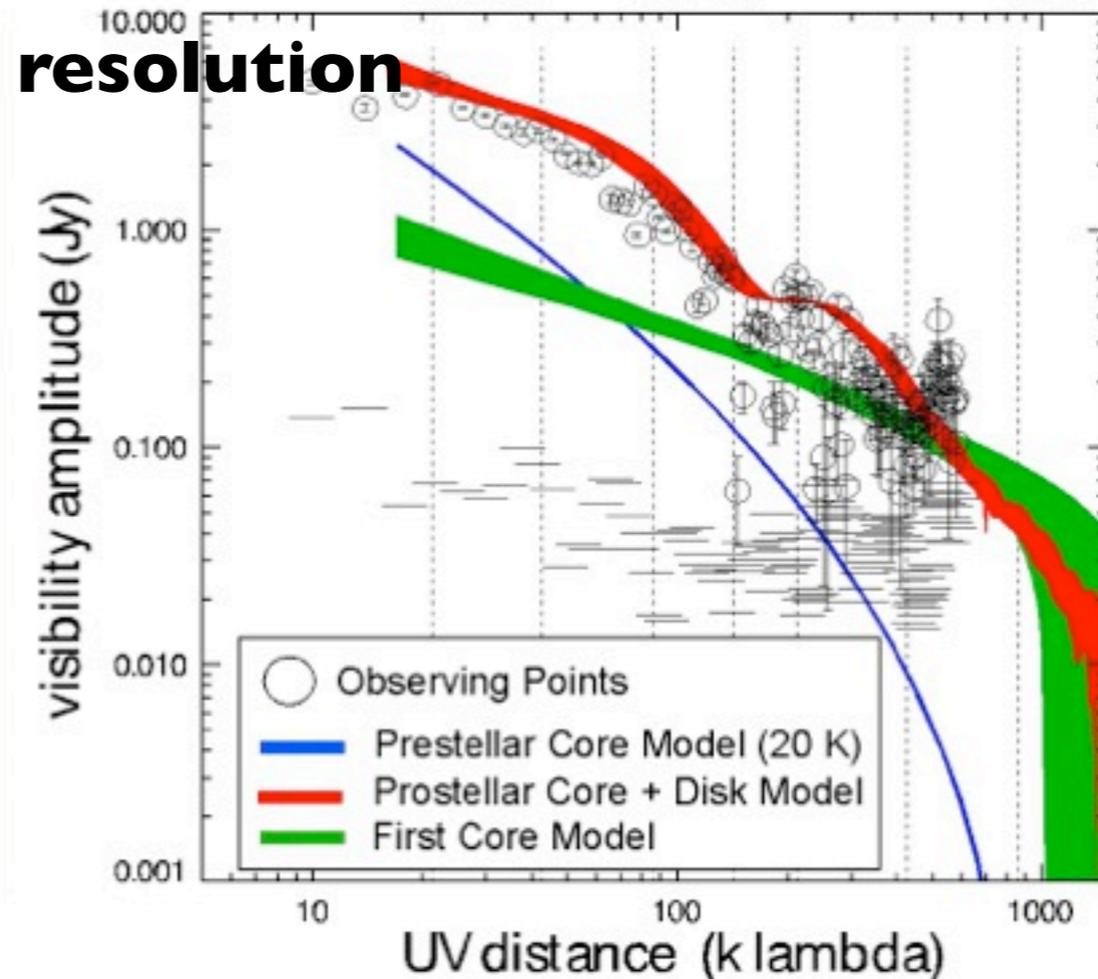
Cores in Orion Molecular cloud



Cores in Orion Molecular cloud

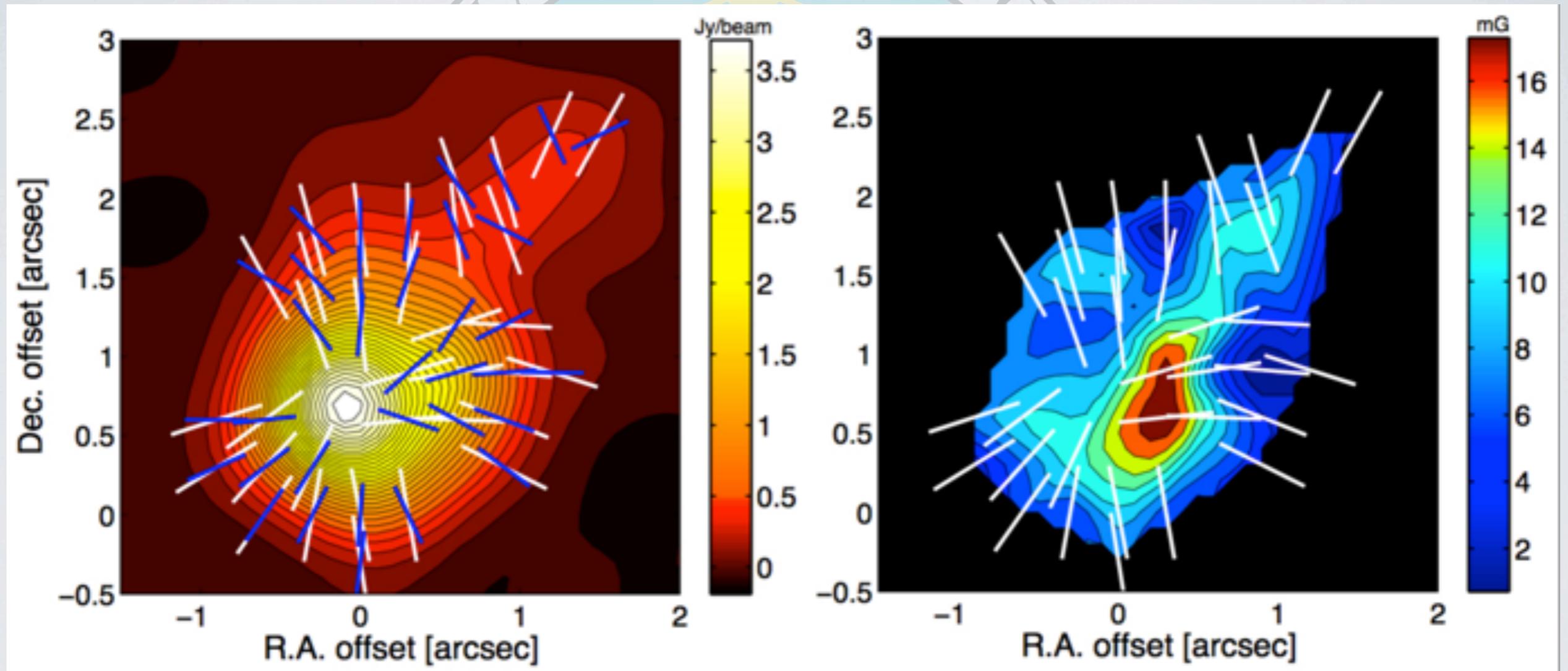


Takahashi et al. (2012)



Magnetic Fields Measurements @ W51 e2

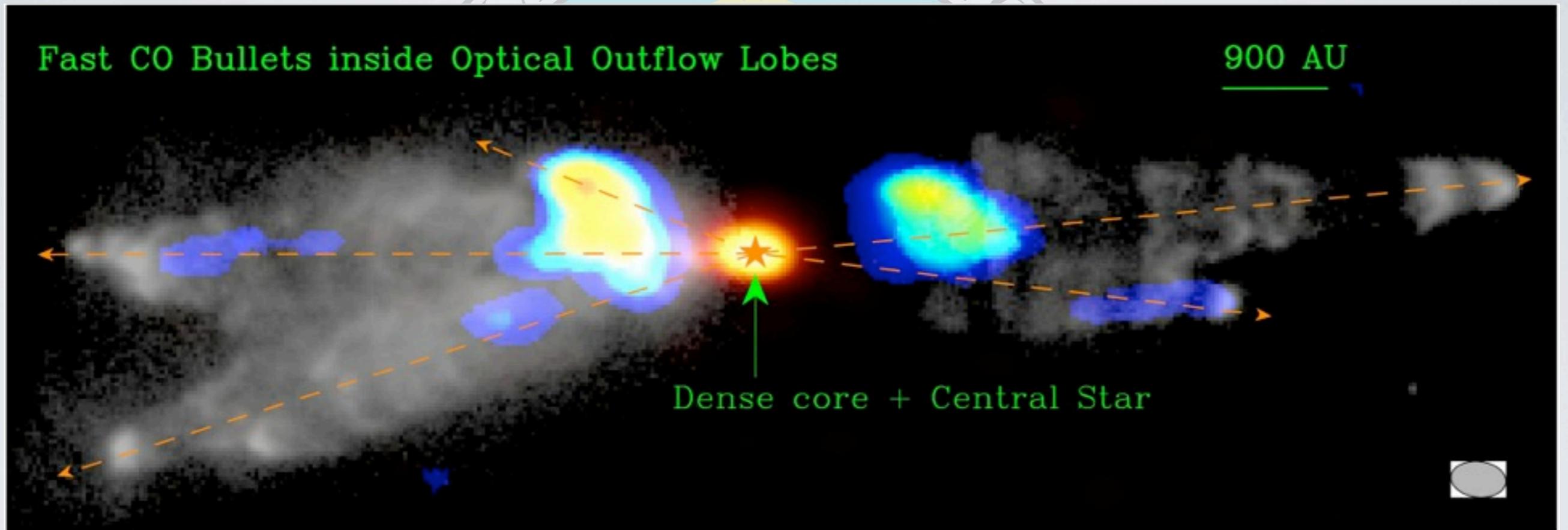
dust continuum @ 345 GHz B-field strength map



white: B-field direction
blue: intensity gradient

Koch et al. 2012, ApJ, 747, 79

Multiple Fast Molecular Outflows in the PPN CRL 618



Multiple fast molecular outflows with two different dynamical ages oriented along the different optical lobes

Lee C.-F. et al. 2013, ApJ, in press

The background features a large, semi-transparent watermark of the SMA logo. The logo is circular with a light blue center containing a yellow network diagram of nodes and connections. The outer ring of the logo contains the text "SMITHSONIAN ASTROPHYSICAL OBSERVATORY" at the top and "ACADEMIA SINICA TAIPEI" at the bottom. The acronym "SMA" is prominently displayed in the center of the logo.

The SMA is welcome
to receive your
observing proposals.

<http://sma1.sma.hawaii.edu>