

The next-generation space infrared astronomy mission

SPiCA
Space Infrared Telescope for Cosmology and Astrophysics

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Outline of my Talk

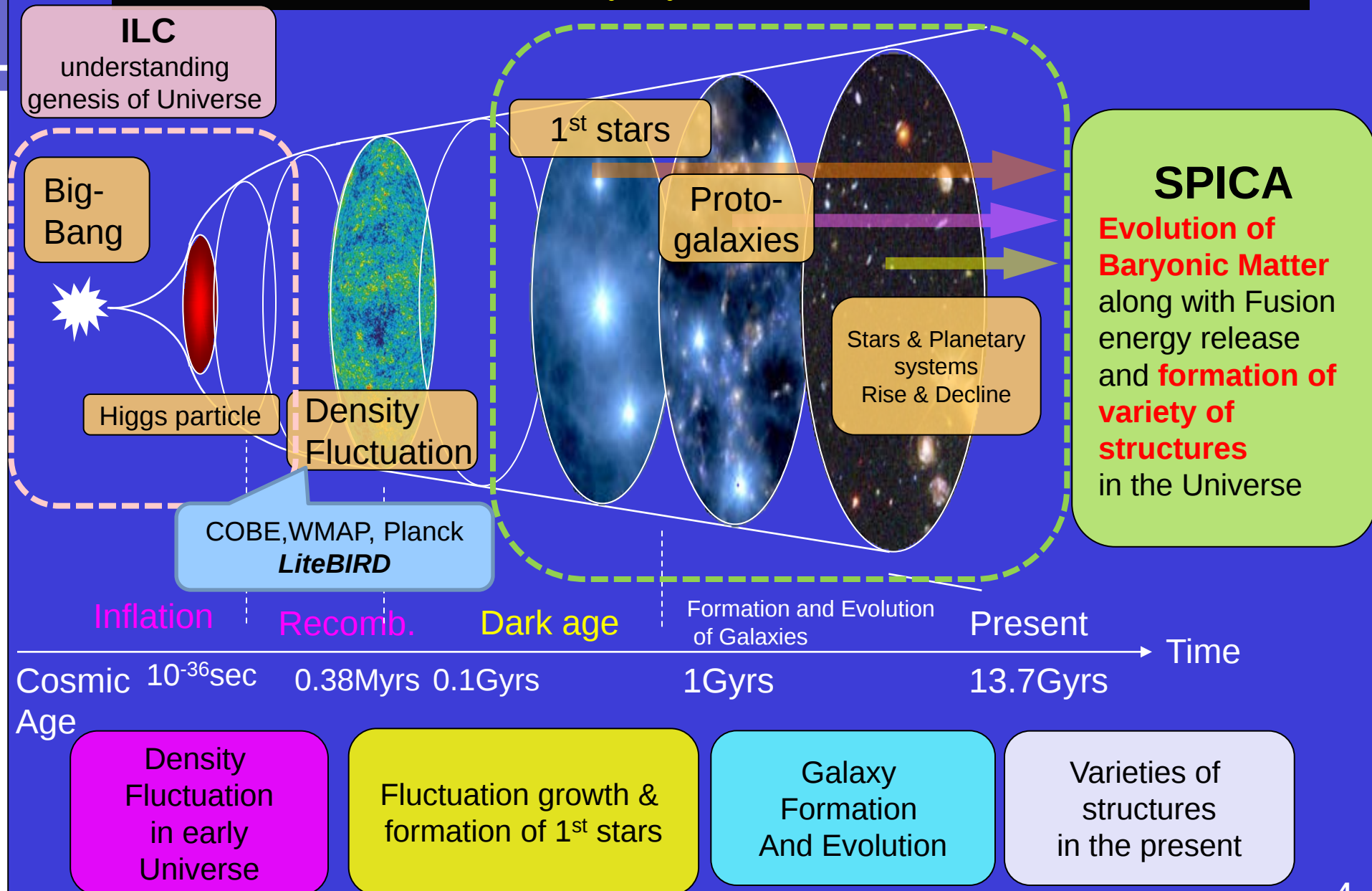
- Unique Scientific Goals of SPICA
- Mission Overview
- Programmatic Status

SPICA Science Goals

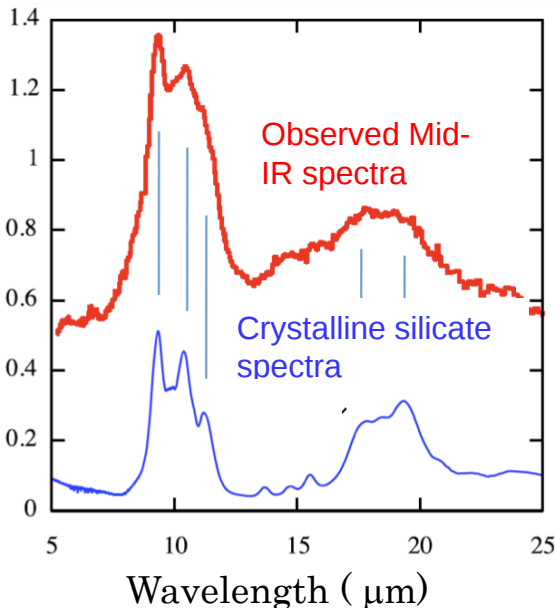
*Unveil the evolution of Baryonic
Matter in the Universe*

*Life cycle of interstellar &
Intergalactic matter
Planetary Systems Formation
Birth and Evolution of Galaxies*

Cosmic History from Genesis of the Universe to Stars & Planetary systems Formation

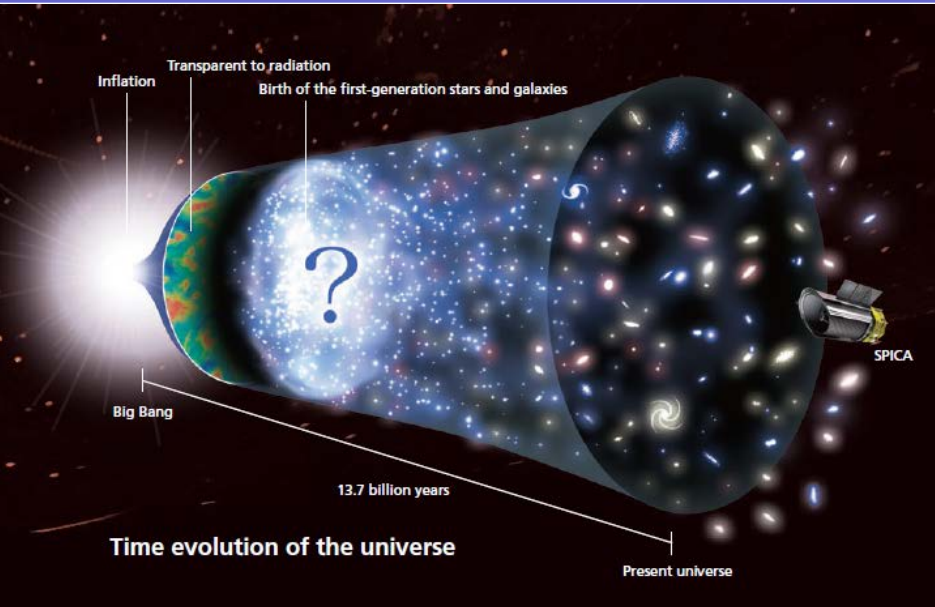


- Abundant spectral lines from atoms, ions and molecules for diagnostic tool.
- Broad-band spectral features provide the information of composition of solid state particles with ices



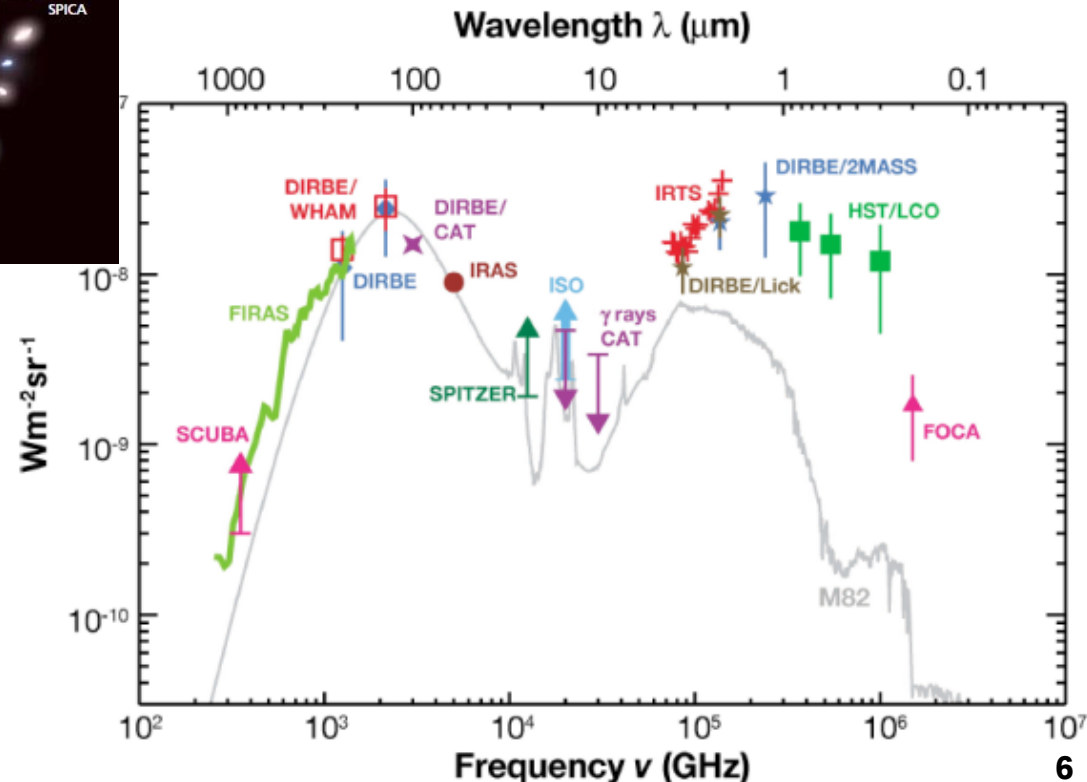
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One of Primary Scientific Goals: the birth & evolution of galaxies



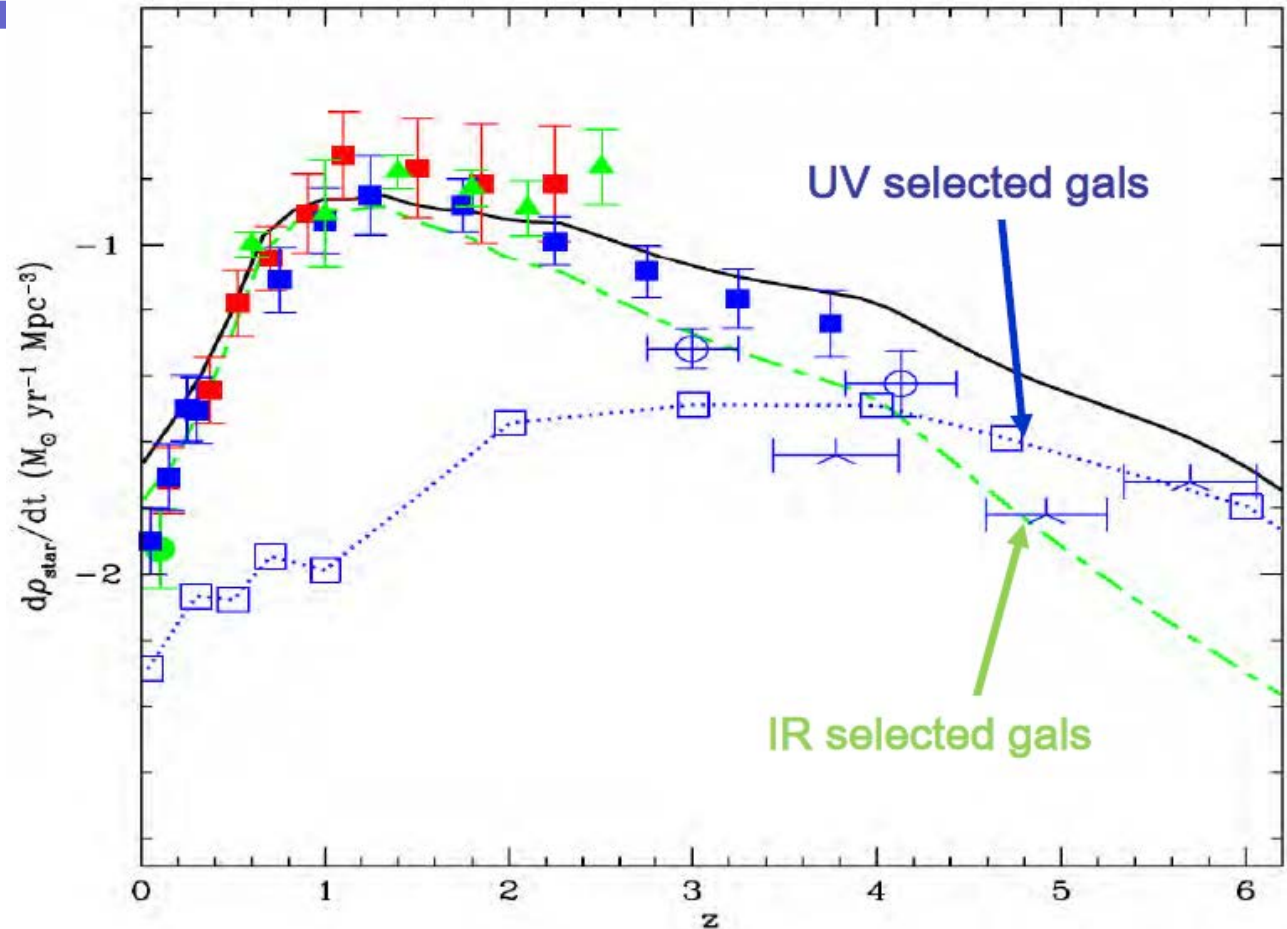
→ Hidden Universe to be
revealed by FIR Observations

- Cosmic IR background tells us that about half of nuclear fusion energy is hidden by dust



Cosmic Star-formation History revealed with Herschel

The bulk of SF activity at $z < 3-4$ appears to be produced by strongly dust-extinguished galaxies



Herschel: COSMOS+WIDE

Vaccari et al. 2013; Marchetti et al. 2013

250 μm

GOODS-N: 250/350/500 μm

350 μm

500 μm

But, What are they ?
Why this happens?

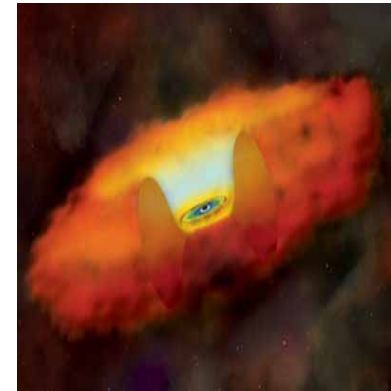
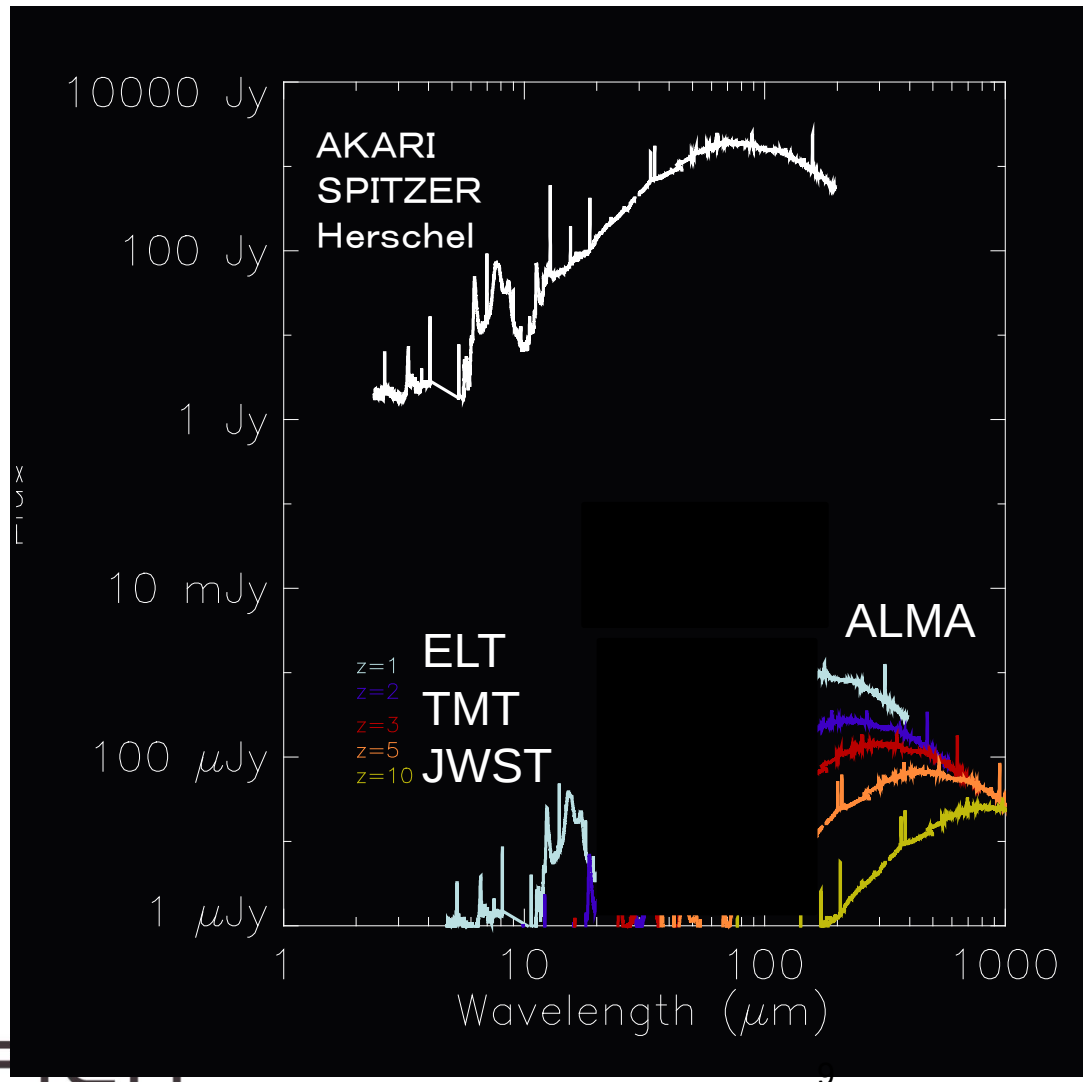


Spectroscopy!!

10 arcmin



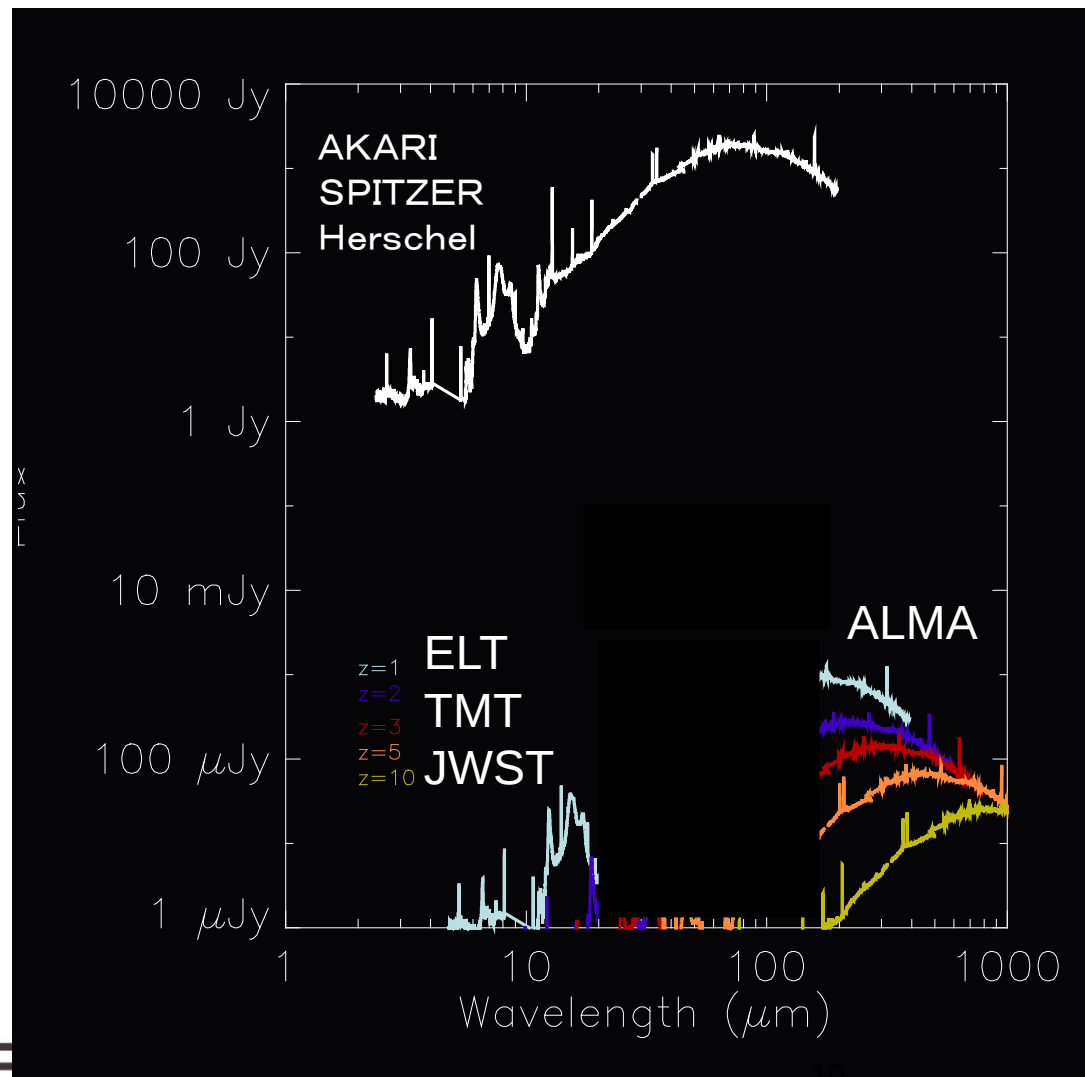
Characterizing Hidden Nature of Galaxies



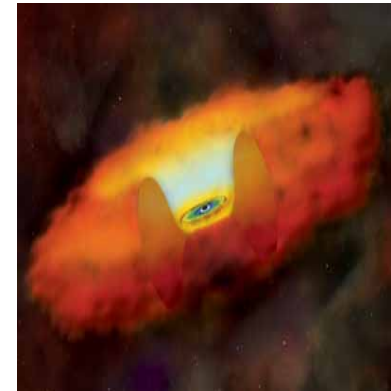
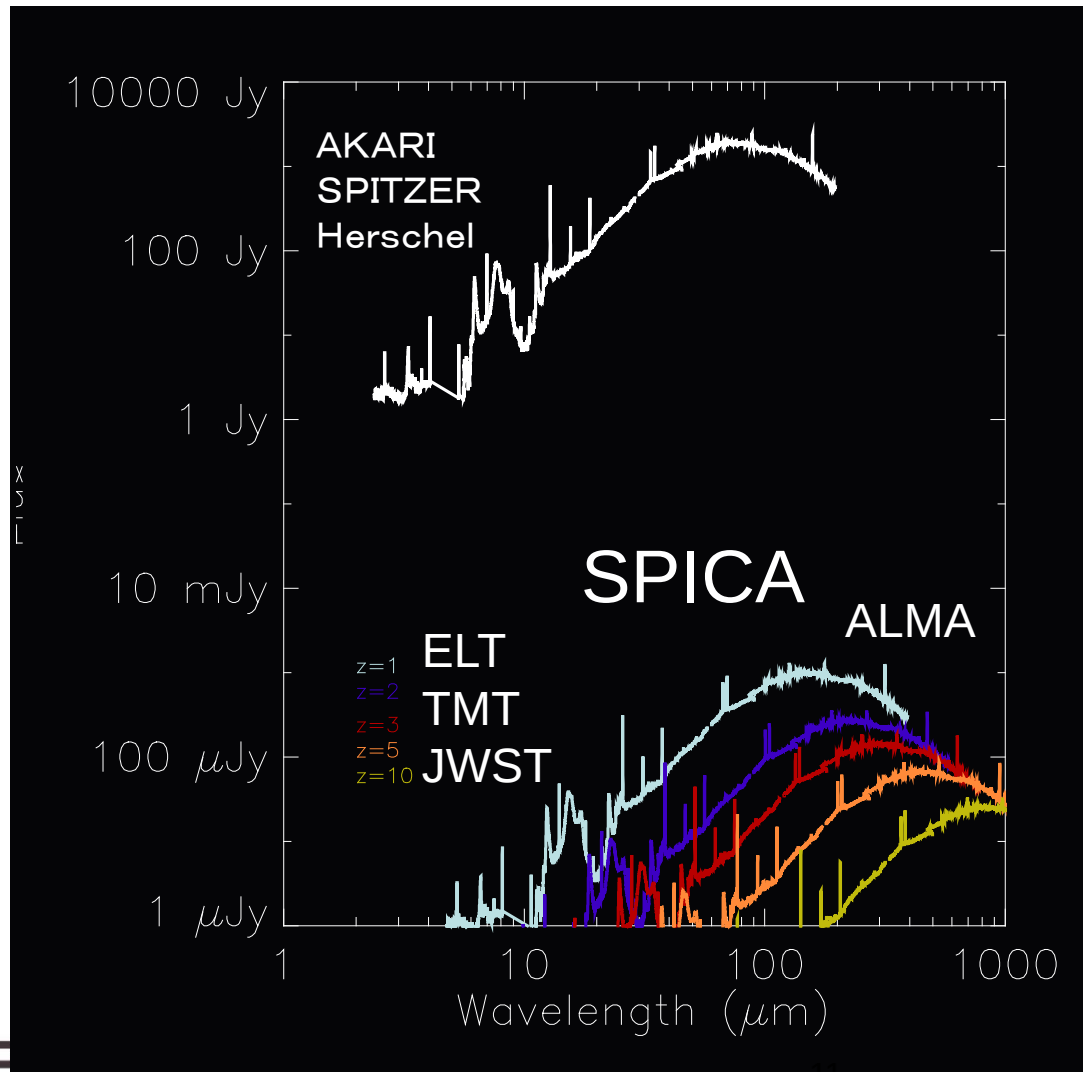
Revealing true nature
of obscured galaxies,
Starburst and /or Super
Massive Black holes

**SPICA
observations
are essential**

Characterizing Hidden Nature of Galaxies



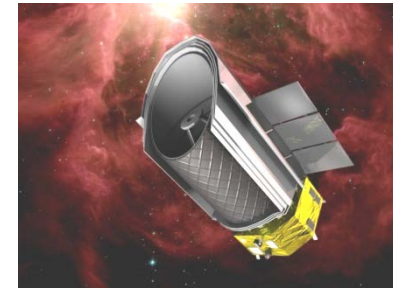
Characterizing Hidden Nature of Galaxies



Revealing true nature
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Mission Overview

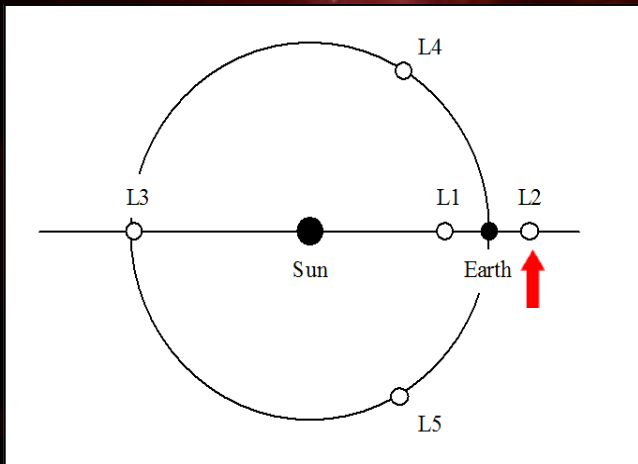


Unveil the evolution of Matter In the Universe

SPICA
Space Infrared Telescope for Cosmology and Astrophysics

SPICA Mission Overview

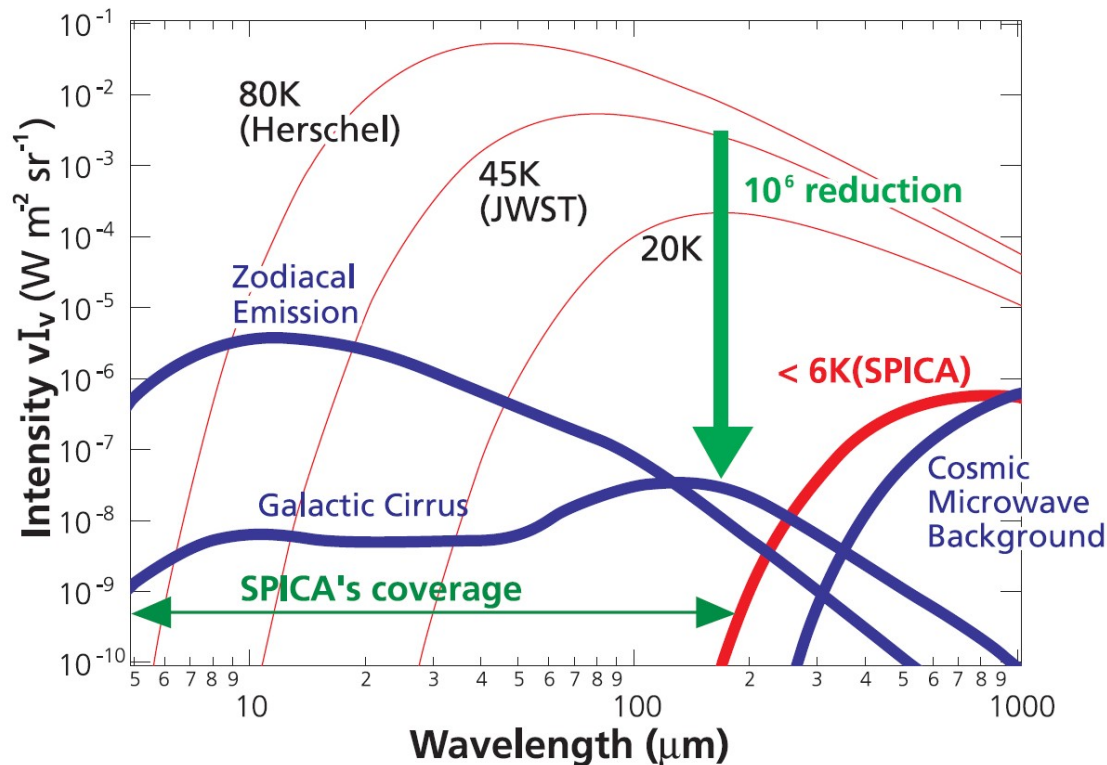
- **Telescope 3.2m (EPD 3.0m), 6K**
- **Orbit: Sun-Earth L2 Halo**
- **Mission Life: 3 years (nominal), 5 years (goal)**
- **Weight: 3.7 t**



- **Next-Generation Infrared Space Observatory with superior Sensitivity and good spatial resolution in mid- and Far-IR (5-210 μ m)**
- **International Mission**
 - **Japan, Europe, Korea, Taiwan**

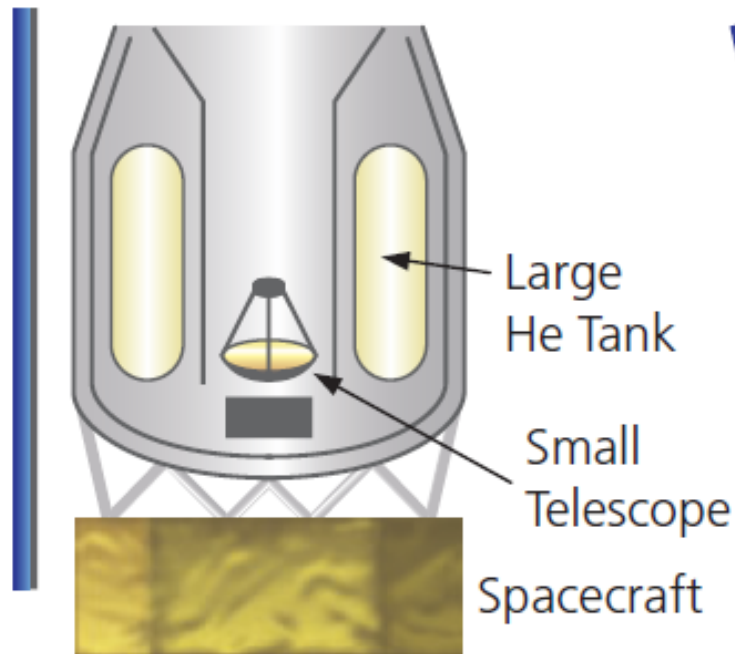
Requirements: Cooled Telescope

- $T < 10$ K is required to improve sensitivity
 - Background Radiation can be reduced by a factor of **one million** !

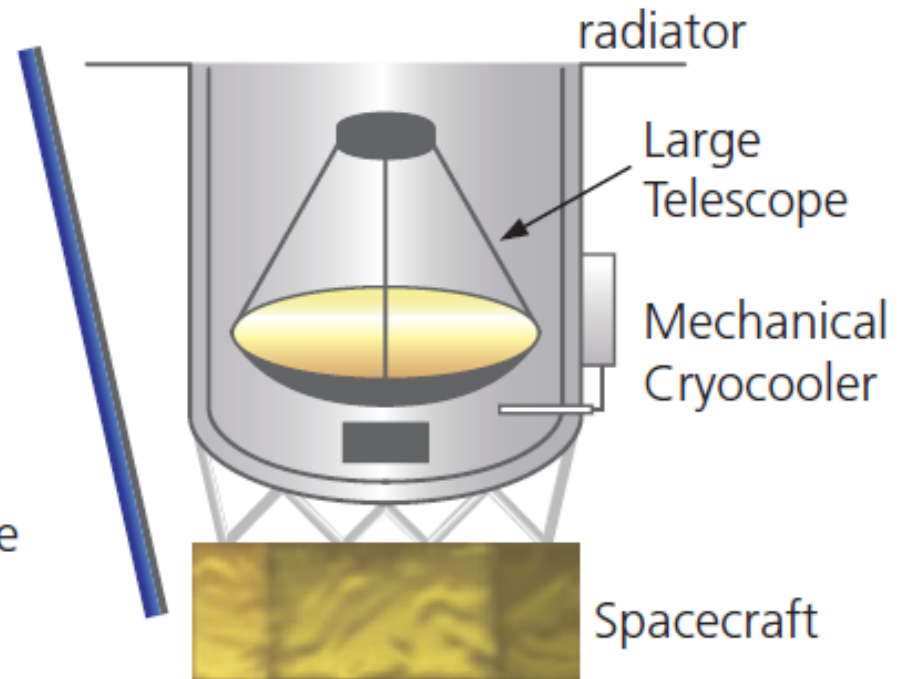


Cryogen-free mission

Today's Space Telescopes

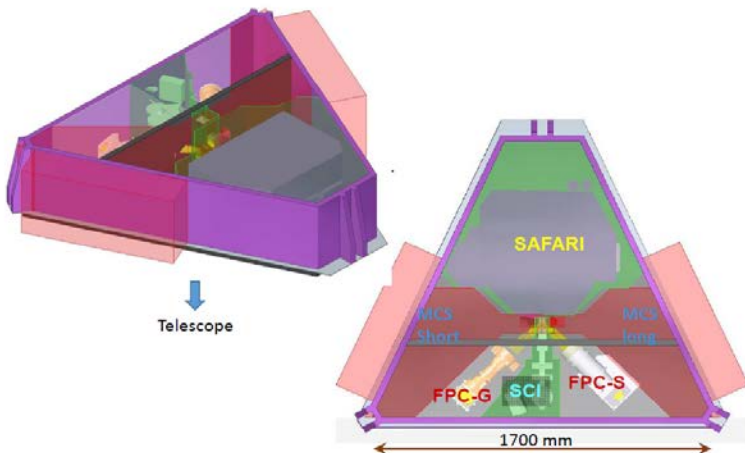
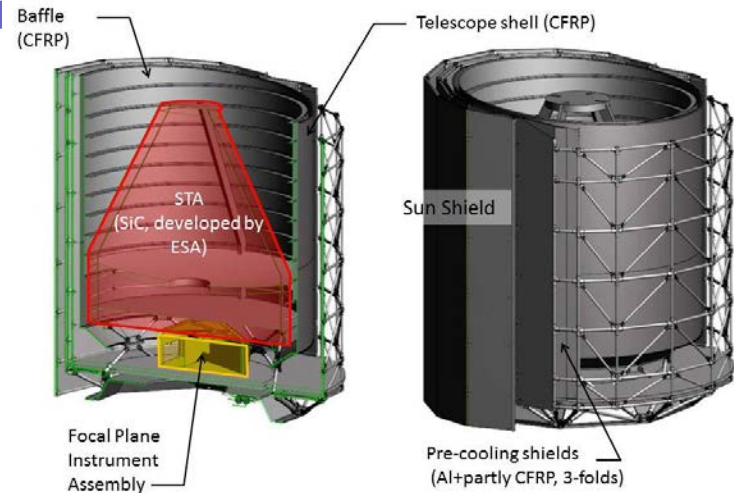


SPICA new design



Lighter and Larger

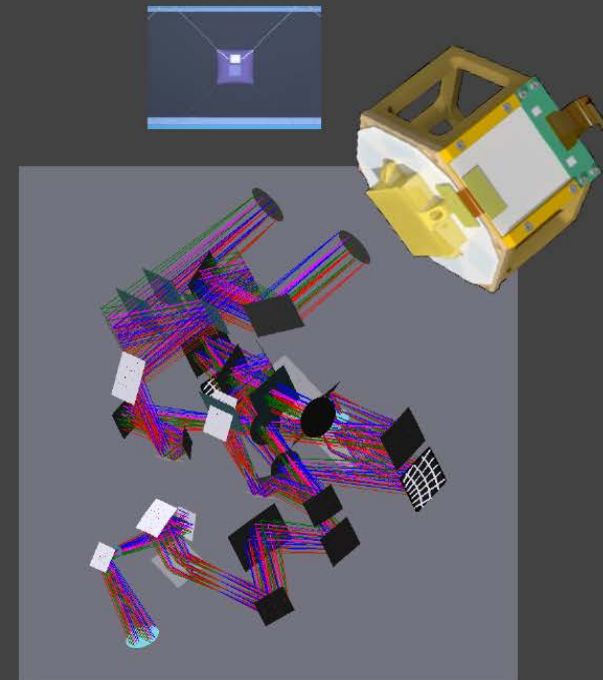
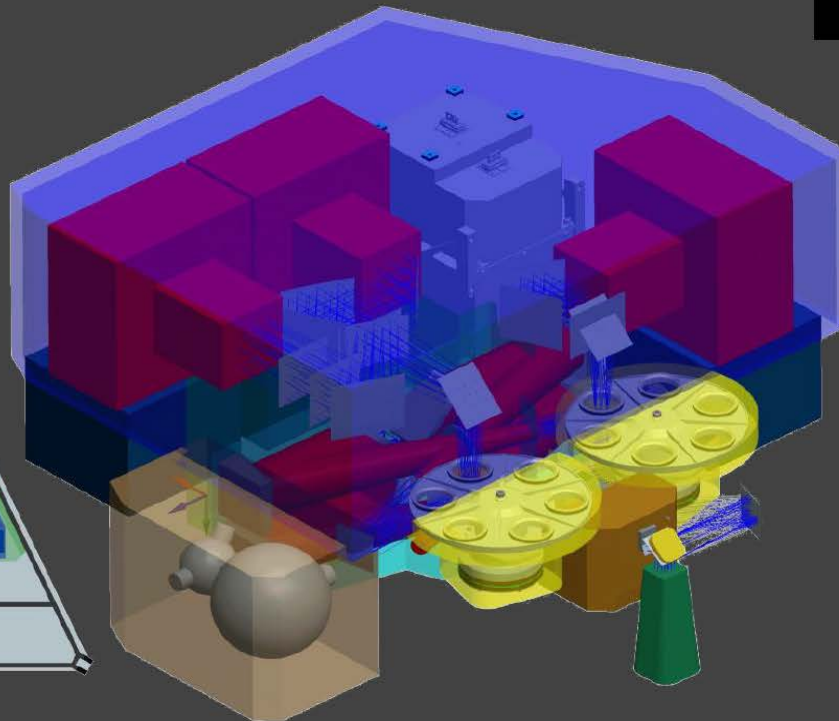
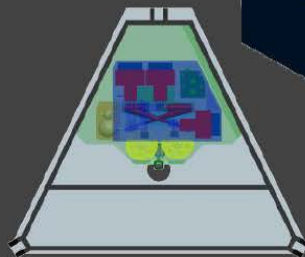
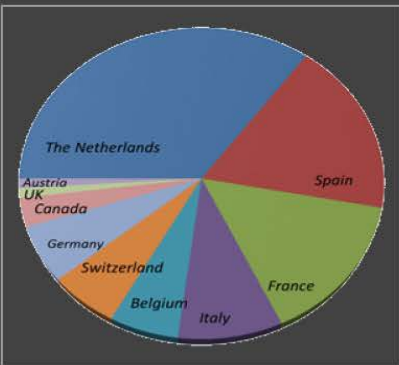
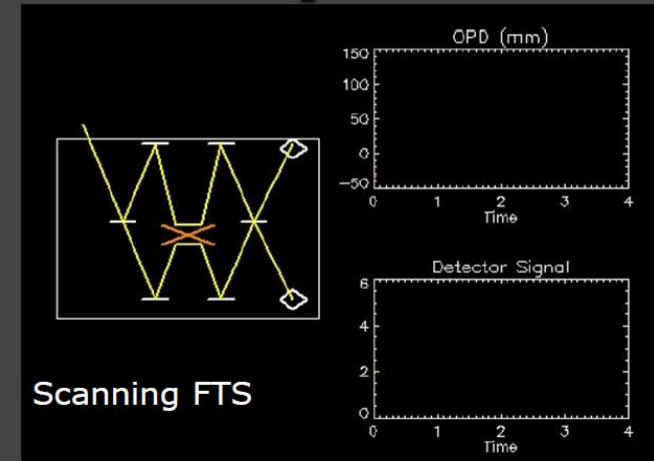
SPICA Focal Plane Instruments



- SAFARI
 - Far-infrared (34-210 μ m) imaging spectrometer
 - P.I. SRON (Netherlands) with SAFARI Consortium
- MCS
 - Mid-infrared (5-38 μ m) camera & spectrometer
 - P.I. JAXA, Universities, and ASIAA (Taiwan)
- SCI
 - SPICA coronagraphic instrument (4-28 μ m)
 - P.I. JAXA with Nagoya Univ.
- FPC
 - Near-infrared (0.7-5 μ m) camera and spectrometer
 - P.I. KASI (Korea)

The SAFARI instrument - summary

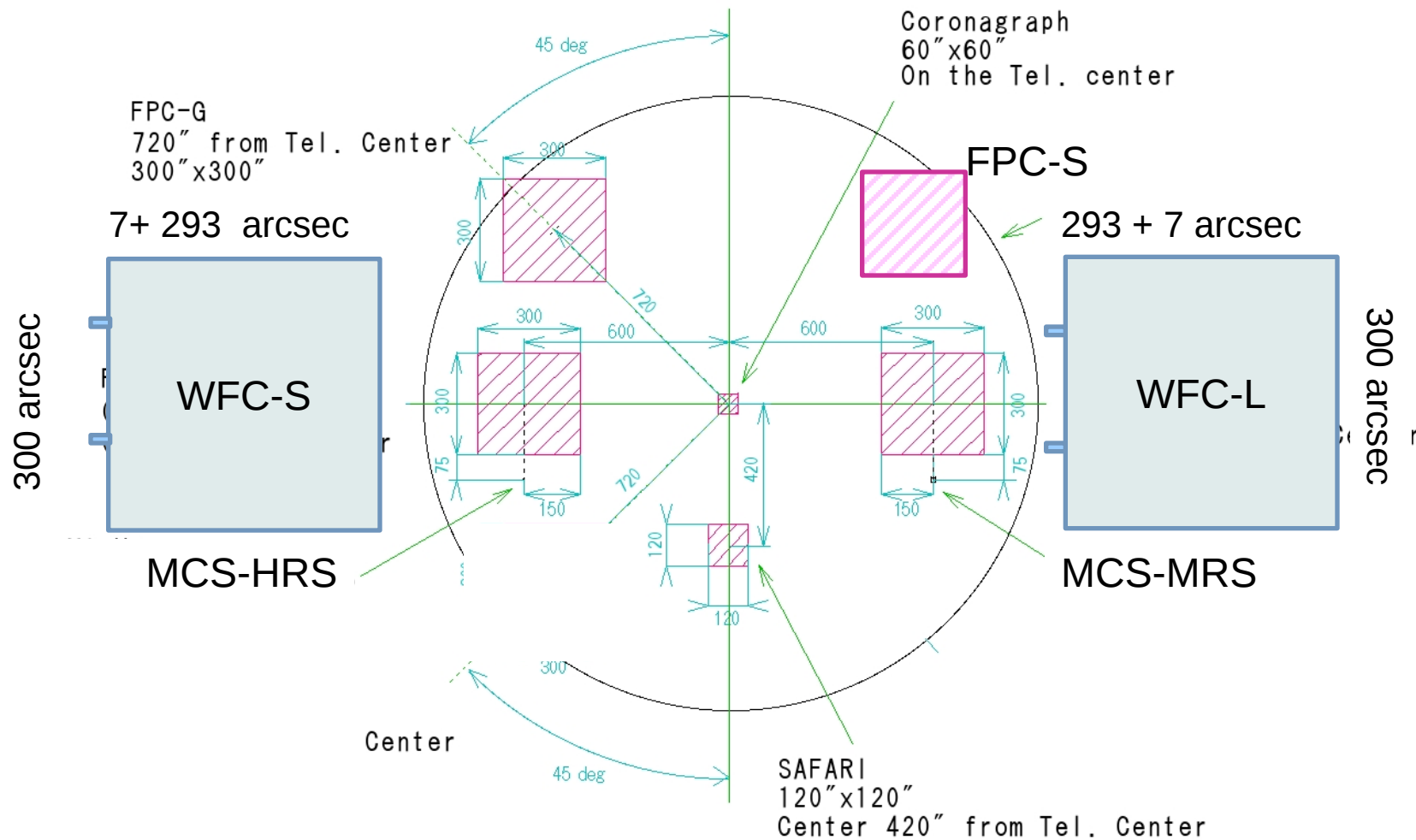
- Scanning Fourier Transform Spectrometer with 2'x2' FoV
- Simultaneously observing in 3 bands (34-210 μ m)
- Ultra sensitive TES detectors/SQUID read out at 50 mK
→ almost **200 times** more sensitive than Herschel
- Frequency Domain Multiplexing
- To be built by an SRON-led consortium
 - ~15 institutes in Europe, Canada, Japan - cost ~170M€



For MCS & FPC ..

- See next two presentations by:
 - “The Mid-Infrared Camera and Spectrometer for SPICA: general overview and Taiwan's contribution” by Ciska KEMPER (ASIAA)
 - “Korean Contribution to SPICA: Development of NIR Instrument, FPC” by Woong-Seob JEONG (KASI)

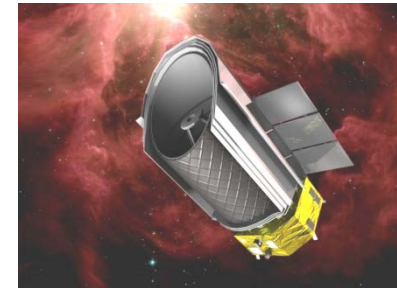
Field-of Views



Updated on June 2013

Seen from a telescope side

Programmatic Aspects

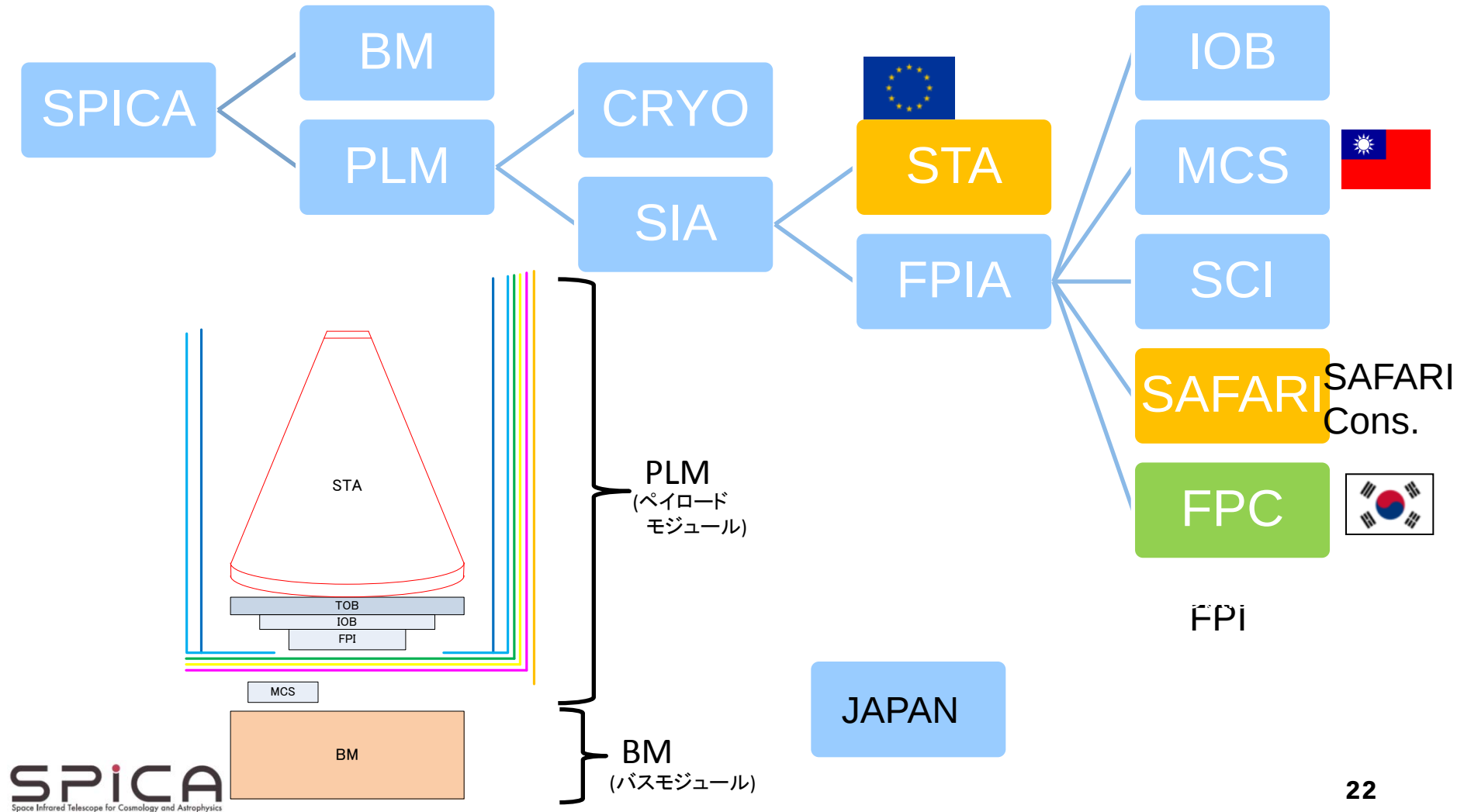


Situation in Japan

- SPICA FY 2014 budgetary proposal (Ministry of education proposal, end of August)
 - Only R&D activity was funded
 - Project itself was not funded
- Space Policy Committee, Cabinet Office of Japan Government, released new long-term plan of Space Science program of Japan (late Sep., 2013)
 - Strategic projects (300M\$ class), one in 5 years (large projects need more discussions)
 - Small satellites (100-150M\$), three in 5 years
 - Small projects (10M\$/year), flexible
 - Astrophysics in the next 5 years
 - ASTRO-H (Launch FY 2015)
 - **SPICA**

SPICA has the top priority among future science satellite candidates, keeping consistency with the new long-term plan

International Collaboration = *Original Scheme* =



A New plan proposed

- Dedicated Discussions were held (among the director level) in order to establish a more feasible plan both from programmatic & technical points of view
- Proposal
 - To increase the role of European contribution, while keeping SPICA as a JAXA-led project
- Implication
 - Increases technical feasibility by using European technical heritage (e.g. Herschel , Planck)
 - No major change in the SPICA program as a whole (science goals, mission specifications). No change in the importance of Korea and Taiwan participation; essentially important
 - Major changes in the approval processes and schedule
 - Re-entering **open competition in the ESA Cosmic Vision program (M-class mission, M4)**
 - The project starts in 2017, and the launch in 2026 in the nominal track

International SPICA Team

- 17 countries, regions and one International org.



Summary

- What's SPICA: next-generation space IR observatory, with cold ($<6\text{K}$) 3.2m diameter telescope
- *Much higher sensitivity in the thermal infrared than Herschel Space Observatory*
- SPICA can explore the evolution of Baryonic matter of various form in the history of the Universe
- Due to the budgetary situation in Japan, the international collaboration scheme needs to be revisited to increase the role of European contribution, while keeping SPICA as a JAXA-led project.
- Thus SPICA re-enters open competition in the ESA Cosmic Vision program (M-class mission, M4)
- In order to realize SPICA in 2020s, the international community support and collaboration with namely Asian partners are essential.

Wavelength

Development of Space IR Astronomy

Herschel (observatory)

2009- 2013 (ESA)

Diameter 3.5m ϕ

Wavelength: 55~700 μ m

**Unique Space IR
Observatory for
the Resolution
of Matter in the
Universe**



AKARI (surveyor)

2006- 2011 (Japan)

Diameter: 0.685m ϕ

Wavelength: 2~180 μ m

Improvement of
Observation
technology



SPICA (Observatory)

2026- (Japan (Asia)& Europe)

Diameter: 3.2m ϕ

Wavelength: 5~210 μ m

**New Generation
All-Sky Survey**

JWST (Observatory)

2018- (USA)

Diameter: 6.5m ϕ

Wavelength: 0.6~28 μ m



WISE (surveyor)

2009-2011 (USA)

Diameter: 0.4m ϕ

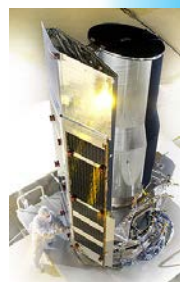
Wavelength: 3~25 μ m

SPITZER (Observatory)

2003- (USA)

(2009 Liq. He boiling-off)

Diameter: 0.85m ϕ , Wavelength: 3~180 μ m



Improvement of
Observation
technology

ISO (Observatory)

1995-1998 (ESA)

Diameter: 0.60m ϕ

Wavelength: 2.4~240 μ m



**1st insight of evolution of matter
by Detailed Observations**



IRTS (surveyor)

1995 (Japan/US)

Diameter: 0.15m ϕ

Wavelength: 1~700 μ m

Initial Survey Era



IRAS (surveyor)

1983 (USA/UK/Netherlands)

Diameter: 0.57m ϕ

Wavelength: 8~100 μ m

Mid-IR

Near-IR

Submm

Far-IR

100 μ m

10 μ m

1 μ m

1980

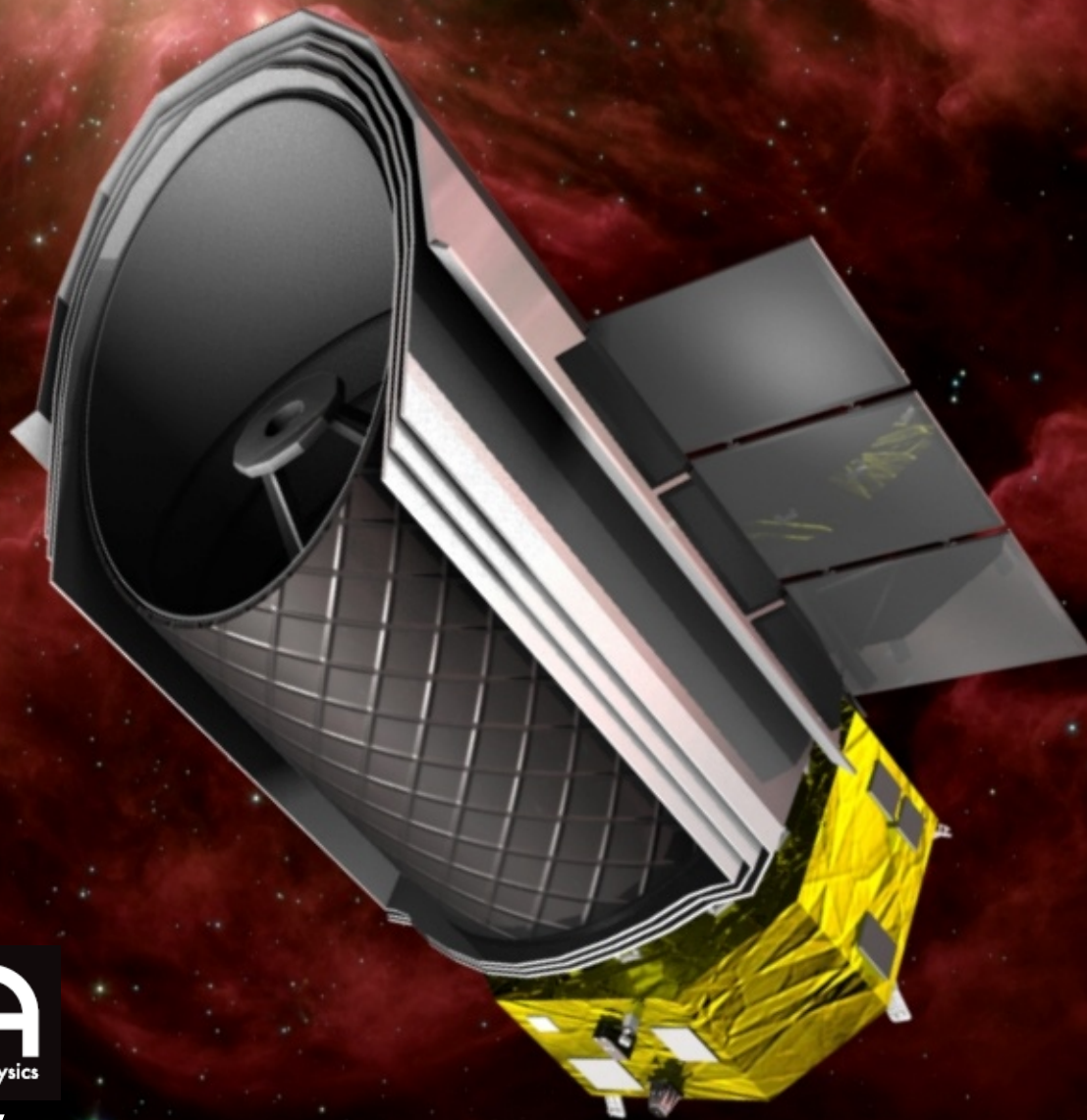
1990

2000

2010

2020

Years



SPICA

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Space Odyssey