
QSO IDENTIFICATION WITH BATC MULTIBAND DATA

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Research Goal

**A new trial for identifying QSO
with BATC multiband data**

Traditional Methods

- Radio observation
 - Quasar was discovered through the optical identification of strong radio source, after that lots of radio-loud quasars are detected

 - X-ray detection
 - Investigation of strong x-ray sources

 - Optical identification
 - Selection of QSO candidates from color-color diagram and follow-up spectroscopic observations
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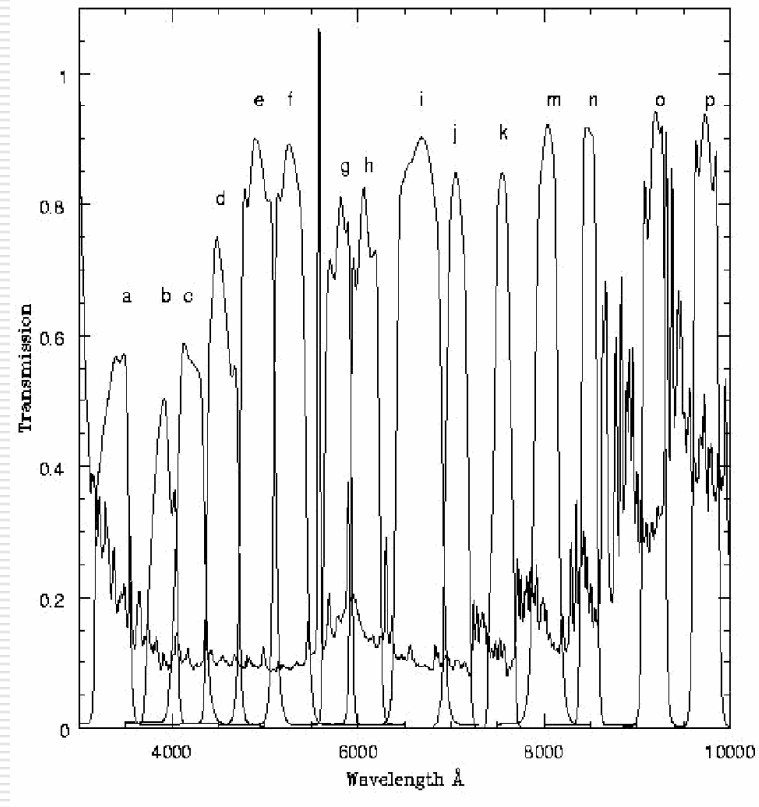
My Selection Criterion

Spectral Energy Distribution

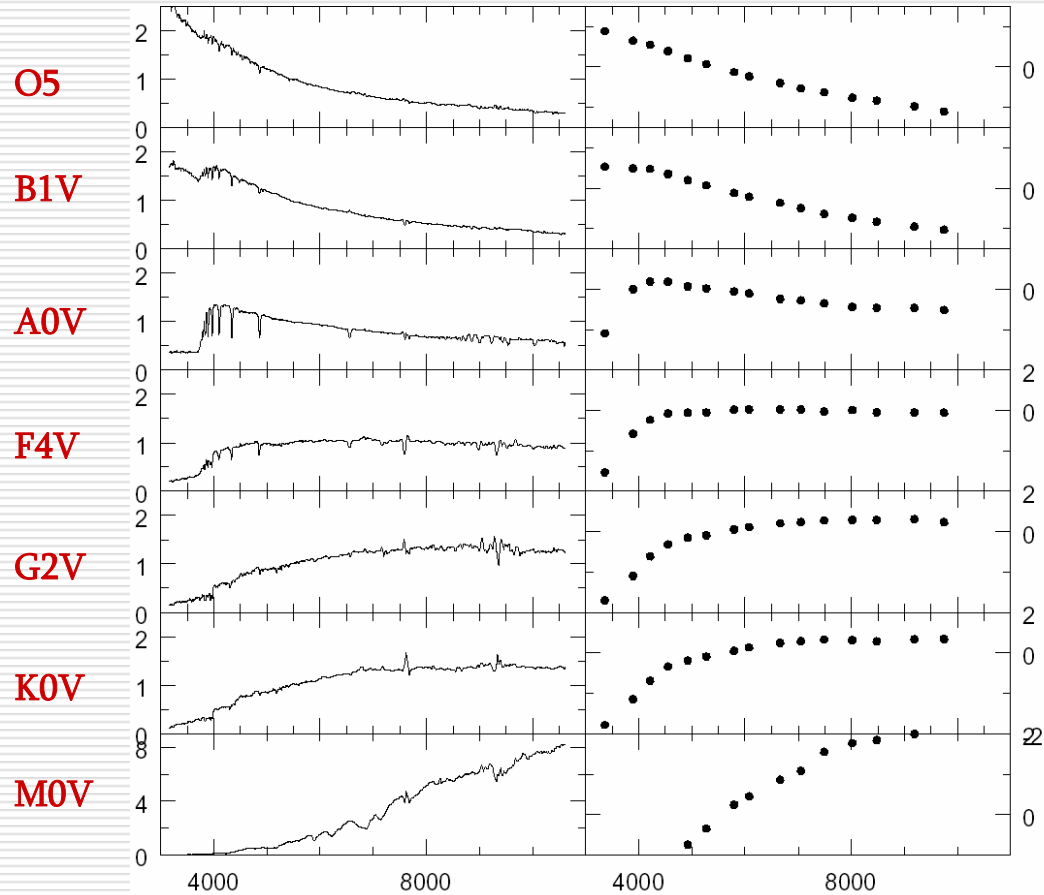
obtained from spectrophotometry

BATC Color Survey

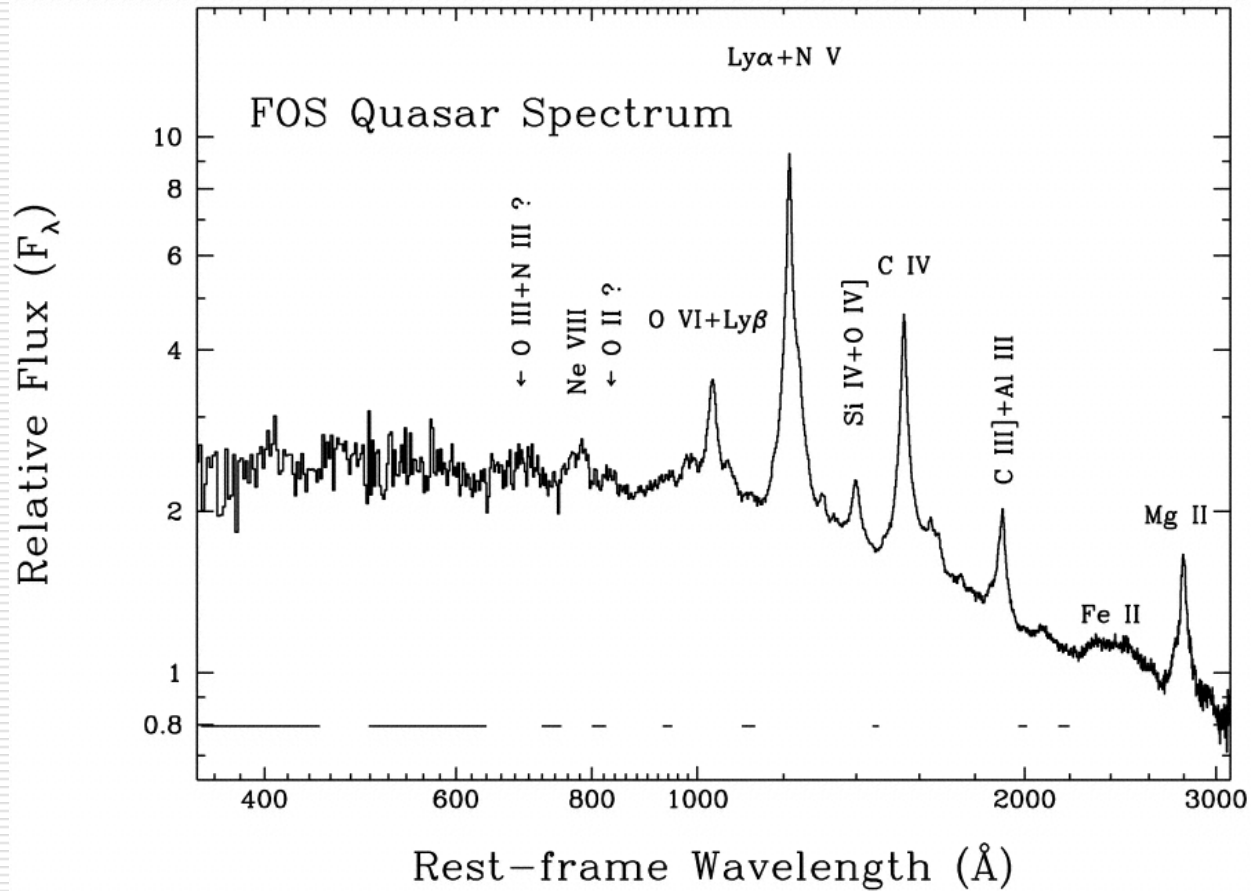
- ❑ BATC(Beijing-Arizona-Taipei-Connecticut) color survey uses 15 intermediate-band filters.
- ❑ Filters are designed to avoid strong night sky emission lines
- ❑ Spectral coverage :
3200Å ~ 9900Å
- ❑ FOV ~ 1° x 1°



Stellar SED from BATC System

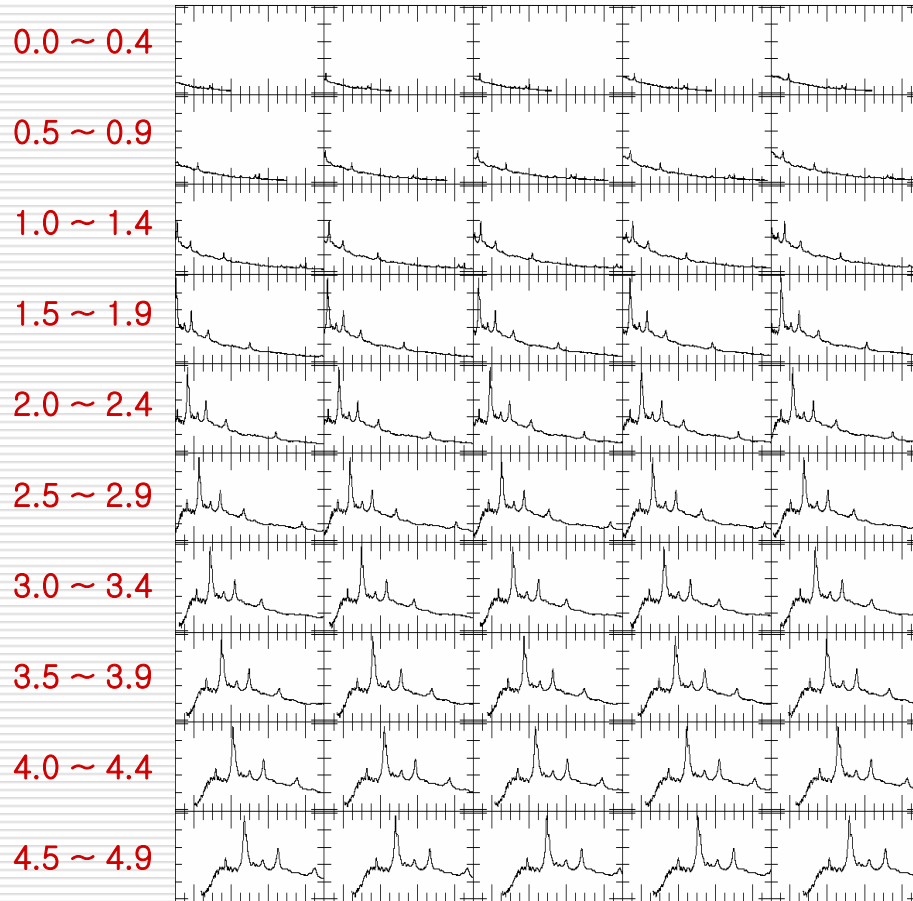


QSO Composite Spectrum

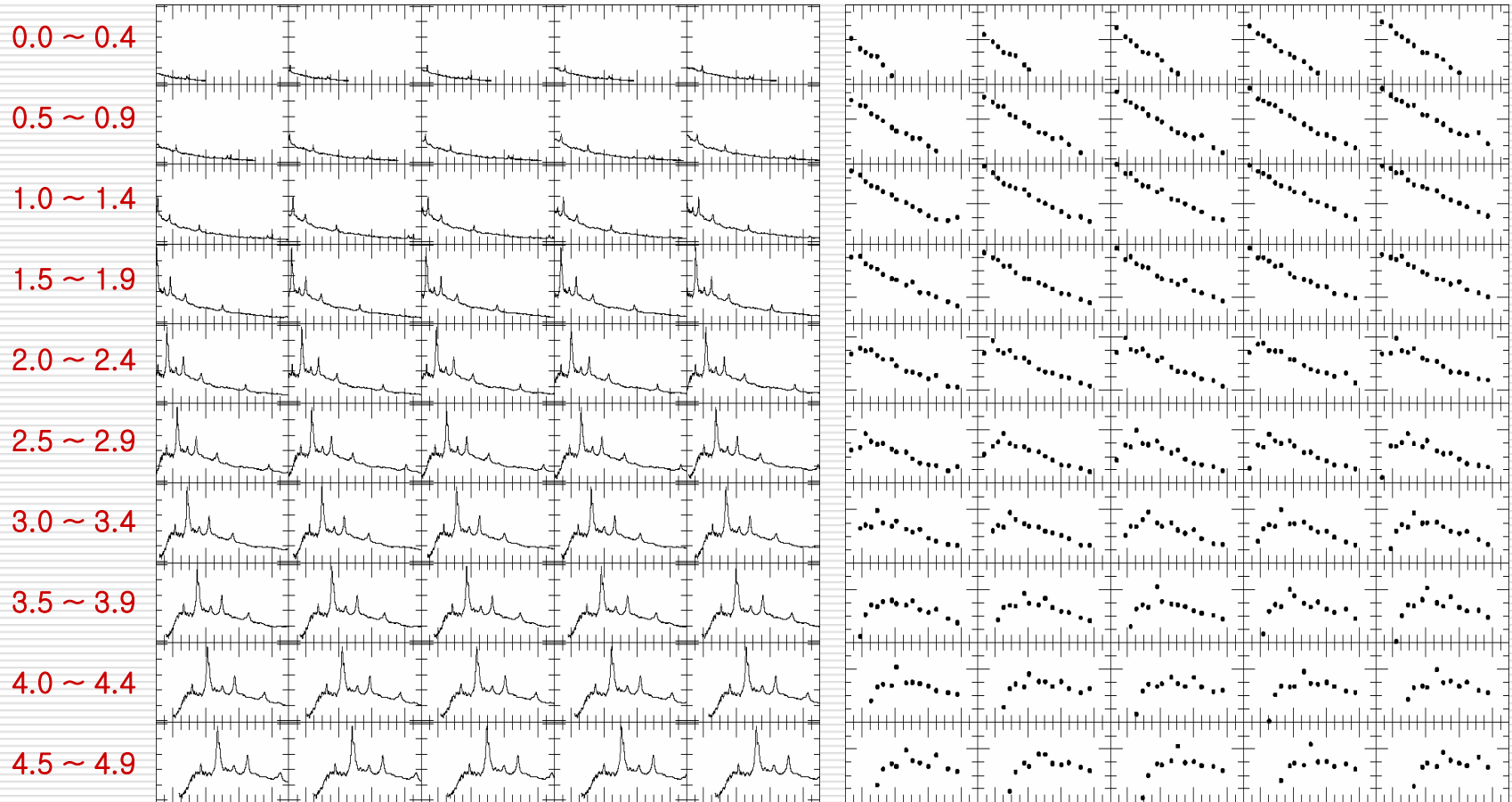


Zheng et al. 1997

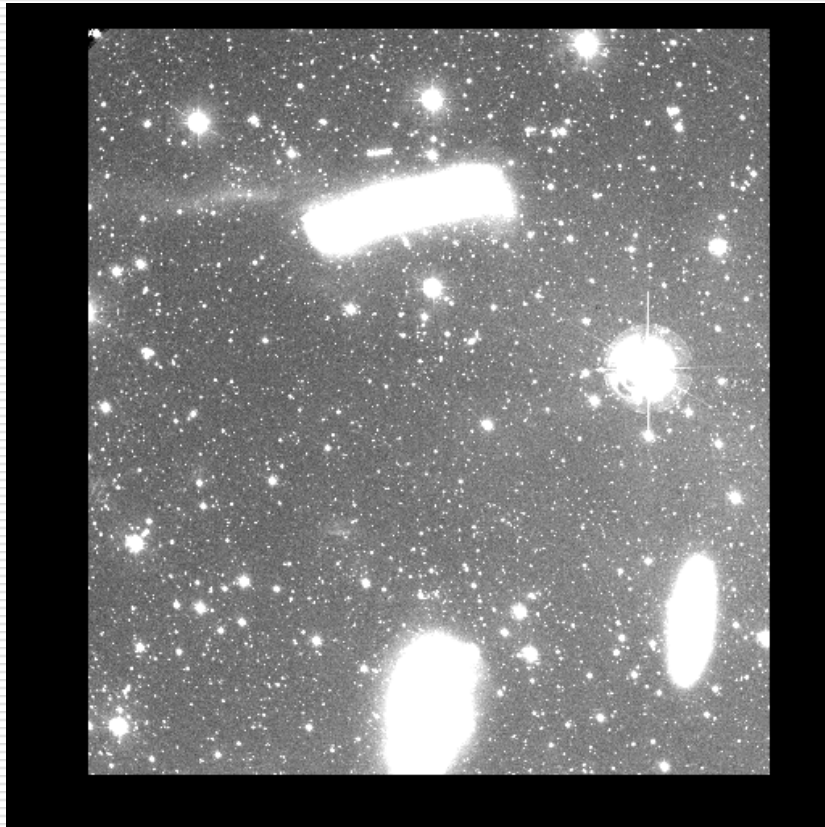
Redshift Simulation for QSO Spectra



Redshift Simulation for QSO Spectra



Leo Triplet System



- Leo Triplet is composed of three galaxies (NGC3628,NGC3627, NGC3623) interacting each other for their tidal forces
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Data Reduction

Image Stacking for High S/N



Photometry with SExtractor



Standardization

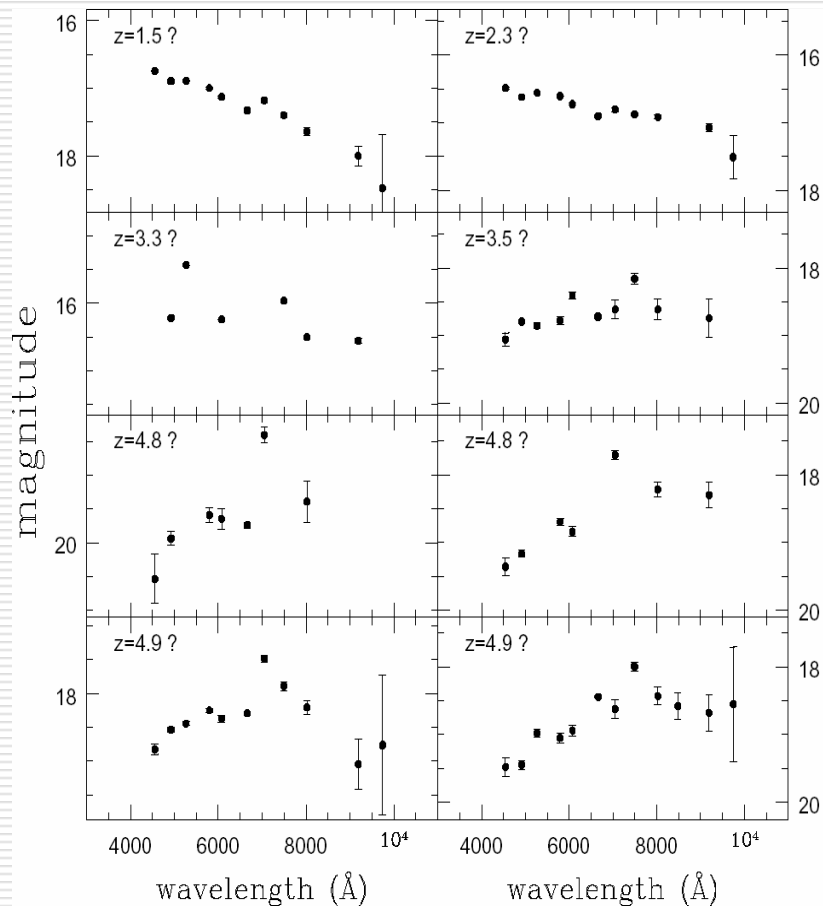


Star Match



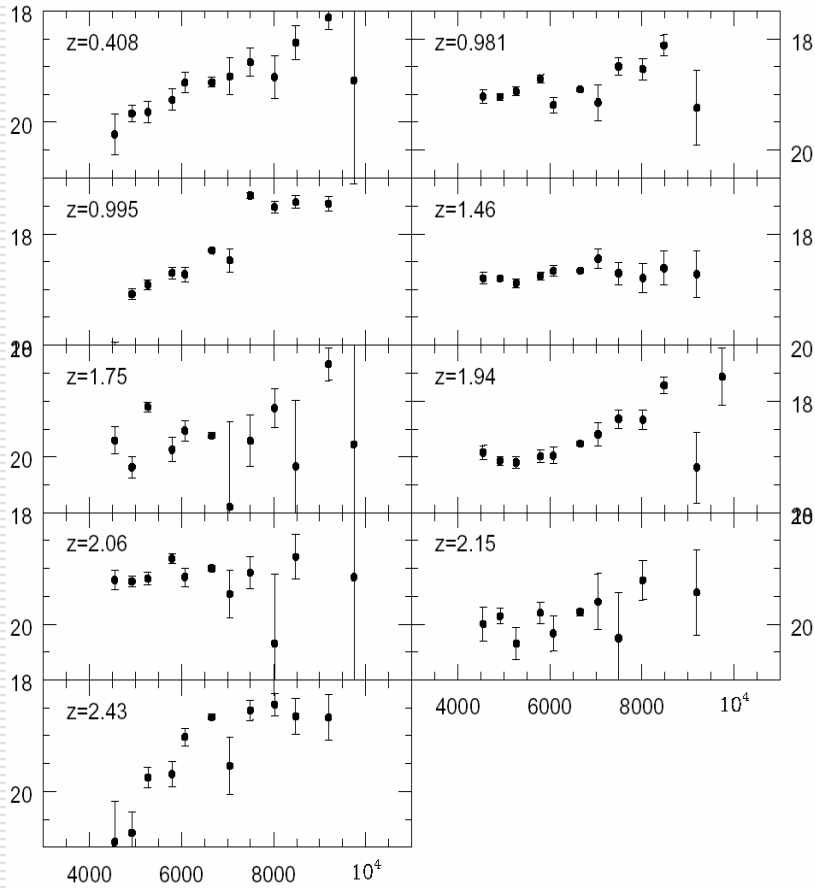
SED Construction

New QSO candidates



- More than 17 new QSO candidates
- They are selected based on their slope and emission lines in the SED
- Follow-up observations are planned

Known QSOs in Leo Triplet field



- No general agreement with model predictions for given redshifts
- Some of them may not be real QSOs
- Some of them may have different redshift

Summary

- We performed QSO identification with BATC multiband data
 - 17 new QSO candidates were selected by examination of SED and follow-up observations are planned
 - Known properties of QSOs are not enough to identify them
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