# Black Holes，Big and Small Impact on Galaxy Formation 



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Miyoshi et al. (1995)
Herrnstein et al. (2005)


Barth, Ho et al. (2001)

The "Nuker" Team


## Correlation Between Black Hole Mass and Bulge Mass



Gebhardt et al. (2000); Ferrarese \& Merritt (2000); Gültekin et al. (2009)

bulge velocity dispersion

## $\mathrm{T}=0 \mathrm{Myr}$




Gaiblier et al. (2012)


Crotton et al. (2006)


Schawinski et al. (2010)


## Standard "Paradigm"

OAll bulges contain BHs
$\bigcirc M_{\bullet} \sim M_{\text {bulge }}^{1.0} \quad\left\langle M_{\bullet} / M_{\text {bulge }}\right\rangle \sim 0.1 \%-0.2 \%$
$\bigcirc M_{\bullet} \propto \sigma^{4}$
$\bigcirc M_{\bullet}-\sigma$ relation tighter than $M_{\bullet}-M_{\text {bulge }}$ relation
No strong dependence on galaxy mass or type
$\bigcirc$ Mild to strong evolution with redshift
$\bigcirc$ AGN feedback engineers BH-host correlations

## Recent Developments

Kormendy \& Ho (2013, ARA\&A):
Coevolution (or Not) of Supermassive Black Holes and Galaxies











Table 2 Supermassive black holes detected dynamically in 45 elliptical galaxies (December 2012)


Table 3 Supermassive black holes detected dynamically in spiral and S0 galaxies galaxies (21 with classical bulges; 22 with pseudobulges; December 2012)

| Object (1) | Type (2) | Distance (Mpc) <br> (3) | $K_{s}$ (4) | $M_{K s T}$ ${ }_{(5)}$ | $M_{K s, \text { bulge }}$ (6) | $M_{K s, \text { disk }}$ (7) | $B / T$ $(8)$ | $P B / T$ ${ }^{(9)}$ | $\begin{gathered} M_{V T} \\ (10) \end{gathered}$ | $M_{V, \text { bulge }}$ <br> (11) | $M_{V, \text { disk }}$ <br> (12) | $\left(V-K_{s}\right)_{0}$ <br> (13) | $(B-V)_{0}$ <br> (14) | $\begin{gathered} \log M_{\text {bulge }} \\ \left(M_{\odot}\right) \\ (15) \end{gathered}$ | $\begin{gathered} M_{\bullet}\left(\text { low } M_{\bullet}-\text { high } M_{\bullet}\right) \\ \left(M_{\odot}\right) \\ (16) \end{gathered}$ | $\begin{gathered} \sigma_{e} \\ \left(\mathrm{~km} \mathrm{~s}^{-1}\right) \\ (17) \end{gathered}$ | $\begin{gathered} V_{\text {circ }} \\ \left(\mathrm{km} \mathrm{~s}^{-1}\right) \\ (18) \end{gathered}$ | Flags <br> M $M_{\bullet}$ <br> (19) | Source $(20)$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| M 31 | Sb | 0.7748 | 0.573 | -23.89 | -22.62 | -23.85 | 0.31 | 0.0 | -21.14 | -19.64 | -20.83 | 2.746 | 0.865 | $10.35 \pm 0.09$ | $1.43(1.12-2.34) \times 10^{8}$ | $169 \pm 8$ | $250 \pm 20$ | 10 | Bender + 2005 |
| M 81 | Sb | 3.6048 | 3.831 | -23.98 | -22.81 | -23.53 | 0.34 | 0.0 | -21.06 | -19.89 | -20.61 | 2.913 | 0.879 | $10.42 \pm 0.09$ | $6.5(5 .-9.) \times 10^{7}$ | $143 \pm 7$ | $240 \pm 10$ | 1,20 | see notes |
| NGC 52 | S0 | 24.222 | 7.163 | -24.78 | -24.69 | -22.04 | 0.92 | 0.0 | -21.86 | -21.77 | -19.12 | 2.923 | 0.977 | $11.26 \pm 0.09$ | $8.67(8.21-9.61) \times 10^{8}$ | $247 \pm 12$ |  | 10 | Krajnović + 2009 |
| NGC 821 | S0 | 23.442 | 7.715 | -24.17 | -24.11 | -20.92 | 0.95 | 0.0 | -21.19 | -21.13 | -17.94 | 2.980 | 0.893 | $10.98 \pm 0.09$ | $1.65(0.92-2.39) \times 10^{8}$ | $209 \pm 10$ |  | 11 | Schulze + 2011 |
| NGC 1023 | SB0 | 10.812 | 6.238 | -23.95 | -22.93 | -23.41 | 0.39 | 0.0 | -20.99 | -19.96 | -20.45 | 2.964 | 0.946 | $10.53 \pm 0.09$ | $4.13(3.71-4.56) \times 10^{7}$ | $205 \pm 10$ | $251 \pm 15$ | 10 | Bower + 2001 |
| NGC 1194 | S0/ | 57.989 | 9.758 | -24.08 | -23.33 | -23.33 | 0.5 | 0.0 | -21.16 | -20.40 | -20.40 | 2.925 | 0.893 | $10.64 \pm 0.09$ | $7.08(6.76-7.41) \times 10^{7}$ | $148 \pm 24$ | $203 \pm 16$ | 30 | Kuo + 2011 |
| NGC 1277 | S0/ | 73. 10 | 9.813 | $-24.63$ | -23.98 | -23.76 | 0.55 | 0.0 | -21.31 | -20.67 | -20.45 | 3.313 | 0.985 | $11.00 \pm 0.09$ | $1.7(1.4-2.0) \times 10^{10}$ | $333 \pm 17$ |  | 11 | van den Bosch +2012 |
| NGC 2549 | S0/ | 12.702 | 8.046 | -22.49 | -20.99 | -22.18 | 0.25 | 0.0 | -19.55 | -18.05 | -19.24 | 2.942 | 0.912 | $9.71 \pm 0.09$ | $1.45(0.31-1.65) \times 10^{7}$ | $145 \pm 7$ |  | 10 | Krajnović + 2009 |
| NGC 3115 | S0/ | 9.542 | 5.883 | -24.03 | -23.91 | -21.53 | 0.90 | 0.0 | -21.13 | -21.01 | -18.63 | 2.901 | 0.928 | $10.92 \pm 0.09$ | $8.97(6.20-9.54) \times 10^{8}$ | $230 \pm 11$ | $315 \pm 20$ | 11 | Emsellem +1999 |
| NGC 3245 | S0 | 21.382 | 7.862 | -23.80 | -23.41 | -22.49 | 0.70 | 0.0 | -20.88 | -20.50 | -19.58 | 2.914 | 0.888 | $10.69 \pm 0.09$ | $2.39(1.63-2.66) \times 10^{8}$ | $205 \pm 10$ |  | 20 | Barth + 2001 |
| NGC 3585 | S0 | 20.512 | 6.703 | -24.88 | -24.80 | -21.99 | 0.93 | 0.0 | -21.93 | -21.85 | -19.04 | 2.951 | 0.913 | $11.26 \pm 0.09$ | $3.29(2.71-4.74) \times 10^{8}$ | $213 \pm 11$ | $280 \pm 20$ | 10 | Gultekin +200 |
| NGC 3998 | S0 | 14.302 | 7.365 | -23.42 | -23.24 | -21.36 | 0.85 | 0.0 | -20.63 | -20.46 | -18.57 | 2.786 | 0.936 | $10.67 \pm 0.09$ | $8.45(7.79-9.15) \times 10^{8}$ | $275 \pm 7$ |  | 11 | Walsh +2012 |
| NGC 3998 | S0 | 14.302 | 7.365 | -23.42 | -23.24 | -21.36 | 0.85 | 0.0 | -20.63 | -20.46 | -18.57 | 2.786 | 0.936 | $10.67 \pm 0.09$ | $2.27(1.43-3.28) \times 10^{8}$ | $275 \pm 7$ |  | 20 | DeFrancesco +2006 |
| NGC 4026 | S0 | 13.352 | 7.584 | -23.05 | -22.51 | -22.03 | 0.61 | 0.0 | -20.01 | -19.47 | -18.99 | 3.043 | 0.900 | $10.33 \pm 0.09$ | $1.80(1.45-2.40) \times 10^{8}$ | $180 \pm 9$ | $300 \pm 20$ | 10 | Gultekin + 2009a |
| NGC 4258 | SABbc | 7.276 | 5.464 | -23.85 | -21.55 | -23.71 | 0.12 | 0.0 | -20.94 | -18.64 | -20.80 | 2.907 | 0.676 | $9.86 \pm 0.09$ | $3.78(3.74-3.82) \times 10^{7}$ | $115 \pm 10$ | $208 \pm 6$ | 30 | Section 3.3 |
| NGC 4342 | S0 | 22.913 | 9.023 | -22.78 | -22.40 | -21.48 | 0.70 | 0.01 | -19.50 | -19.11 | -18.19 | 3.287 | 0.932 | $10.31 \pm 0.09$ | $4.53(3.05-7.18) \times 10^{8}$ | $225 \pm 11$ |  | 10 | Cretton + 1999b |
| NGC 4526 | S0/ | 16.442 | 6.473 | -24.61 | -24.15 | -23.47 | 0.65 | 0.0 | -21.44 | -20.98 | -20.30 | 3.170 | 0.941 | $11.02 \pm 0.09$ | $4.51(3.48-5.91) \times 10^{8}$ | $222 \pm 11$ | $290 \pm 20$ | 40 | Davis +2013 |
| NGC 4564 | So | 15.942 | 7.937 | -23.09 | -22.65 | -21.88 | 0.67 | 0.0 | -20.06 | -19.62 | -18.85 | 3.028 | 0.899 | $10.38 \pm 0.09$ | $8.81(6.38-11.26) \times 10^{7}$ | $162 \pm 8$ |  | 1 | Schulze + 2011 |
| NGC 4594 | Sa | 9.872 | 4.625 | -25.36 | -25.28 | -22.55 | 0.925 | 0.01 | -22.38 | -22.30 | -19.57 | 2.980 | 0.934 | $11.47 \pm 0.09$ | $6.65(6.24-7.05) \times 10^{8}$ | $240 \pm 12$ | $360 \pm 10$ | 11 | Jardel + 2011 |
| NGC 4596 | SB0 | 16.534 | 7.463 | -23.64 | -22.21 | -23.29 | 0.27 | 0.0 | -20.72 | -19.30 | -20.38 | 2.913 | 0.920 | $10.20 \pm 0.09$ | $7.67(4.43-11.41) \times 10_{7}^{7}$ | $136 \pm 6$ | $230 \pm 30$ | 20 | Sarzi +2001 |
| NGC 7457 | S0 | 12.532 | 8.179 | -22.33 | -20.82 | -22.02 | 0.25 | 0.0 | -19.45 | -17.94 | -19.14 | 2.880 | 0.844 | $9.56 \pm 0.09$ | $0.90(0.36-1.43) \times 10^{7}$ | $67 \pm 3$ | $145 \pm$ | 11 | Schulze + 2011 |
| Galaxy | Sb | 0.00828 |  | -23.7 | -21.9 | -23.5 | 0.0 | 19 | -20.8 | -18.9 | -20.5 | 2.980 |  | $10.09 \pm 0.10$ | $4.30(3.94-4.66) \times 10^{6}$ | $105 \pm 20$ | $220 \pm 20$ | 10 | Genzel + 2010 |
| Circinus | SABb: | 2.82 | 4.71 | -22.85 | -21.55 | -22.47 | 0.0 | 0.30 | -19.80 | -18.49 | -19.41 | 3.052 | 0.410 | $9.63 \pm 0.14$ | $1.14(0.94-1.34) \times 10^{6}$ | $79 \pm 3$ | $155 \pm 5$ | 30 | Greenhill + 2003 |
| NGC 1068 | Sb | 15.99 | 5.788 | -25.23 | -24.25 | -24.66 | 0.0 | 0.41 | -22.23 | -20.92 | -21.84 | 3.000 | 0.710 | $10.92 \pm 0.10$ | $8.39(7.95-8.83) \times 10^{6}$ | $151 \pm 7$ | $283 \pm 9$ | 0 | Lodato+2003, Huré+2002 |
| NGC 1300 | SBbc | 21.59 | 7.564 | -24.11 | -21.71 | -23.98 | 0.0 | 0.11 | -21.32 | -18.92 | -21.19 | 2.791 | 0.653 | $9.84 \pm 0.10$ | $7.55(3.89-14.75) \times 10^{7}$ | $88 \pm 3$ | $220 \pm 10$ | 20 | Atkinson +2005 |
| NGC 2273 | SBa | 29.59 | 8.480 | -23.89 | -22.07 | -23.67 | 0.0 | 0.19 | -20.88 | -19.06 | -20.66 | 3.007 | 0.827 | $10.08 \pm 0.09$ | $8.61(8.15-9.07) \times 10^{6}$ | $125 \pm 9$ | $220 \pm 6$ | 30 | Kuo + 2011 |
| NGC 2748 | Sc | 23.49 | 8.723 | -23.13 | -20.56 | -23.02 | 0.0 | 0.094 | $-20.27$ | -17.70 | -20.16 | 2.862 | 0.707 | $9.41 \pm 0.10$ | $4.44(2.62-6.20) \times 10^{7}$ | $115 \pm 5$ | $150 \pm 10$ | 20 | Atkinson +2005 |
| NGC 2787 | SB0/a | 7.452 | 7.263 | -22.14 | -21.06 | -21.64 | 0.11 | 0.26 | -19.10 | -18.02 | -18.60 | 3.038 | 0.944 | $9.78 \pm 0.09$ | $4.07(3.55-4.47) \times 10^{7}$ | $189 \pm 9$ | $226 \pm 10$ | 20 | Sarzi +2001 |
| NGC 3227 | SBa | 23.752 | 7.639 | -24.25 | -21.83 | -24.12 | 0.0 | 0.108 | -21.55 | -19.13 | -21.43 | 2.696 | 0.800 | $9.99 \pm 0.09$ | $2.10(0.98-2.79) \times 10^{7}$ | $133 \pm 12$ | $250 \pm 10$ | 10 | Davies + 2006 |
| NGC 3368 | SABab | 10.622 | 6.320 | -23.99 | -22.48 | -23.68 | 0.0 | 0.25 | -21.14 | -19.63 | -20.82 | 2.854 | 0.838 | $10.26 \pm 0.09$ | $7.66(6.13-9.19) \times 10^{6}$ | $125 \pm 6$ | $204 \pm 5$ | 10 | Nowak + 2010 |
| NGC 3384 | SB0 | 11.492 | 6.750 | -23.65 | -22.56 | -23.15 | 0.0 | 0.37 | -20.55 | -19.46 | -20.05 | 3.105 | 0.906 | $10.34 \pm 0.09$ | $1.08(0.59-1.57) \times 10^{7}$ | $146 \pm 7$ | $160 \pm 10$ | 11 | Schulze + 2011 |
| NGC 3393 | SABa | 49.29 | 9.059 | -24.45 | -23.03 | -24.11 | 0.0 | 0.27 | -21.48 | -20.05 | -21.14 | 2.968 | 0.813 | $10.48 \pm 0.09$ | $1.57(0.58-2.55) \times 10^{7}$ | $148 \pm 10$ |  | 30 | Kondratko +2008, Huré+201 |
| NGC 3489 | SABa | 11.982 | 7.370 | -23.29 | -22.15 | -22.82 | 0.11 | 0.24 | -20.17 | -19.03 | -19.70 | 3.120 | 0.815 | $10.11 \pm 0.09$ | $5.94(5.11-6.78) \times 10^{6}$ | $113 \pm 4$ |  | 10 | Nowak + 2010 |
| NGC 3945 | SB0 | 19.59 | 7.526 | -23.93 | -22.88 | -23.41 | 0.04 | 0.34 | -20.95 | -19.90 | -20.43 | 2.980 | 0.925 | $10.50 \pm 0.09$ | $8.8(0.00-25.5) \times 10^{6}$ | $192 \pm 10$ |  | 10 | Gültekin + 2009b |
| NGC 4388 | SBbc | 16.534 | 8.004 | -23.10 | -20.55 | -22.99 | 0.0 | 0.096 | -20.14 | -17.60 | -20.03 | 2.955 | 0.711 | $9.41 \pm 0.10$ | $7.31(7.13-7.48) \times 10^{6}$ | $99 \pm 10$ | $200 \pm 10$ | 30 | Kuo + 2011 |
| NGC 4736 | Sab | 5.002 | 5.106 | -23.39 | -22.29 | -22.91 | 0.0 | 0.36 | -20.68 | -19.58 | -20.20 | 2.710 | 0.735 | $10.13 \pm 0.10$ | $6.77(5.21-8.33) \times 10^{6}$ | $120 \pm 6$ | $181 \pm 10$ | 10 | Gebhardt + 2013 |
| NGC 4826 | Sab | 7.272 | 5.330 | -23.99 | -22.24 | -23.75 | 0.0 | 0.20 | -20.98 | -19.23 | -20.74 | 3.009 | 0.803 | $10.14 \pm 0.09$ | $1.56(1.17-1.95) \times 10^{6}$ | $104 \pm 3$ | $155 \pm 5$ | 10 | Gebhardt + 2013 |
| NGC 4945 | Scd | 3.58 | 4.438 | -23.38 | -20.50 | -23.30 | 0.0 | 0.07 | -20.58 | -17.70 | -20.50 | 2.801 | 1.20 | $9.35 \pm 0.12$ | $1.35(0.87-2.03) \times 10^{6}$ | $134 \pm 20$ | $174 \pm 10$ | 30 | Greenhill + 1997b |
| NGC 6264 | SBb | 147.69 | 11.407 | -24.5 | -22.6 | -24.3 | 0.0 | 0.17 |  |  |  |  |  | $10.36 \pm 0.09$ | $3.08(3.04-3.12) \times 10^{7}$ | $158 \pm 15$ |  | 30 | Kuo + 2011 |
| NGC 6323 | SBab | 113.49 | 10.530 | -24.80 | -21.55 | -24.75 | 0.0 | 0.050 |  |  |  |  |  | $9.94 \pm 0.09$ | $1.01(1.00-1.03) \times 10^{7}$ | $158 \pm 26$ |  | 30 | Kuo + 2011 |
| NGC 7582 | SBab | 22.39 | 7.316 | $-24.43$ | -21.96 | -24.31 | 0.0 | 0.103 | -21.78 | -19.31 | -21.66 | 2.649 | 0.738 | $10.02 \pm 0.10$ | $5.51(4.56-6.81) \times 10^{7}$ | $156 \pm 19$ | $226 \pm 10$ | 20 | Wold +2006 |
| IC 2560 | SBbc | 37.2 | 8.694 | -24.19 | -22.05 | -24.02 | 0.0 | 0.14 | -21.65 | -19.51 | -21.48 | 2.541 | 0.886 | $10.12 \pm 0.09$ | $5.01(0.00-5.72) \times 10^{6}$ | $141 \pm 10$ | $196 \pm 3$ | 30 | Yamauchi +2012 |
| UGC 3789 | SABab | 49.9 | 9.510 | -24.03 | -22.79 | -23.61 | 0.0 | 0.32 | -21.13 | -19.89 | -20.71 | 2.9 | 0.86 | $10.39 \pm 0.09$ | $9.65(8.10-11.20) \times 10^{6}$ | $107 \pm 12$ | $273 \pm 20$ | 30 | Kuo+2011, Huré+2011 |

 Column 2 is the galaxy Hubble type, mostly from RC3 with a few corrections from Kormendy \& Bender (2013b). Column 3 is the assumed distance (see notes to Table 2). For our Galaxy, Circinus, NGC 4945, and UGC 3789, see supplemental notes on individual galaxies


Table 3 Supermassive black holes detected dynamically in spiral and S0 galaxies galaxies ( 21 with classical bulges; 22 with pseudobulges; December 2012)

| Object (1) |  | Distanc (Mpc) (3) |  | $M_{K s T}$ | $M_{K s, \text { bulge }}$ <br> (6) | $M_{K s, \text { disk }}$ <br> (7) |  | $P B / T$ <br> (9) | $M_{V T}$ <br> (10) | $M_{V, \text { bulge }}$ <br> (11) | $M_{V, \text { disk }}$ <br> (12) | $\left(V-K_{s}\right)_{0}$ <br> (13) | $(B-V)_{0}$ <br> (14) | $\begin{gathered} \log M_{\text {bulge }} \\ \left(M_{\odot}\right) \\ (1.5) \end{gathered}$ | $\begin{gathered} M_{\bullet}\left(\text { low } M_{\bullet}-\text { high } M_{\bullet}\right) \\ \left(M_{\odot}\right) \\ (16) \end{gathered}$ | $\begin{gathered} \sigma_{e} \\ \left(\mathrm{~km} \mathrm{~s}^{-1}\right) \\ (17) \end{gathered}$ |  | Flags M $M$ • (19) | Source <br> (20) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| M 31 <br> M 81 <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC <br> NGC |  | $\begin{aligned} & 19.5 \\ & 20.3 \\ & 19.1 \\ & \hline 20.5 \\ & 19.4 \\ & 21.8 \\ & 21.1 \end{aligned}$ |  | $\begin{aligned} & 2.9 \\ & 2.9 \\ & 2.8 \\ & 2.9 \\ & 3.0 \\ & 3.0 \\ & 2.7 \end{aligned}$ |  |  |  |  | $\begin{array}{r} 11 . \\ 10 . \\ 9 . \\ 10 . \\ 9 . \\ 10 . \end{array}$ | $\begin{aligned} & 7 \pm \\ & 0 \pm \\ & 6 \pm \\ & 9 \pm \\ & 3 \pm \\ & 2 \pm \\ & 4 \pm \end{aligned}$ | $\begin{aligned} & 09 \\ & 09 \\ & 09 \\ & \hline 10 \\ & 14 \\ & 10 \\ & 10 \end{aligned}$ | $\begin{aligned} & 6.6 \\ & 7.6 \\ & 0.9 \\ & 4.3 \\ & 1.1 \\ & 8.3 \\ & 7.5 \end{aligned}$ | $\begin{aligned} & (6.24 \\ & (4.4 \\ & (0.36 \\ & \hline(3.94 \\ & (0.94 \\ & (7.95 \\ & (3.85 \end{aligned}$ | $\begin{array}{lr} - & 7 \\ - & 11 \\ - & 1 \\ \hline- & 4 . \\ - & 1 . \\ - & 8 . \\ - & 14 . \end{array}$ | 5) $\times 10^{8}$ <br> 1) $\times 10^{7}$ <br> 3) $\times 10^{7}$ <br> 6) $\times 10^{6}$ <br> 4) $\times 10^{6}$ <br> 3) $\times 10^{6}$ <br> 5) $\times 10^{7}$ | $\begin{aligned} 240 & \pm \\ 136 & \pm \\ 67 & \pm \\ 105 & \pm \\ 79 & \pm \\ 151 & \pm \\ 88 & \pm \end{aligned}$ | 12 6 3 20 3 7 3 | $\begin{aligned} & 360 \\ & 230 \\ & 145 \\ & 220 \\ & 155 \\ & 283 \\ & 220 \end{aligned}$ | $\begin{aligned} & \pm 10 \\ & \pm 30 \\ & \pm \quad 6 \\ & \pm 20 \\ & \pm \quad 5 \\ & \pm 9 \\ & \pm 10 \end{aligned}$ |
| Galaxy | Sbc | 0.008 |  | -23.7 | -21.9 | -23.5 | 0.0 | 0.19 | -20.8 | -18.9 | -20.5 | 2.980 |  | $10.09 \pm 0.10$ | $4.30(3.94-4.66) \times 10^{6}$ | $105 \pm 20$ | $220 \pm 20$ | 10 | Genzel + 2010 |
| Circinus | SABb: | 2.82 | 4.71 | -22.85 | -21.55 | -22.47 | 0.0 | 0.30 | -19.80 | -18.49 | -19.41 | 3.052 | 0.410 | $9.63 \pm 0.14$ | $1.14(0.94-1.34) \times 10^{6}$ | $79 \pm 3$ | $155 \pm 5$ | 30 | Greenhill +2003 |
| NGC 1068 | Sb | 15.99 | 5.788 | -25.23 | -24.25 | -24.66 | 0.0 | 0.41 | -22.23 | -20.92 | -21.84 | 3.000 | 0.710 | $10.92 \pm 0.10$ | $8.39(7.95-8.83) \times 10^{6}$ | $151 \pm 7$ | $283 \pm 9$ | 30 | Lodato+2003, Huré+2002 |
| NGC 1300 | SBbc | 21.5 | 7.564 | -24.11 | -21.71 | -23.98 | 0.0 | 0.11 | -21.32 | -18.92 | -21.19 | 2.791 | 0.653 | $9.84 \pm 0.10$ | $7.55(3.89-14.75) \times 10^{7}$ | $88 \pm 3$ | $220 \pm 10$ | 20 | Atkinson +2005 |
| NGC 2273 | SBa | 29.5 | 8.480 | -23.89 | -22.07 | -23.67 | 0.0 | 0.19 | -20.88 | -19.06 | -20.66 | 3.007 | 0.827 | $10.08 \pm 0.09$ | $8.61(8.15-9.07) \times 10^{6}$ | $125 \pm 9$ | $220 \pm 6$ | 30 | Kuo + 2011 |
| NGC 2748 | Sc | 23.49 | 8.723 | -23.13 | -20.56 | -23.02 | 0.0 | 0.094 | -20.27 | -17.70 | -20.16 | 2.862 | 0.707 | $9.41 \pm 0.10$ | $4.44(2.62-6.20) \times 10^{7}$ | $115 \pm 5$ | $150 \pm 10$ | 20 | Atkinson +2005 |
| NGC 2787 | SB0/a | 7.45 | 7.263 | -22.14 | -21.06 | -21.64 | 0.11 | 0.26 | -19.10 | -18.02 | -18.60 | 3.038 | 0.944 | $9.78 \pm 0.09$ | $4.07(3.55-4.47) \times 10^{7}$ | $189 \pm 9$ | $226 \pm 10$ | 20 | Sarzi +2001 |
| NGC 3227 | SBa | 23.75 | 7.639 | -24.25 | -21.83 | -24.12 | 0.0 | 0.108 | -21.55 | -19.13 | -21.43 | 2.696 | 0.800 | $9.99 \pm 0.09$ | $2.10(0.98-2.79) \times 10^{7}$ | $133 \pm 12$ | $250 \pm 10$ | 10 | Davies + 2006 |
| NGC 3368 | SABab | 10.62 | 6.320 | -23.99 | -22.48 | -23.68 | 0.0 | 0.25 | -21.14 | -19.63 | -20.82 | 2.854 | 0.838 | $10.26 \pm 0.09$ | $7.66(6.13-9.19) \times 10^{6}$ | $125 \pm 6$ | $204 \pm 5$ | 10 | Nowak +2010 |
| NGC 3384 | SB0 | 11.49 | 6.750 | -23.65 | -22.56 | -23.15 | 0.0 | 0.37 | -20.55 | -19.46 | -20.05 | 3.105 | 0.906 | $10.34 \pm 0.09$ | $1.08(0.59-1.57) \times 10^{7}$ | $146 \pm 7$ | $160 \pm 10$ | 11 | Schulze +2011 |
| NGC 3393 | SABa | 49.29 | 9.059 | -24.45 | -23.03 | -24.11 | 0.0 | 0.27 | -21.48 | -20.05 | -21.14 | 2.968 | 0.813 | $10.48 \pm 0.09$ | $1.57(0.58-2.55) \times 10^{7}$ | $148 \pm 10$ |  | 30 | Kondratko+2008, Huré+2011 |
| NGC 3489 | SABa | 11.98 | 7.370 | -23.29 | -22.15 | -22.82 | 0.11 | 0.24 | -20.17 | -19.03 | -19.70 | 3.120 | 0.815 | $10.11 \pm 0.09$ | $5.94(5.11-6.78) \times 10^{6}$ | $113 \pm 4$ |  | 10 | Nowak +2010 |
| NGC 3945 | SB0 | 19.59 | 7.526 | -23.93 | -22.88 | -23.41 | 0.04 | 0.34 | -20.95 | -19.90 | $-20.43$ | 2.980 | 0.925 | $10.50 \pm 0.09$ | $8.8(0.00-25.5) \times 10^{6}$ | $192 \pm 10$ |  | 10 | Gültekin + 2009b |
| NGC 4388 | SBbc | 16.53 | 8.004 | -23.10 | -20.55 | -22.99 | 0.0 | 0.096 | -20.14 | -17.60 | -20.03 | 2.955 | 0.711 | $9.41 \pm 0.10$ | $7.31(7.13-7.48) \times 10^{6}$ | $99 \pm 10$ | $200 \pm 10$ | 30 | Kuo + 2011 |
| NGC 4736 | Sab | 5.002 | 5.106 | -23.39 | -22.29 | -22.91 | 0.0 | 0.36 | -20.68 | -19.58 | -20.20 | 2.710 | 0.735 | $10.13 \pm 0.10$ | $6.77(5.21-8.33) \times 10^{6}$ | $120 \pm 6$ | $181 \pm 10$ | 10 | Gebhardt + 2013 |
| NGC 4826 | Sab | 7.272 | 5.330 | -23.99 | -22.24 | -23.75 | 0.0 | 0.20 | -20.98 | -19.23 | -20.74 | 3.009 | 0.803 | $10.14 \pm 0.09$ | $1.56(1.17-1.95) \times 10^{6}$ | $104 \pm 3$ | $155 \pm 5$ | 10 | Gebhardt + 2013 |
| NGC 4945 | Scd | 3.58 | 4.438 | -23.38 | -20.50 | -23.30 | 0.0 | 0.07 | -20.58 | -17.70 | -20.50 | 2.801 | 1.20 | $9.35 \pm 0.12$ | $1.35(0.87-2.03) \times 10^{6}$ | $134 \pm 20$ | $174 \pm 10$ | 30 | Greenhill +1997 b |
| NGC 6264 | SBb | 147.69 | 11.407 | -24.5 | -22.6 | -24.3 | 0.0 | 0.17 |  |  | ... | ... |  | $10.36 \pm 0.09$ | $3.08(3.04-3.12) \times 10^{7}$ | $158 \pm 15$ |  | 30 | Kuo + 2011 |
| NGC 6323 | SBab | 113.49 | 10.530 | -24.80 | -21.55 | -24.75 | 0.0 | 0.050 |  |  |  |  |  | $9.94 \pm 0.09$ | $1.01(1.00-1.03) \times 10^{7}$ | $158 \pm 26$ |  | 30 | Kuo + 2011 |
| NGC 7582 | SBab | 22.39 | 7.316 | -24.43 | -21.96 | -24.31 | 0.0 | 0.103 | -21.78 | -19.31 | -21.66 | 2.649 | 0.738 | $10.02 \pm 0.10$ | $5.51(4.56-6.81) \times 10^{7}$ | $156 \pm 19$ | $226 \pm 10$ | 20 | Wold + 2006 |
| IC 2560 | SBbc | 37.29 | 8.694 | -24.19 | -22.05 | -24.02 | 0.0 | 0.14 | -21.65 | -19.51 | -21.48 | 2.541 | 0.886 | $10.12 \pm 0.09$ | $5.01(0.00-5.72) \times 10^{6}$ | $141 \pm 10$ | $196 \pm 3$ | 30 | Yamauchi +2012 |
| UGC 3789 | SABab | 49.9 | 9.510 | -24.03 | -22.79 | -23.61 | 0.0 | 0.32 | -21.13 | -19.89 | -20.71 | 2.9 | 0.86 | $10.39 \pm 0.09$ | $9.65(8.10-11.20) \times 10^{6}$ | $107 \pm 12$ | $273 \pm 20$ | 30 | Kuo+2011, Huré+2011 |

 Column 3 is the assumed distance (see notes to Table 2). For our Galaxy, Circinus, NGC 4945, and UGC 3789, see supplemental notes on individual galaxies


 galaxy bulge


## M• $-\sigma$ Relation

## $M_{\bullet}-\sigma$ Relation

$$
\frac{M_{\bullet}}{10^{9} M_{\odot}}=\left(0.309_{-0.033}^{+0.037}\right)\left(\frac{\sigma}{200 \mathrm{~km} \mathrm{~s}^{-1}}\right)^{4.38 \pm 0.29} \text { intrinsic scatter }=0.28
$$

## $M_{\bullet}-M_{\text {bulge }}$ Relation

## $M_{\bullet}-M_{\text {bulge }}$ Relation

$$
\frac{M_{\bullet}}{10^{9} M_{\odot}}=\left(0.49_{-0.05}^{+0.06}\right)\left(\frac{M_{\text {bulge }}}{10^{11} M_{\odot}}\right)^{1.16 \pm 0.08} ; \text { intrinsic scatter }=0.29 \text { dex. }
$$

## $M_{\bullet}-M_{\text {bulge }}$ Relation

$$
\frac{M_{\bullet}}{10^{9} M_{\odot}}=\left(0.49_{-0.05}^{+0.06}\right)\left(\frac{M_{\text {bulge }}}{10^{11} M_{\odot}}\right)^{1.16 \pm 0.08} ; \text { intrinsic scatter }=0.29 \text { dex. }
$$



## Virial BH Masses for Type 1 AGNs

$$
M_{\mathrm{virial}}=f R V^{2} / G
$$

$f$ geometric fudge factor
$R$ BLR radius
$V$ BLR velocity dispersion

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## Virial BH Masses for Type 1 AGNs



Peterson et al. (2002)
$M_{\text {virial }}=f R V^{2} / G$
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## Virial BH Masses for Type 1 AGNs

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Ho \& $\operatorname{Kim}(2014,2015)$


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Ho \& $\operatorname{Kim}(2014,2015)$


## Virial BH Masses for Type 1 AGNs

$$
M_{\text {virial }}=f R V^{2} / G
$$

$f$ geometric fudge factor $R$ BLR radius
$V$ BLR velocity dispersion
M. can be estimated to an accuracy of $\sim 0.3-0.5 \mathrm{dex}$ for $z \approx 0-6$

## Correlation Between Black Hole Mass and Bulge Mass



## Correlation Between Black Hole Mass and Bulge Mass







## NGC 205

M32

## NGC 205

M32

M31 G1
(Mayall II)



Mortlock et al. (2011)


J. Wise $\mathcal{E}$ T. Abel





Filippenko \& Ho (2003); Barth et al. (2004); Ho (2008, ARA\&A)


Filippenko \& Ho (2003); Barth et al. (2004); Ho (2008, ARA\&A)


NGC 4395
Sdm

NGC 4395

Sdm

$$
M_{0}=10^{4}-10^{5} M_{\odot}
$$

## POX 52

Sph or dE

## POX 52

Sph or $d E$

$$
M_{\bullet}=1.6 \times 10^{5} M_{\odot}
$$



Greene \& Ho (2004, 2007a,b); Dong, Ho et al. (2012)


Greene \& Ho (2004, 2007a,b); Dong, Ho et al. (2012)


## HST/ACS

Greene, Ho \& Barth (2008); Jiang, Greene \& Ho (2011a, b)


Moran et al. (2014); cf. Barth, Greene E Ho (2008)


$$
M_{g}>-18 \mathrm{mag} \quad M_{*}=10^{8}-10^{9} M_{\odot}
$$












## Recent Updates

$\bigcirc$ Central BHs detected from $10^{4}-10^{10} M_{\odot}$
$\bigcirc$ All bulges contain BHs, but not all BHs live in bulges
$\bigcirc M_{\bullet} \sim M_{\text {bulge }}^{1.2} \quad\left\langle M_{\bullet} / M_{\text {bulge }}\right\rangle \sim 0.5 \%$
$\bigcirc M_{\bullet} \propto \sigma^{4.4}$
$Q_{\bullet}-\sigma$ and $M_{\bullet}-M_{\text {bulge }}$ relations have similar scatter
$\bigcirc$ Scaling relations only tight for classical bulges and Es
Scaling relations already in place for high-z QSOs
© Mild evolution only for most massive BHs
$\bigcirc$ AGN feedback effective only for classical bulges and Es
Sriverath

## Opportunities with ALMA



BH masses using nuclear (cold) gas disks

ISM content of quasars at all redshifts
$\bigcirc$ Dynamical masses of quasar host galaxies (CO ladder, [C II]) Gas distribution and kinematics


## 

## Thirty Meter Telescope

## Future Directions with TMT




## Future Directions with TMT

© Direct measurement of low-mass BHs in dwarf galaxies
$\bigcirc$ Direct measurement of BH masses in high-z inactive galaxies
$\bigcirc$ Direct measurement of BH-host scaling relations at high-z
© Calibration of BH masses in reverberation-mapped AGNs

- Stellar orbital structure of centers of BCGs, constrain growth mechanism of most massive BHs

