

Circumbinary Ring, Circumstellar Disks and Accretion in the Binary System UY Aurigae

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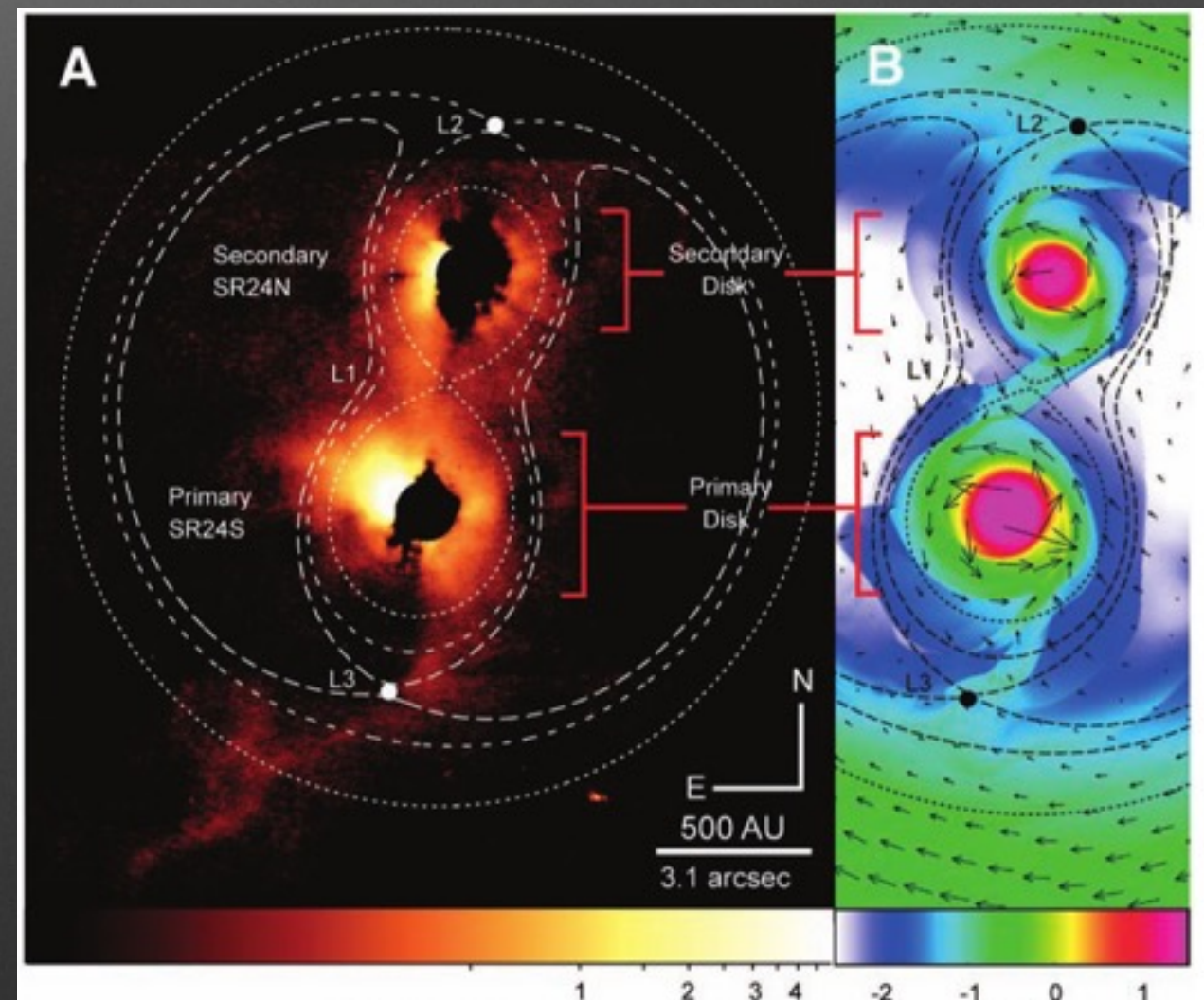
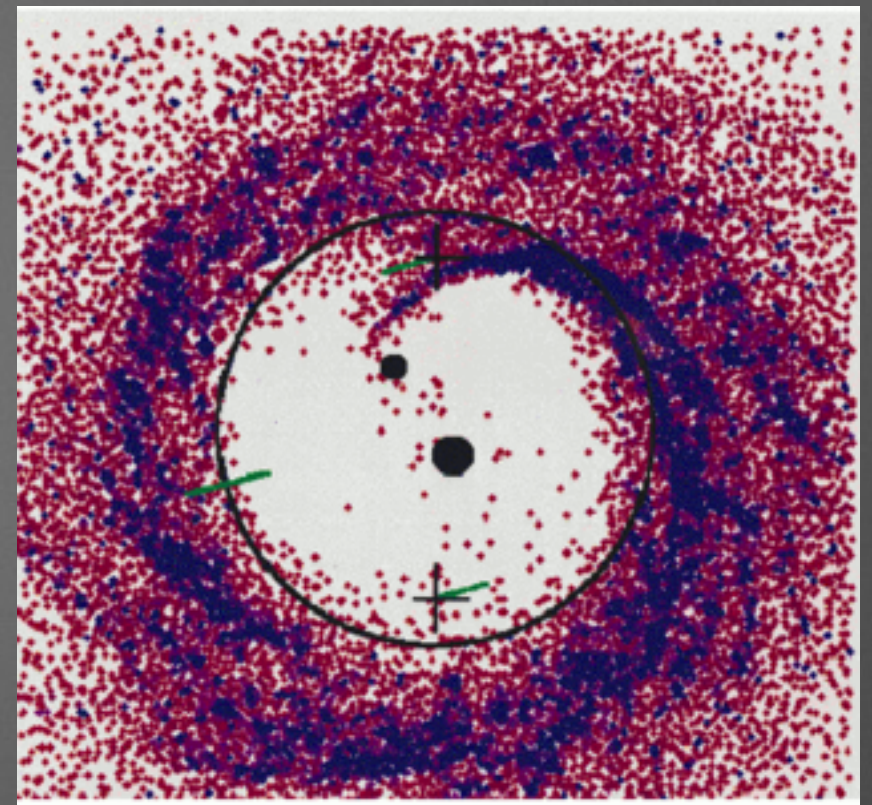
Tang, Dutrey, Guilloteau et al. 2014, ApJ

Image credit: NAOJ based on Pyo+2014

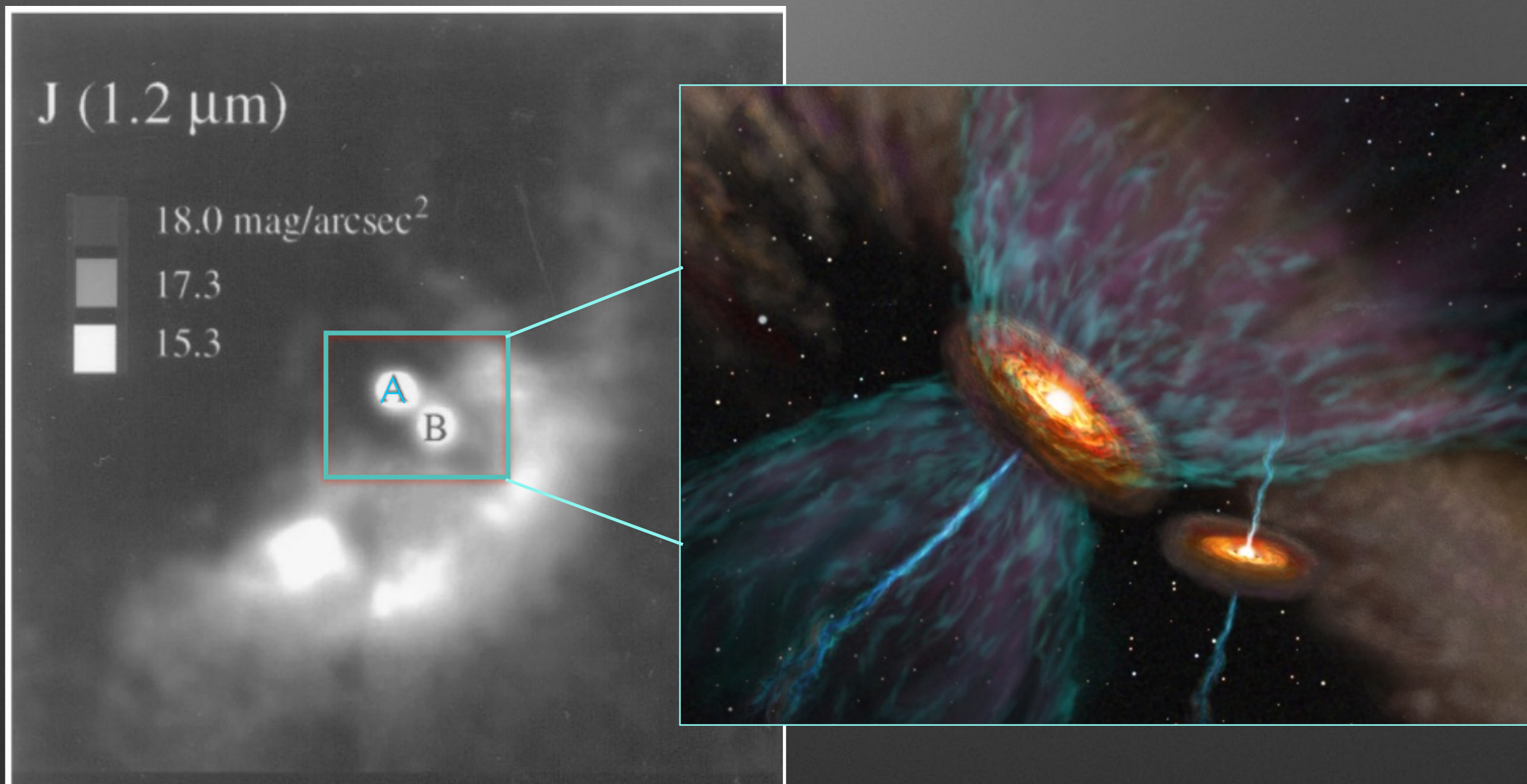


theoretical predictions in the binary system

- In binary system, both circumstellar (CS) disks and circumbinary (CB) disk are expected. CB disk feeds the binary through streams (Artymowicz+1996)
- Accretion rate is larger toward the primary (Ochi 2005) or secondary (Bate 1997) or even long time and short time variability (Hanawa 2010)



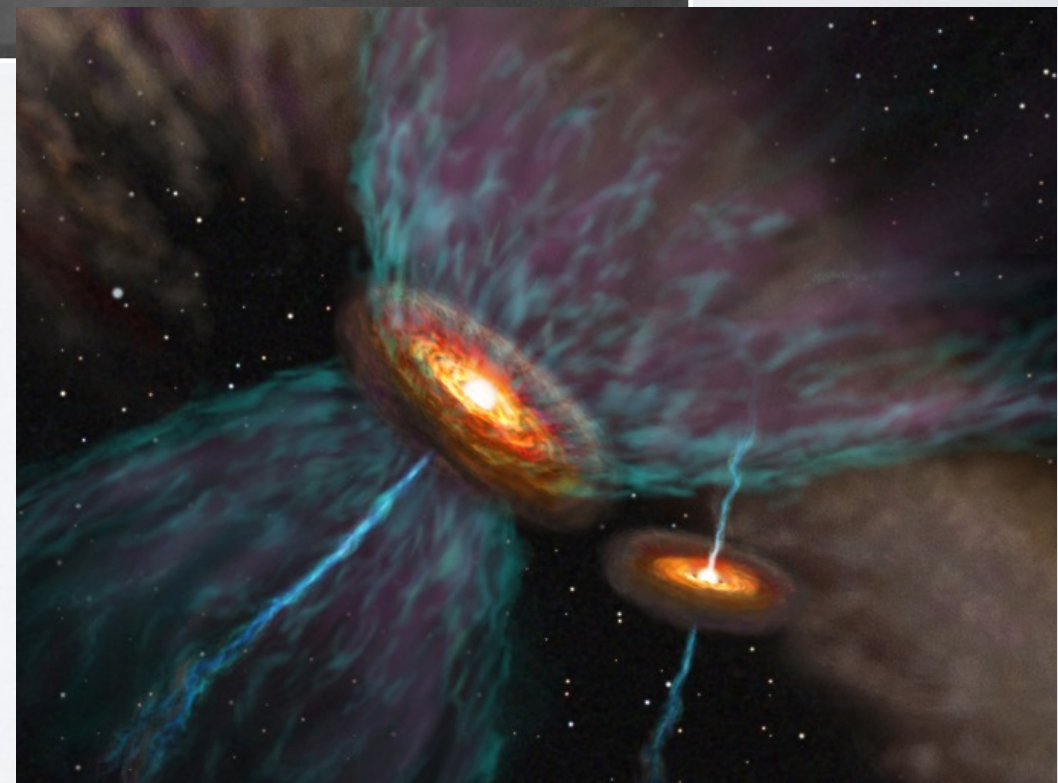
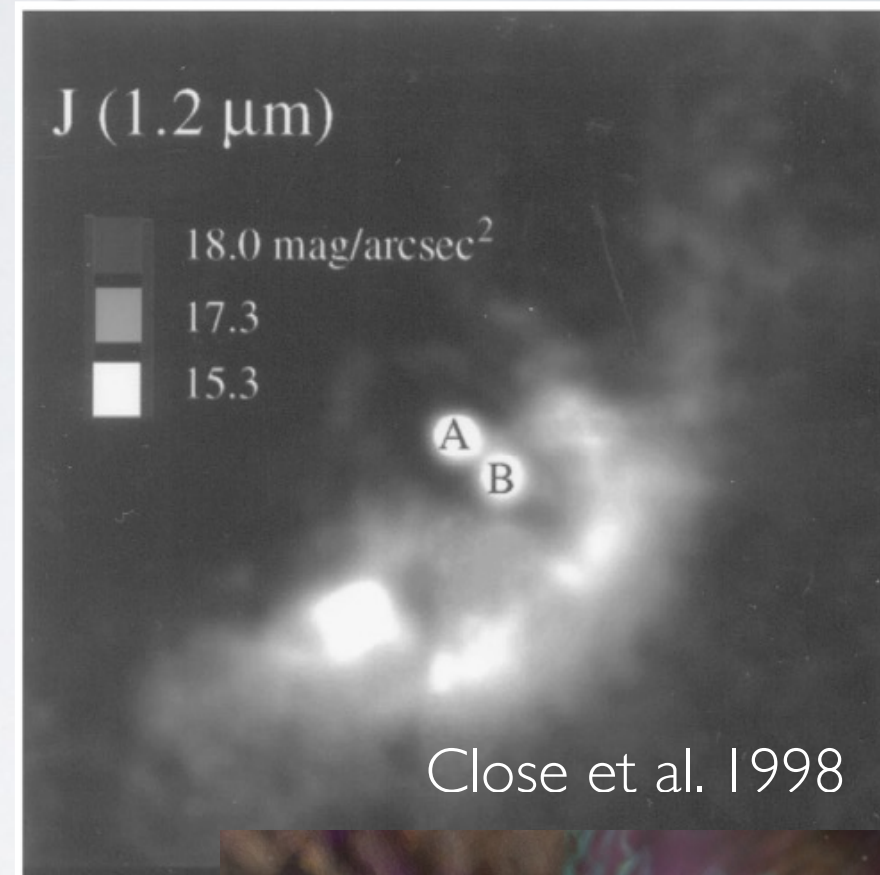
T Tauri binary UY Aur

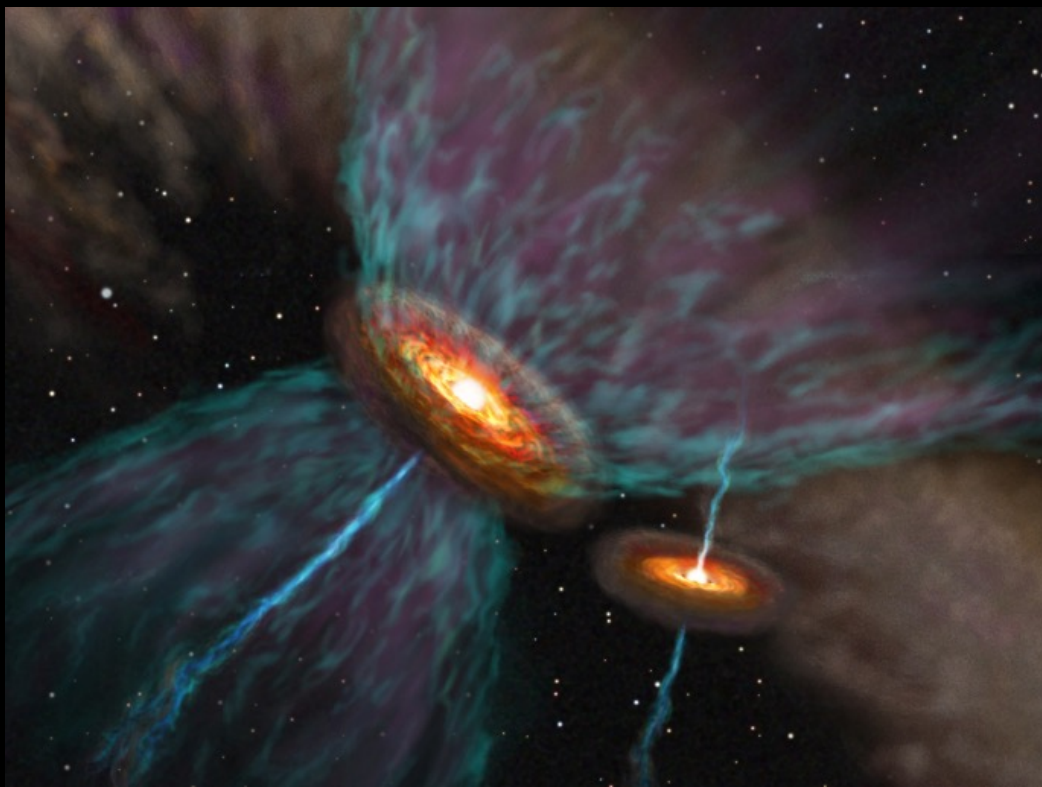


Close et al. (1998)

T Tauri binary UY Aur

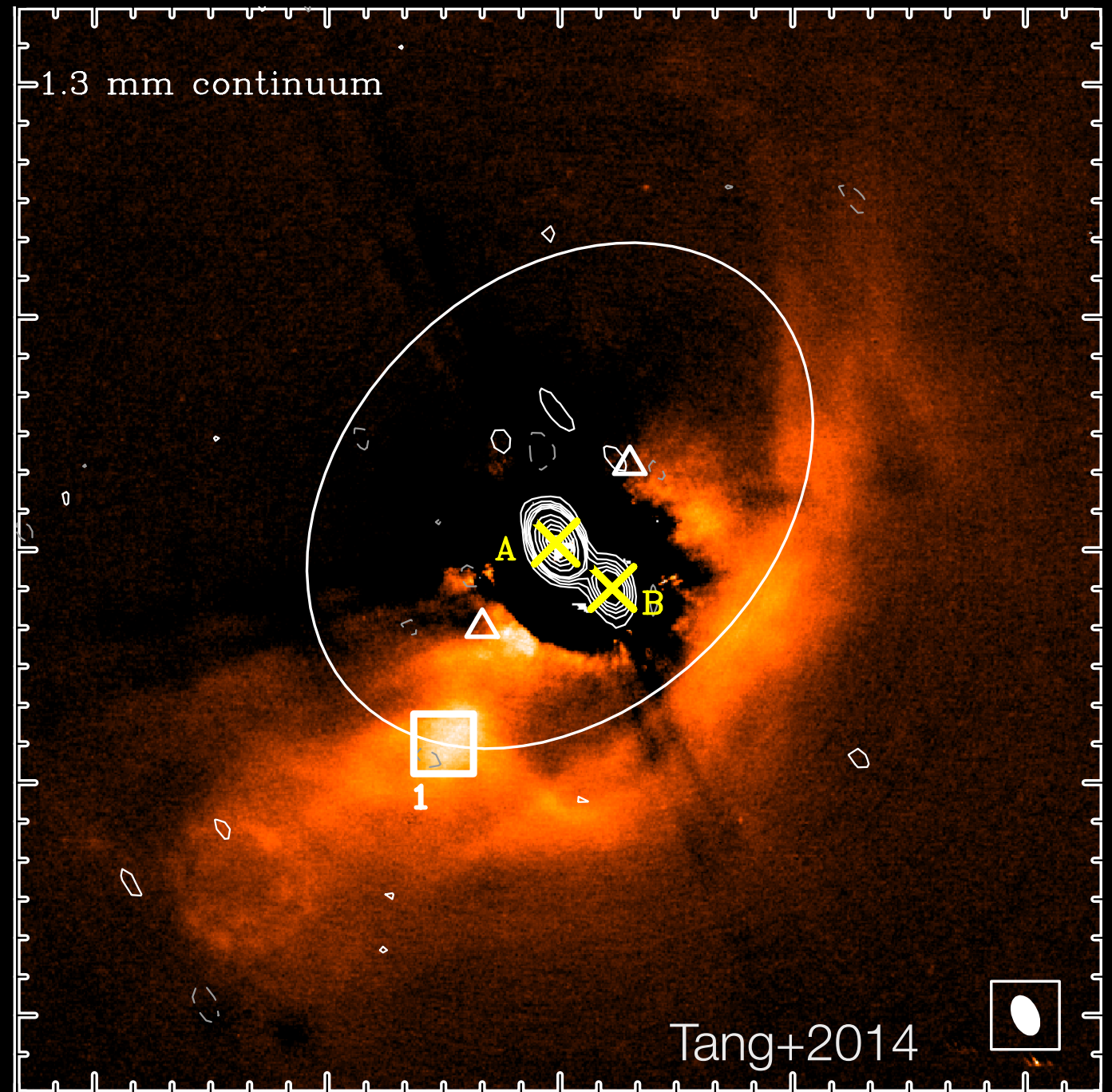
- T-Tauri binary system with a separation 0.88'' (120 AU)
- the primary:
 - K7 (0.95) or M0 (0.6) M_{\odot}
 - brightness is relatively stable
 - Accretion rate: $\sim 2 \times 10^{-8} M_{\odot}/\text{yr}$
 - pristine ISM-like dust
- the secondary:
 - M0 (0.6) or M2.5 (0.25) M_{\odot}
 - **brightness varies by 5 mag. in the R band over 50 yrs**
 - Accretion rate: $\sim 6 \times 10^{-8} M_{\odot}/\text{yr}$
 - silicate dust



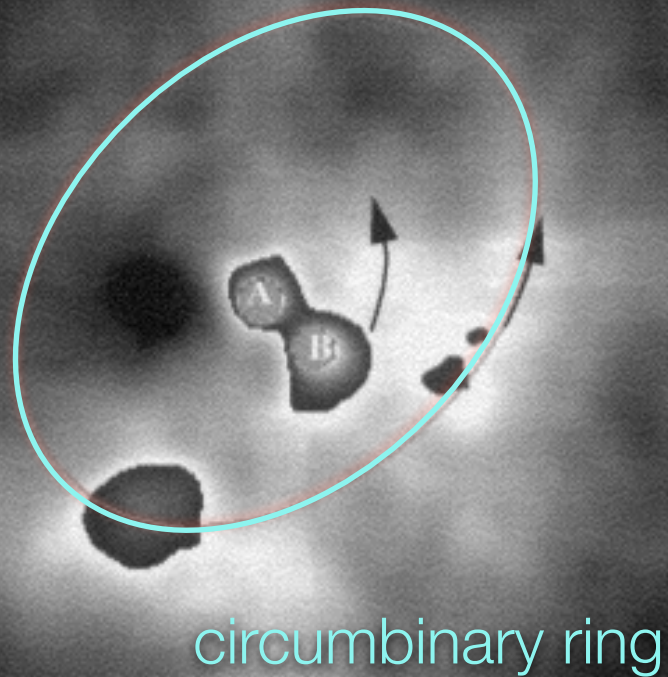


Circumstellar disks

- detected and resolved at 1.3 mm
 - Primary: 0.22'' (30 AU)
 - Secondary: 0.34'' (50 AU)
- Spectral index $\alpha \sim 1.6$ between 6 cm and 0.88 mm
 - a combination of free-free emission with thermal dust emission



K' (2.1 μm)

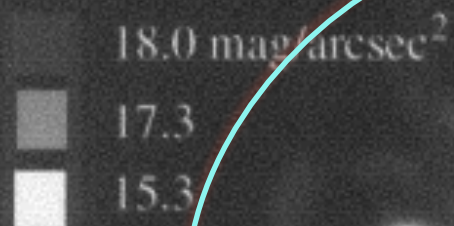


Circumbinary material

In scatter light

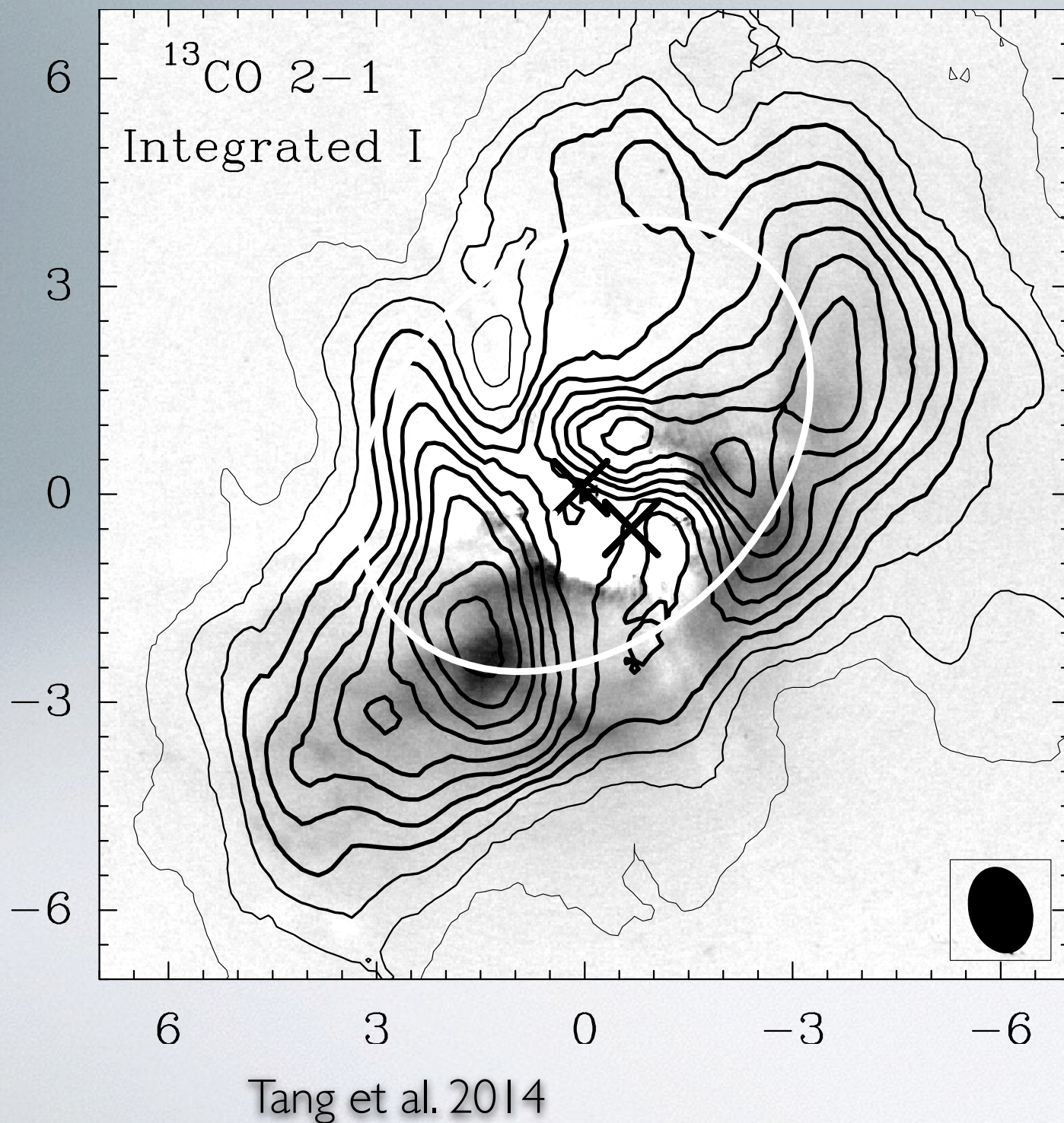
- A circumbinary ring is identified in K' band
 - inclination = 42°
 - PA = 135°
- “spiral-like” structures
 - two accreting spirals toward the binary
- an extended armlike structure

J (1.2 μm)

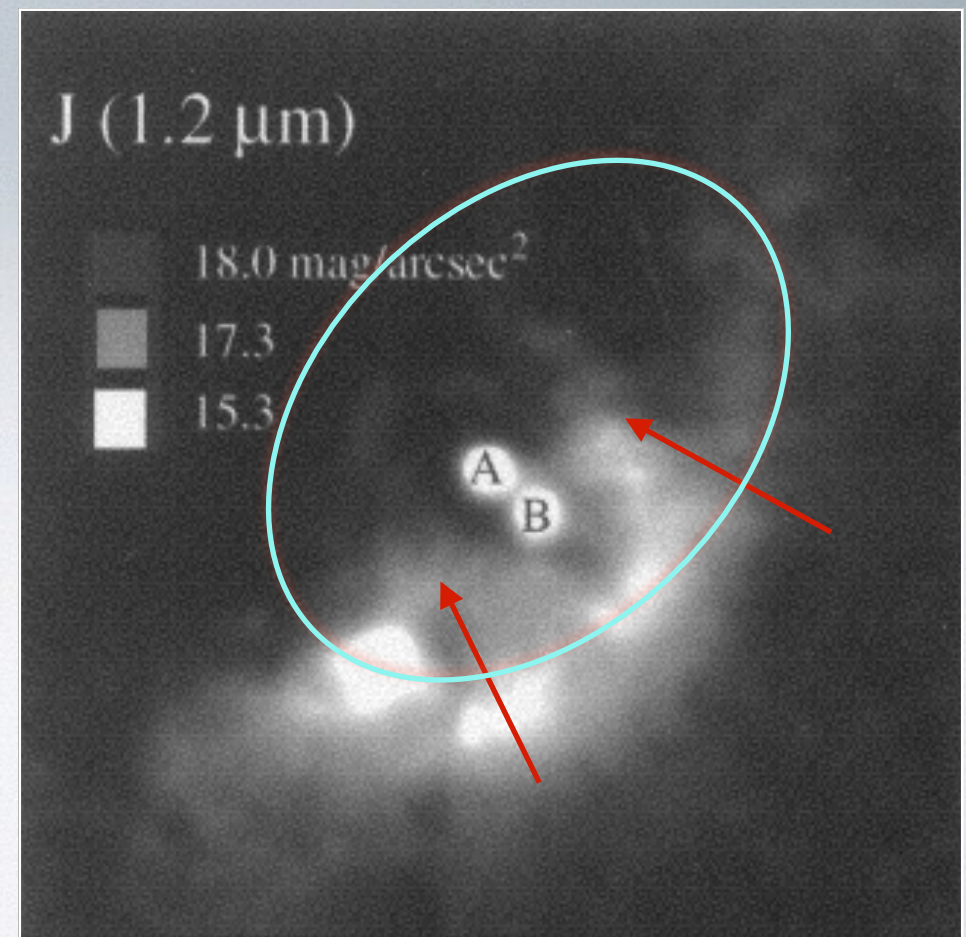


Circumbinary material:

New PdBI ^{13}CO 2-1 observations (contour)



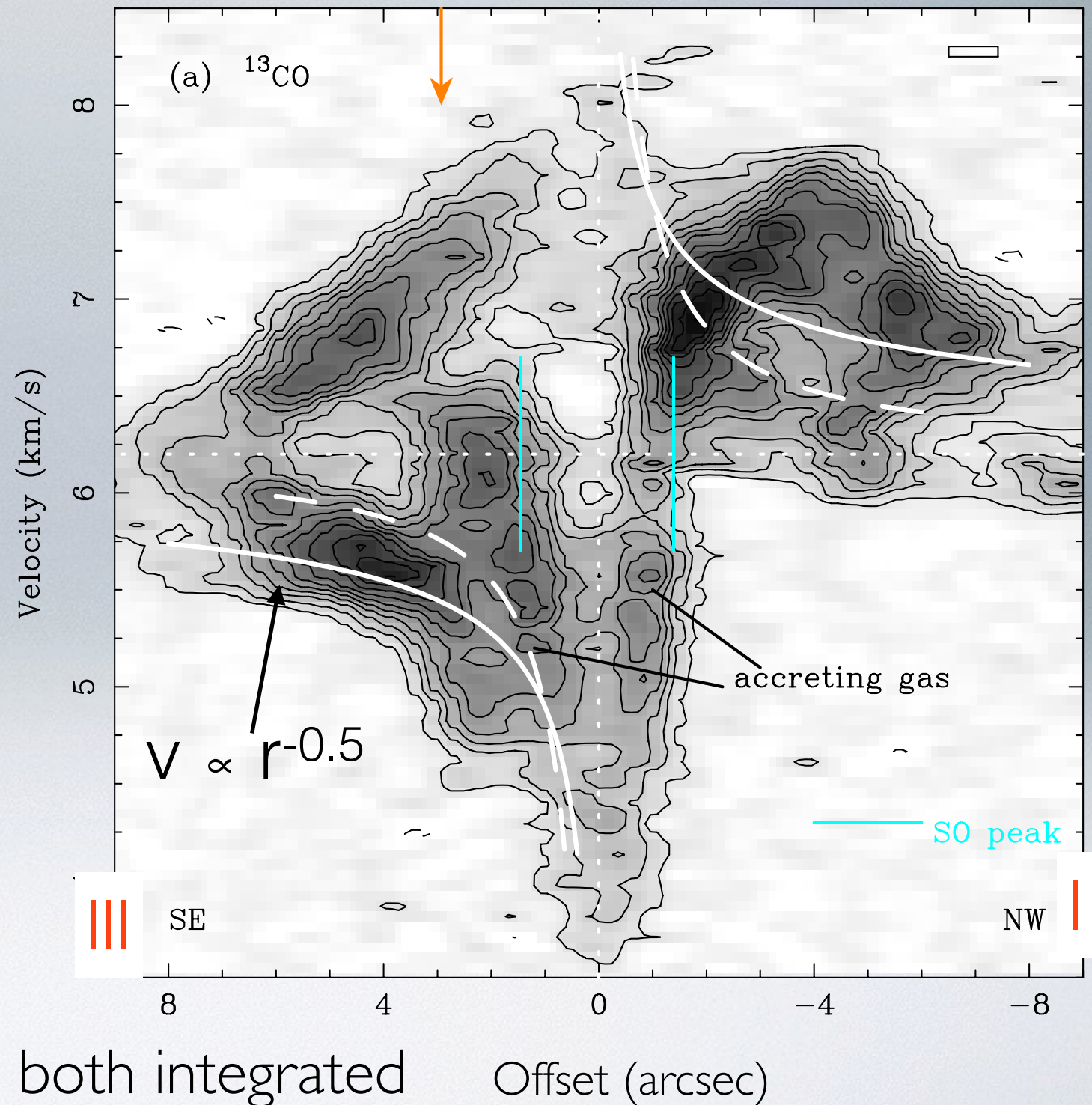
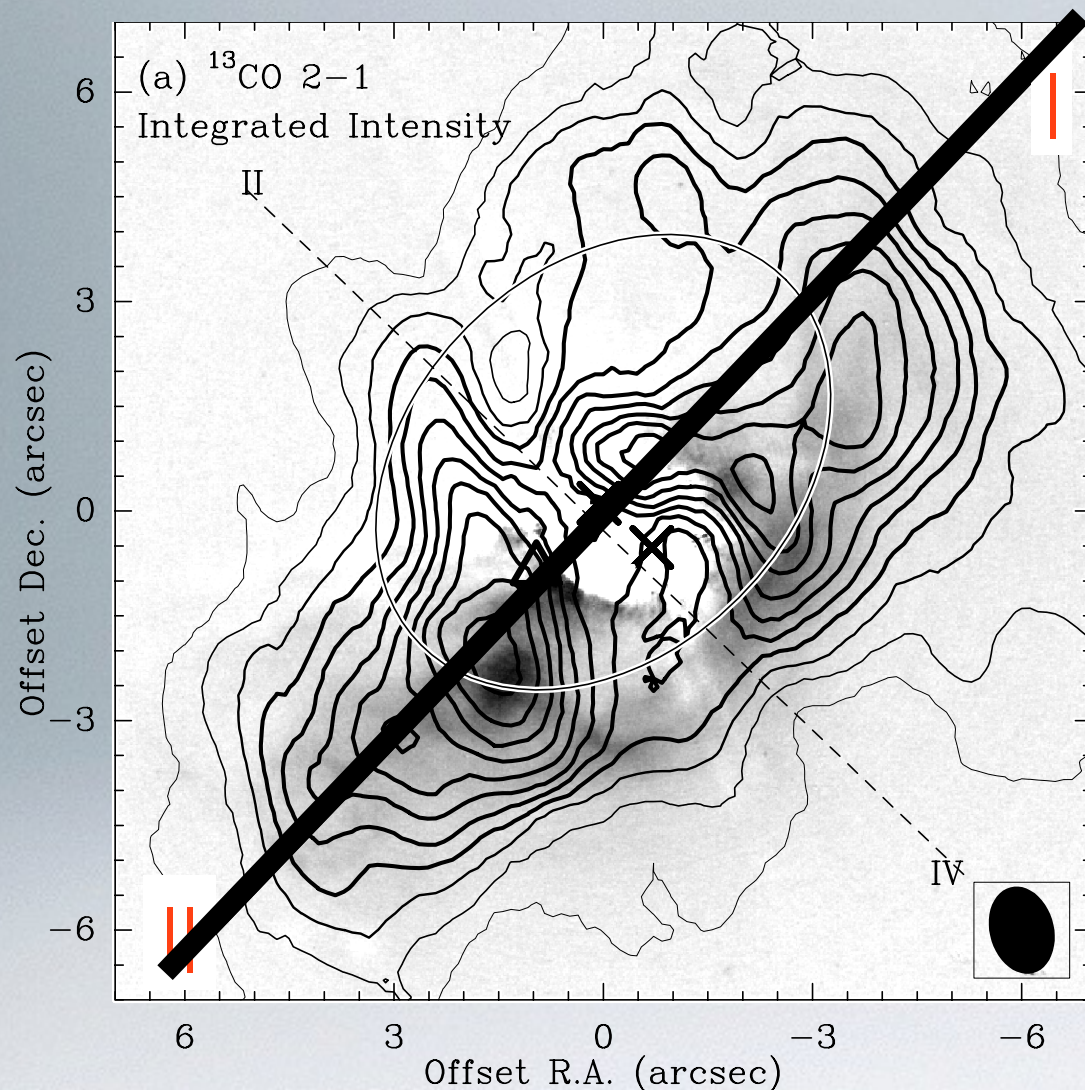
- Circumbinary disk/ring
- spiral-like streamers (accreting spirals)
- extended arm-like structure



Circumbinary material:

New PdBI ^{13}CO 2-1 observations

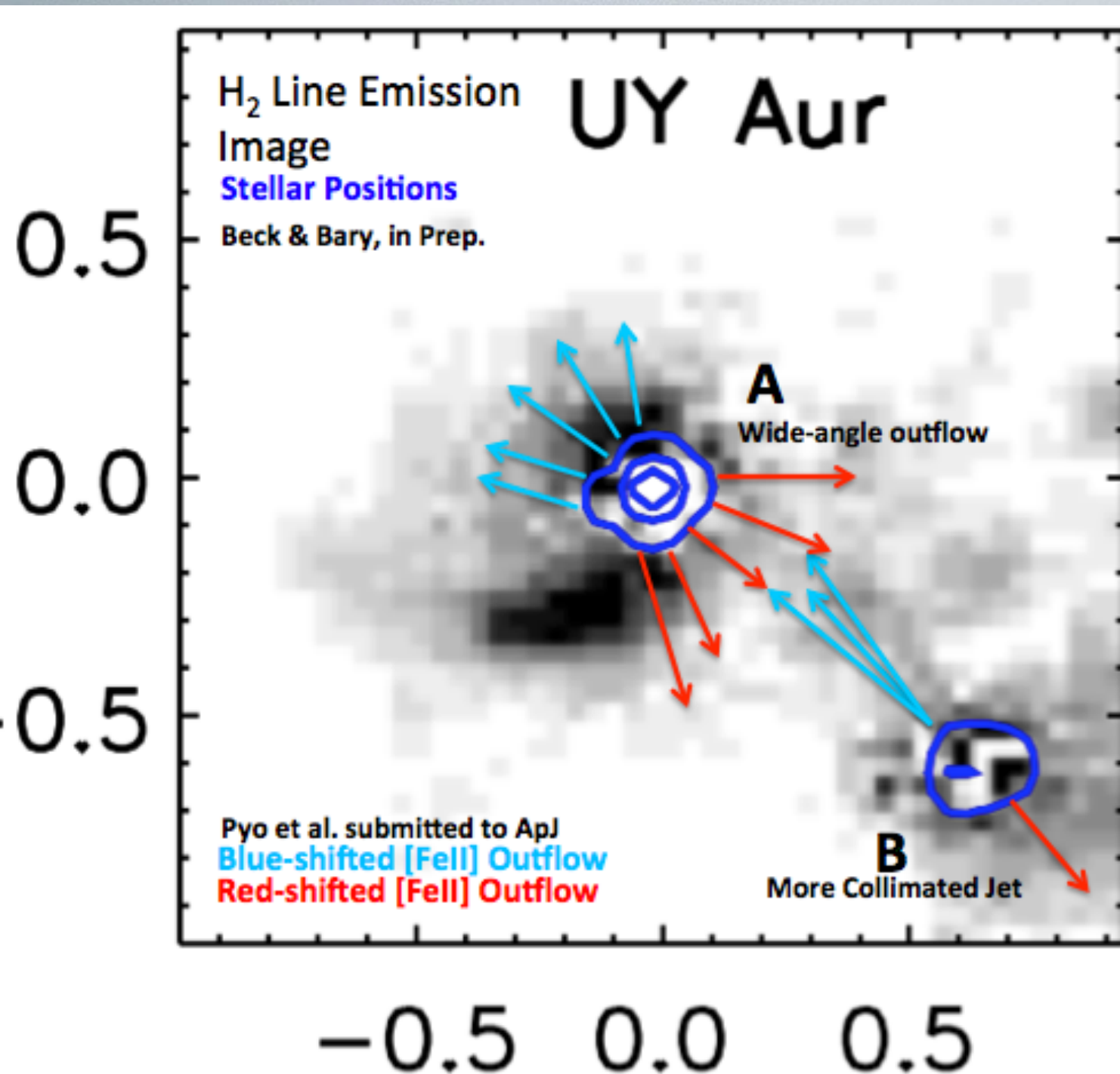
Tang et al. 2014



- kinematics can be analyzed
- evidence of accreting spirals in both integrated intensity map and the position velocity diagram

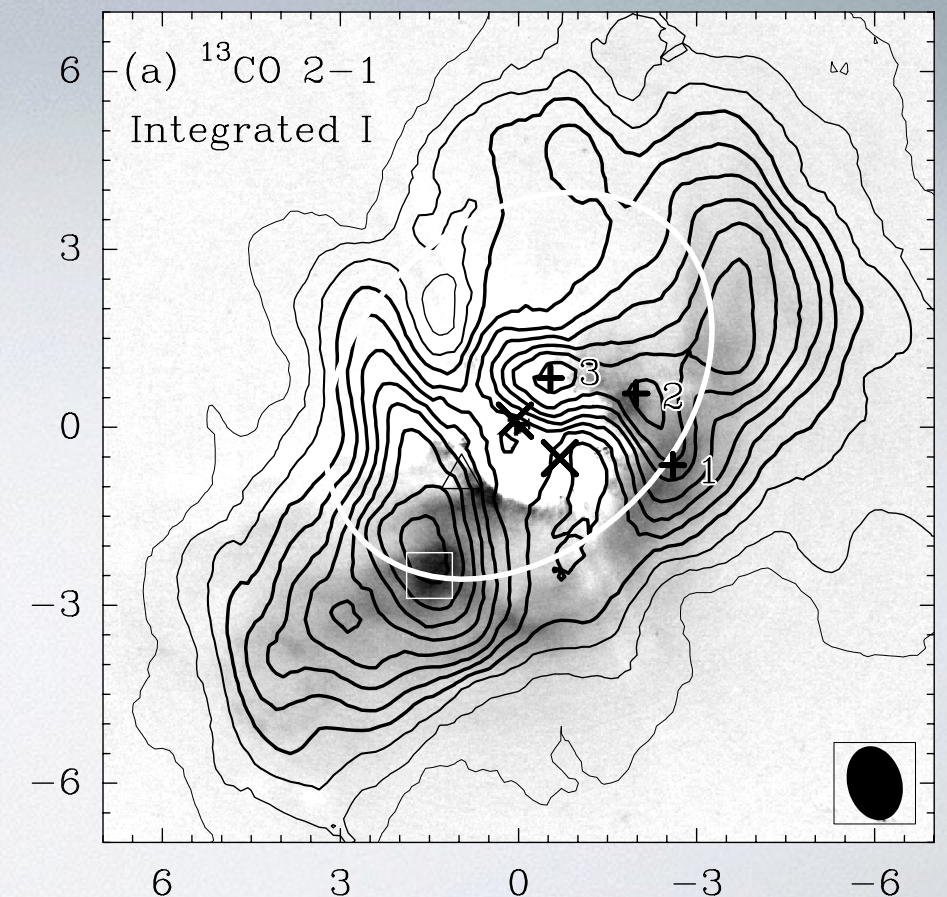
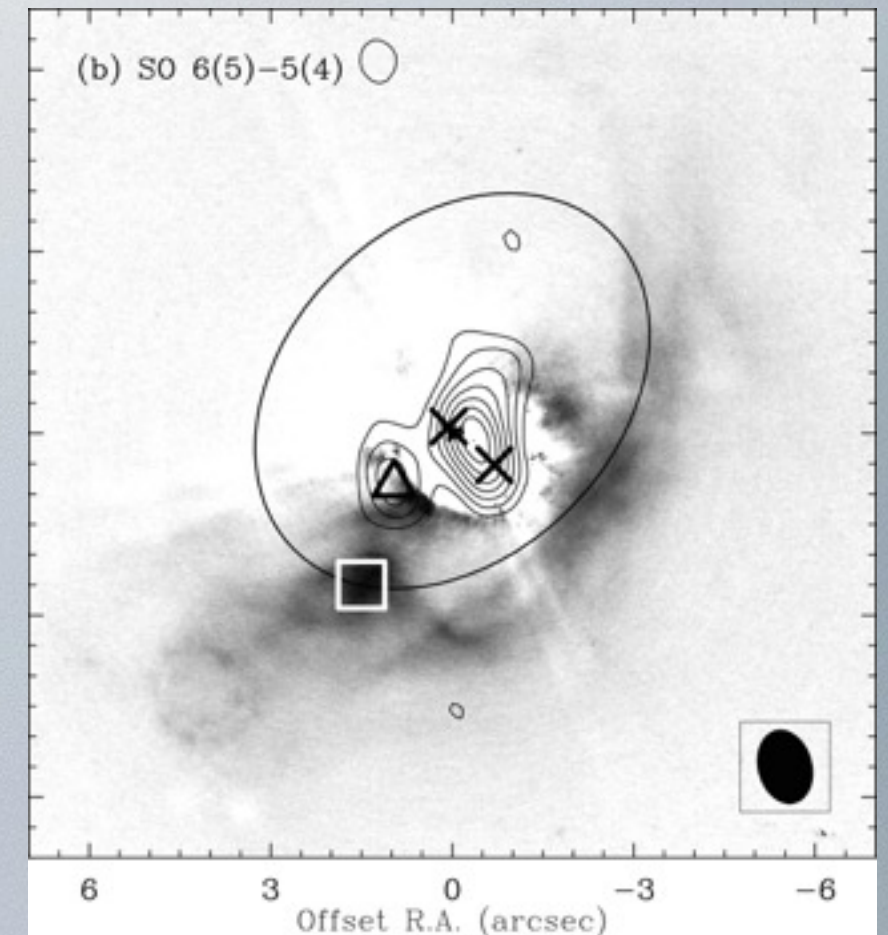
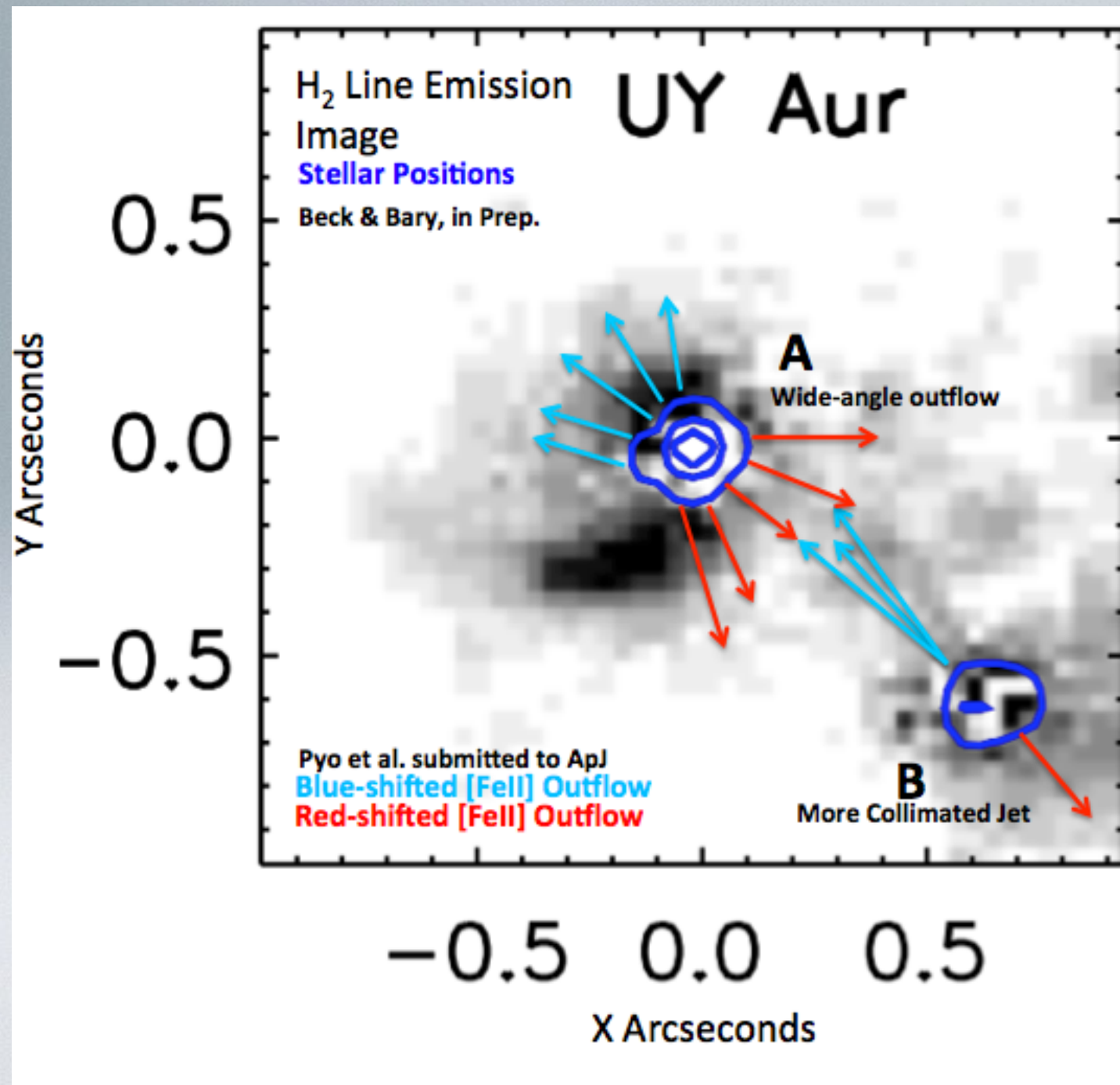
Circumbinary material:

New PdBI ^{13}CO 2-1, SMA CO 3-2, SO 5(6)-4(5)



- the suggested scenario of accreting spirals is further supported by the new H_2 2.12 μm line
- high velocity wings are seen in the spectra of ^{13}CO 2-1 and ^{12}CO 3-2 toward the binary
- ... and, there is SO 5(6)-4(5) detected toward the bases of the accreting spirals

Circumbinary material: ...and, there is SO



Summary

- Circumstellar disks are detected and resolved at 1.3mm
- the T Tauri binary system is still actively interacting with the surrounding envelope through a circumbinary disk
- evidences of streaming gas between the circumbinary disk and the binary stars are found
- the accretion rate toward the primary and secondary can not be determined with current data
- Angular momentum transfer...